

1 Title: The distribution of platinum-group elements and Te, As, Bi, Sb and Se
2 (TABS+) in the Paraná Magmatic Province: Effects of crystal fractionation, sulfide
3 segregation and magma degassing
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11 **<https://doi.org/10.1016/j.lithos.2021.106374> Abstract**

12
13 The concentrations of Te, As, Bi, Sb and Se (TABS+) in magmas from large igneous
14 provinces (LIPs) are of interest because these elements are important in the formation of
15 platinum-group minerals (PGM) found in magmatic Ni-Cu platinum-group element
16 (PGE) deposits. Furthermore, the TABS+ are volatile and hence they may be lost in
17 degassing and have a role to play in the mass extinctions associated with LIPs events
18 because these are mostly toxic elements. However, the concentrations of TABS+ in
19 magmas from LIPs are not well documented due in part to the analytical difficulties and
20 in part to the numerous processes that affect their distribution. We have determined the
21 concentration of TABS+ and PGE in rocks from the Paraná Magmatic Province (PMP)
22 in order to assess the effects of fractional crystallization, sulfide segregation and magma
23 degassing. Decrease in the Rh, Ru and Ir concentrations with Mg and Cr indicate that
24 these elements behaved as compatible elements during crystal fractionation. Based on
25 changes in the Cu/Pd sulfide saturation occurred in some cases resulting into a decrease
26 in Pd and Pt. Arsenic and Sb behave as incompatible elements, whereas Se, Te and Bi
27 show variable behavior. Most lavas from the PMP display negative As, Se and Te
28 anomalies on mantle-normalized patterns coupled with Te/Cu, Se/Cu and As/Th and Sb/Th

29 ratios lower than mantle values, which we attribute to loss of the TABS+ during
30 degassing.

31

32 Key words: Paraná Magmatic Province; Te, As, Bi, Sb and Se (TABS+); Platinum-group
33 element; Fractionation; Sulfide segregation; Degassing

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36

37 **1 – Introduction**

38 The distribution of chalcophile elements such as the platinum-group elements
39 (PGE) in lavas from different geological settings has been extensively investigated
40 over the past decades (Barnes et al., 1985; Bockrath et al., 2004; Pitcher et al., 2009;
41 Hughes et al., 2015). However, recent contributions have also investigated a wider
42 range of chalcophile elements, which include Te, As, Bi, Sb (TABS) and Se (Jenner
43 and O'Neill, 2012; König et al., 2012; Lissner et al., 2014; Mansur et al., 2020a;
44 Maciag and Brenan, 2020). The interest in better understanding the distribution of PGE
45 and TABS in rocks from different environments is the result of their ability to record
46 a wide range of geological processes. These processes include Earth's differentiation,
47 fractional crystallization, crustal assimilation, sulfide segregation, magma degassing
48 (König et al., 2012; Lissner et al., 2014; Edmonds et al., 2018; Wieser et al., 2020;
49 Maciag and Brenan, 2020), and the formation of mineral deposits (Chen et al., 2013;
50 Mansur and Barnes, 2020a, Mansur et al., 2020a).

51 In the current work we present new whole-rock analyses of S and PGE, and the
52 first TABS and Se (TABS)+ analyses for samples from different portions of the Paraná
53 Magmatic Province (PMP). The rocks represent both high- and low-Ti basalts
54 (summarized by Peate 1997) and intrusive rocks from the Ponta Grossa and

55 Florianópolis dyke swarms. The results are presented together with those for samples
56 with previously published PGE results from the PMP (Mincato, 2000; Rocha-Júnior
57 et al., 2012), and also from the Etendeka Province (Maier et al., 2003; Barnes and
58 Mansur, *in press*), which together form a Large Igneous Province (LIP) – the Paraná-
59 Etendeka Magmatic Province (PEMP). Results for TABS+ are also compared with
60 previously obtained values for other LIPs to refine our understanding of their
61 geochemical cycle. This contribution documents the effects of fractional
62 crystallization, sulfide segregation and magma degassing on the distribution of PGE
63 and TABS+ in the PMP.

64

65 **2 – Regional Setting**

66

67 The PEMP is a LIP associated with the Gondwana break-up during Early
68 Cretaceous (~135-130 Ma) and the opening of the South Atlantic Ocean (Peate, 1997;
69 Thiede and Vasconcelos, 2010; Janasi et al., 2011; Pinto et al., 2011a). It comprises
70 mainly subaerial lavas and related dyke swarms at central South America (Brazil,
71 Argentina, Uruguay, and Paraguay) and southwestern Africa (Angola and Namibia).
72 In total, the PEMP covers an area of approximately 920,000 km² (Frank et al., 2009),
73 of which the vast majority is exposed in the Paraná Sedimentary Basin, South America
74 (Fig. 1). The remaining African counterpart of the province comprises the bimodal
75 magmatism exposed in the Etendeka region (Maier et al., 2003; Frank et al., 2009;
76 Marsh and Swart, 2018).

77

78 In Brazil, the volcanic pile is referred to as the Serra Geral Group (Fodor et al.,
79 1989; Rosseti et al., 2018), and reaches up to 1700 m thick in the central portion of the
80 basin (Frank et al., 2009). The rocks are divided on the basis of geochemical criteria,
comprising six basaltic and two silicic magma types. The basaltic magma types are

81 divided into the northern high-Ti Pitanga, Paranapanema and Ribeira, the southern
82 low-Ti Gramado and Esmeralda, and the north-eastern high-Ti-Sr Urubici. Similarly,
83 the silicic magmas are divided into the high-Ti Chapecó, and the low-Ti Palmas
84 (Mantovani et al. 1985; Peate et al., 1992; Guimarães et al., 2019). Recent advances
85 on the understanding of the stratigraphy of the Paraná lavas have allowed for refining
86 these classifications (Polo et al., 2018; Rosseti et al., 2018). For instance, Rosseti et al.
87 (2018) proposed the division of the low-Ti Gramado and Esmeralda magma types,
88 from the base to top, into the Torres, Vale do Sol and Esmeralda Formations.
89 Moreover, it is also possible that the Esmeralda Formation is equivalent to the Barros
90 Cassal sequence, defined by Polo et al. (2018) as an andesitic basalt to dacite sequence
91 emplaced between the Vale do Sol Formation and the upper low-Ti silicic units of
92 rhyolitic composition.

93 Three main basaltic dyke swarms occur along the coastal areas of Brazil and
94 represent, at least partly, the plumbing systems that fed the magmas from the PEMP
95 (Floribal et al., 2014). These dyke swarms coeval to the Paraná lava flows (Janasi et
96 al., 2011; Pinto et al., 2011a; Floribal et al., 2014; Almeida et al., 2018) also show the
97 continuity of radiating structures in the African continent, allowing the genetic link
98 with the Etendeka margin counterparts (Marsh and Swart, 2018; Beccaluva et al.,
99 2020). From north to south the dyke swarms are the NE-SW Santos-Rio de Janeiro,
100 the NW-SE Ponta Grossa, and the nearly N-S Florianópolis (Fig. 1). The Ponta Grossa
101 dyke swarm is dominated by high-Ti basalts compositionally similar to the Pitanga
102 and Paranapanema magma types (Peate et al., 1992; Almeida et al., 2018). The
103 Florianópolis dyke swarm comprises high-Ti-Sr basalts similar to the Urubici magma
104 type (Floribal et al., 2014). The Santos-Rio de Janeiro dyke swarm is relatively more

105 complex and younger, comprising rocks with alkaline and tholeiitic affinities (Bennio
106 et al., 2003).

107

108 **3 – Sample selection and dataset**

109 The set of samples selected for analyses of PGE and TABS+ in this work
110 thought to be representative of the main varieties of basalts present in the PMP and
111 was based on work now under development. We selected 39 samples representative of
112 the different basaltic magma types from the Serra Geral Group and correlated dyke
113 swarms, which from now on we refer as the Paraná Magmatic Province (PMP; Fig. 1).
114 These comprise 8 samples from the high-Ti flows of the northern portion, 12 samples
115 from the low-Ti flows of the southern portion, 14 samples from the Ponta Grossa dyke
116 swarm, and 5 samples from the Florianópolis dyke swarm. All the samples from the
117 northern high-Ti flows represent the Pitanga magma type, whereas samples from the
118 southern low-Ti flows include the Vale do Sol Fm. (n=3), Torres Fm. (n=3), Esmeralda
119 Fm. (n=2) and Barros Cassal sequence (n=4). Samples from the Ponta Grossa dyke
120 swarm (sills, dykes and their corresponding volcanic flows) have geochemical
121 affinities with Pitanga (n= 8) and Paranapanema (n= 6) magma types, whereas those
122 from the Florianópolis dyke swarm present affinities with the Urubici type.

123 The PGE analyses of Rocha-Júnior et al. (2012), obtained using isotopic
124 dilution and allowing for very low detection limits, for the Esmeralda (n= 6),
125 Paranapanema (n= 3), Pitanga (n= 1) and Urubici (n=1) magma types, and the results
126 from Mincato (2000) which are not available in the English literature for the Gramado
127 (n= 16), Pitanga (n= 4), Paranapanema (n= 3) and Urubici (n= 2) types are included
128 for comparison. Results for PGE from Maier et al. (2003) and TABS+ from Barnes
129 and Mansur (*in press*) for the Horingbaai (n= 3) and Tafelberg (n= 3) members of the

130 Etendeka Province are also included. Other contributions have also investigated the
131 distribution of PGE (or at least some of them) in rocks from the PMP (Romanini and
132 de Albuquerque, 2001; Crockett, 2002; Pinto et al., 2011b; Arena et al., 2014).
133 Although some of these contributions did not focus specifically on the controls on the
134 distribution of PGE, we have also compared their results with our values in order to
135 check for major discrepancies.

136

137 **4 – Analytical methods**

138

139 The analyses of whole rocks by X-ray fluorescence and ICP-MS listed in
140 Electronic Supplementary Material (ESM, Table 1) were kindly provided by
141 colleagues, whose names are indicated in ESM (Table 1) and are part of datasets from
142 manuscripts in preparation.

143 Tellurium, As, Bi, Sb and Se analyses were carried out by Hydride Generation-
144 Atomic Fluorescence Spectrometry (HG-AFS) following the technique described by
145 Mansur et al. (2020b), at LabMaTer, Université du Québec à Chicoutimi (UQAC).
146 International reference materials (CH-4 and TDB-1 from Natural Resources Canada;
147 OKUM from IAGEO), and a blank were determined at the same time as the samples,
148 and the results agree with working values (ESM, Table 2). The 3σ detection limits are
149 0.006, 0.003, 0.007, 0.009 and 0.003 ppm for Te, As, Bi, Sb and Se, respectively.

150 Platinum-group elements (Os, Ir, Ru, Rh, Pd, and Pt) and S were also analyzed
151 at LabMaTer, UQAC. The PGE were pre-concentrated by Ni-sulfide fire assay and co-
152 precipitated with Te, and analyzed by solution inductively coupled plasma – mass
153 spectrometry (ICP-MS) using the method described by Savard et al. (2010). Sulfur was
154 determined by combustion and infrared analysis using a HORIBA EMIA-220V

155 analyser using the method described by Bédard et al. (2008). International reference
156 materials (OKUM and GeoPT-18/KPT-1, from IAGEO; LK-NIP-1 and LDI-1, from
157 Geo Labs - OGS) were analyzed with the samples to monitor data quality. The results
158 obtained for the reference materials are consistent with the working values (ESM,
159 Table 3).

160 Whole-rock major oxides and some trace-elements were analyzed by X-ray
161 Fluorescence at the Geonalitica Core Center, Instituto de Geociências, Universidade
162 de São Paulo, Brazil, following the protocol described in Mori et al. (1999). Analyses
163 were performed in a Panalytical AXIOS MAX Advanced spectrometer. Major and
164 minor elements were analyzed from molten tablets obtained from 1g sample and 9g of
165 a mixture of spectroscopic grade lithium tetra and metaborate at 1100-1200° C in a
166 platinum crucible, using a Claisse fusion machine. Trace elements were analyzed from
167 pressed pellets obtained from homogenization of the pulverized sample (previously
168 micronized in a McCrone Micronizer) with wax. Analytical quality was controlled
169 using reference materials JB-1a and JG-1a as unknown samples, and duplicate samples
170 (ESM, Table 4).

171 Rare earth and other trace elements were determined by ICPMS at the
172 Geonalitica Core Center, Instituto de Geociências, Universidade de São Paulo, Brazil,
173 following the protocol described in Navarro et al (2008). Analyses were performed in
174 a Thermo-Analitica iCAP quadrupole analyzer. Powdered samples were dissolved by
175 acid attack with concentrated HF and HNO₃ in microwave oven (aphanitic and fine-
176 grained volcanic rocks) or in Parr bombs (medium-grained intrusive rocks). Analytical
177 quality control was checked by the analyses of reference materials JG-1a and JR-1 as
178 unknown samples, blank and duplicate samples (ESM, Table 4).

179

180 **5 – Results**

181

182 **5.1 – Fractionation of the magmas**

183

184 The rocks from the PMP are classified as high-Ti and low-Ti magma series
185 (Mantovani et al. 1985; Peate et al., 1992; Guimarães et al., 2019). Samples from the
186 high-Ti series have Ti concentrations greater than 10,000 ppm and Ti/Y ratios above
187 350, whereas most low-Ti rocks have Ti contents lower than 10,000 ppm and Ti/Y
188 ratios ranging from 200 to 400 (Fig. 2a and 2b). Moreover, high-Ti rocks have higher
189 La/Yb and Sm/Yb ratios, ranging from 7 to 13 and 2 to 4, respectively, than low-Ti
190 rocks in which these range from 2 to 10 and 1 to 2.5, respectively (Fig. 2c and 2d).
191 The Urubici type magmas have similar Ti contents and Ti/Y ratios to high-Ti rocks,
192 but even higher La/Yb and Sm/Yb ratios, ranging from 15 to 18 and 3.8 to 4.5,
193 respectively (Fig 2).

194 The distribution of Mg# (i.e. ($[Mg]/[Mg+Fe] \times 100$)) and other major and
195 minor elements (Al_2O_3 , Fe_2O_3 , CaO, Ni, Cr and Cu) are shown in Fig. 3. Samples from
196 the high-Ti and Urubici varieties have lower Mg# relative to low-Ti magma type,
197 whereas the highest Mg# are found in picrites from the Etendeka Province (Fig. 3).
198 There is a positive correlation between Al_2O_3 , CaO, Ni and Cr with Mg#, and a
199 negative correlation between Fe_2O_{3T} and Cu with Mg# (Fig. 3). These variations
200 suggest progressive crystallization of clinopyroxene and plagioclase during
201 fractionation of these liquids. Also, the increase in Al_2O_3 and CaO and decrease in
202 Fe_2O_3 from Etendeka picrites into the more primitive magma composition (i.e. higher
203 Mg#) from the PMP support the early olivine removal from these liquids. The decrease
204 in Cr values with progressive fractionation may reflect the crystallization of accessory
205 chromite, especially from more primitive magmas. The Cu values increase with

206 decreasing Mg#, thus reflecting its incompatible behaviour during silicate
207 crystallization. However, there are a few samples at Mg# close to 35 that display Cu
208 depletion, thus supporting that fractional crystallization was at least partially
209 accompanied by sulfide removal.

210 The results for PGE are broadly in agreement with previously published values
211 (Mincato, 2000; Crockett, 2002; Romanini and de Albuquerque, 2001; Pinto et al.,
212 2011b; Rocha-Júnior et al., 2012; Arena et al., 2014). Given that no major
213 discrepancies were found between our results and literature values, and that we do not
214 aim to provide a review of PGE results within PMP, we acknowledge previous studies
215 but only include the data from Mincato (2000) and Rocha-Júnior et al. (2012) in our
216 plots. These contributions provide representative examples from literature values and
217 cover a similar sampling extension to our samples. Platinum and Pd concentrations
218 vary from 3 to 20 ppb and although scattered, there is a slightly increase in
219 concentrations following fractional crystallization (Fig. 4a and 4b). The exceptions are
220 a few high-Ti rocks from the Pitanga magma type located at the northern portion of
221 the PMP, which have Pt and Pd concentrations below 1 ppb, most likely reflecting the
222 effect of sulfide removal. Rhodium and Ru concentrations are mostly below 0.7 and
223 0.3 ppb, respectively, and decrease during fractional crystallization (Fig. 4c and 4d).
224 Many of our samples have Ir, and especially Os, concentrations below their detection
225 limits, and most samples with detectable amounts have less than 0.5 and 0.2 ppb,
226 respectively (Fig. 4e and 4f). The exceptions are picrites from the Etendeka Province,
227 which have Ir and Os concentrations ranging from 0.5 to 2 ppb. Thus, Ir and Os
228 concentrations decrease following progressively fractional crystallization (Fig. 4e and
229 4f).

230 Concentrations of Se, Te and Bi do not correlate with Mg# (Fig. 5). Selenium
231 and Te concentrations fall mainly below 0.1 and 0.01 ppm, respectively, with the
232 exception of dykes from the Ponta Grossa dyke swarm and the Etendeka Province,
233 which have slightly greater contents (Fig. 5a and 5b). Bismuth concentrations are more
234 widespread, and although concentrations in samples from the Ponta Grossa dyke
235 swarm are also among the highest values, there is a vast overlap with other samples
236 (Fig. 5c). It is noteworthy that Se, Te and Bi concentrations found in some samples
237 from the PMP, especially from the Ponta Grossa dyke swarm, are among the highest
238 values compared to other LIP (Fig. 5a to 5c), which could be related to the more
239 fractionated character of rocks from the PMP. In contrast to other TABS+, As and
240 especially Sb, show an incompatible behaviour, and concentrations increase following
241 progressively lower Mg#. Arsenic concentrations vary mainly from 0.05 to 0.5 ppm,
242 whereas Sb values vary from 0.03 to 0.2 ppm (Fig. 5d and 5e). The increase in Sb
243 concentrations is also followed by other incompatible elements (e.g. Nb in Fig. 5f),
244 which support its incompatible behaviour during fractional crystallization.

245

246 **5.2 – The PGE distribution**

247 Samples from both high- and low-Ti types from the PMP display highly
248 fractionated mantle-normalized PGE patterns, with strong IPGE (Ir, Os and Ru) and
249 Rh depletion relative to Pt and Pd (Fig. 6). A few high-Ti samples are also depleted in
250 Pt and Pd relative to Cu (Fig. 6a), and this feature is also observed in the data from
251 Mincato (2000). Slightly negative Ru anomalies are observed in few mantle-
252 normalized patterns from both high- and low-Ti rocks. There is no major variation in
253 mantle-normalized patterns from different formations within low-Ti rocks (Fig. 6b).
254 In contrast, the Ponta Grossa dyke swarm patterns for the Paranapanema type magmas
255 are slightly enriched relative to those for the Pitanga type magmas (Fig. 6c). Samples

256 from the Florianópolis dyke swarm also display IPGE-depleted patterns, but no clear
257 negative Ru anomalies (Fig. 6d).

258 Platinum and Pd show a broad positive correlation, but the Pt/Pd ratios are variable
259 within the PMP (Fig. 7a). The Pt/Pd ratios of our dataset vary from 0.5 to 2 with no
260 clear distinction between high-Ti and low-Ti rocks. A clear distinction, however, is
261 observed among the high-Ti basalts, with our Paranapanema samples showing lower
262 Pt/Pd (mostly <1), which reflect high Pd contents (>10 ppb), whereas Pitanga spreads
263 over a roughly trend at higher (1.5-2) Pt/Pd. High Pd contents are also shown by the
264 Paranapanema samples analysed by Mincato (2000), that have Pt/Pd ratios even lower
265 than ours, but not by the Paranapanema flows analysed by Rocha-Júnior et al. (2012)
266 (Fig. 7a). Therefore, it is not possible to clearly separate high-Ti and low-Ti rocks
267 based only on their Pt/Pd ratios. Palladium also shows a moderate positive correlation
268 with Rh (Fig. 7b), but no correlation with Ir. Palladium/Ir ratios vary widely within the
269 PMP, ranging from ~ 25 to 300 in both low-Ti and high-Ti rock types, however, for
270 most of the samples the ratio varies from 50 to 200 (Fig. 7c). Rhodium shows no clear
271 correlation with IPGE (Fig. 7d), whereas Ru and Os display a positive correlation (Fig.
272 7e). Iridium and Os show no correlation but is noteworthy that concentrations of both
273 elements are close to their detection limits (Fig. 7f).

274

275 **5.3 – The TABS+ distribution**

276

277 The distribution of TABS+ can be readily assessed using the mantle-normalized
278 diagram (Fig. 8), proposed by Barnes and Mansur (*in press*). The use of the diagram is
279 detailed by the authors and only a summary is provided here. In order to assess the effect
280 of sulfide liquid segregation, or sulfide liquid accumulation, the elements are arranged in

281 order of partition coefficient into a sulfide liquid (Liu and Brenan, 2015). To take into
282 consideration the effects of degree of partial melting and/or crystal fractionation Th, Nb
283 were added to the diagram as representatives of the lithophile incompatible elements. The
284 TABS+ and S are volatile (Lodders, 2003), and therefore subaerial lavas could be
285 depleted in these elements during degassing. In order to take this into consideration,
286 chalcophile elements that are less volatile (Cu, Pd, and Pt) were added to the plot at the
287 position of their relative partition coefficients into base metal sulfide liquid.

288 Most of the samples show similar patterns with overall negative slopes
289 decreasing from Th and Nb mantle-normalized values in the 50 to 200 range to Pd and
290 Pt values in the 1 to 3 range. A few high-Ti Pitanga magma type are the exceptions
291 and are more strongly depleted in Pt and Pd (Fig. 8a). Most samples also show negative
292 As, Se and Te anomalies. In samples from the Ponta Grossa dyke swarm these negative
293 anomalies are variable and tend to be more pronounced in volcanic rocks relative to
294 the subvolcanic ones (Fig. 8c). The patterns for samples from different localities do,
295 however, show differences in Th to Nb ratios. Most of the high-Ti rocks are not
296 enriched in Th (Fig. 8a, 8c and 8d), whereas low-Ti types are enriched in Th (Figs.
297 8b), and this Th enrichment is best explained by crustal assimilation (Peate and
298 Hawkesworth, 1996). Few low-Ti and most Urubici type lavas display slightly positive
299 Bi anomalies (Fig. 8b and 8d).

300

301 **6 – Discussion**

302

303 **6.1 - The effect of chromite and olivine crystallization**

304

305 The behaviour of PGE during the fractionation of mafic rocks has been
306 extensively considered in different geological contexts worldwide (Barnes et al., 1985;
307 Maier et al., 2003; Bockrath et al., 2004; Pitcher et al., 2009; Hughes et al., 2015).

308 More primitive lavas display relatively flat mantle-normalized PGE patterns. In
309 contrast, progressively more fractionated lavas show an IPGE (Ru, Ir and Os) depletion
310 relative to PPGE (Pt, Pd and Rh), and the PPGE/IPGE ratios tend to increase
311 antithetically to Mg# (Barnes et al., 1985; Day et al., 2013; Arguin et al., 2016).
312 Although sulfide saturation would be expected during fractional crystallization, sulfide
313 removal is not thought to be responsible for fractionating PPGE from IPGE, as both
314 have similar partition coefficients into the sulfide liquid (Mungall and Brenan, 2014;
315 Liu and Brenan, 2015). Thus, alternative processes must take place to explain the
316 preferential IPGE removal during fractional crystallization.

317 Some authors have highlighted the potential association of PGE-alloys (mainly
318 IPGE-alloys) and chromite grains in volcanic rocks (Locmelis et al., 2011; Arguin et
319 al., 2016). Indeed, experimental studies support crystallization of some PGE in the
320 boundary layer of chromite (Finnigan et al., 2008). Moreover, several authors also
321 demonstrate the preferential partition of IPGE not only into chromite, but also into
322 olivine (Righter et al., 2004; Pagé et al., 2012; Arguin et al., 2016; Park et al., 2017).
323 Therefore, the early olivine and chromite removal from primitive magmas could
324 explain the IPGE-depleted patterns observed in more fractionated rocks. For instance,
325 the rocks from the PMP fall in the lowermost depleted end of a positive correlated
326 trend observed between Cr and IPGE for magmas from different LIPs (represented by
327 Ir in Fig. 9a). This suggests that chromite removal during early fractional
328 crystallization may be at least partially accountable for the IPGE depletion observed
329 in the PMP rocks.

330 In contrast to other LIPs, there is a lack of the more primitive end members in
331 the continental flood basalts from the PMP. Although picrites are observed in the
332 Etendeka Province (Thompson et al., 2001; Maier et al., 2003; Jennings et al., 2019),

333 these primitive rocks are not found in the South American counterpart of the province
334 (Peate et al., 1992; Rosseti et al., 2018; Beccaluva et al., 2020). However, modelling
335 results reveal that the early olivine crystallization, and probably chromite, have
336 occurred during fractionation of the PMP magmas (Jennings et al., 2019; Beccaluva et
337 al., 2020). Therefore, it is likely that the IPGE depletion observed in the PMP magmas
338 reflects the formation of IPGE-alloys in response to chromite crystallization, or also
339 the partition of Rh and IPGE into chromite and olivine lattice upon early fractionation.
340 Thus, the progressive fractionation would lead to progressively lower Rh and IPGE
341 concentrations (Fig. 9a), but would not decrease Pt and Pd, leading to progressively
342 higher Pd/Ir ratios (Fig. 9b). Also, higher Rh and IPGE values, accompanied by higher
343 Cr contents in picrites from the Etendeka Province (e.g. Ir in Fig. 9a), and lower Pd/Ir
344 ratios (Fig. 9b), provide additional support for the role of chromite in removing IPGE
345 during fractional crystallization.

346 The shapes of mantle-normalized PGE patterns also argue for IPGE removal
347 in response to chromite crystallization. Arguin et al. (2016) estimated the partition
348 coefficients for Rh and IPGE between chromite/melt and demonstrated that Ru has a
349 higher partition coefficient into chromite relative to Rh, Ir and Os. The authors then
350 argue that chromite crystallization would lead to a more pronounced depletion of Ru
351 relative to Rh, Ir and Os in the melt. This would be recorded by negative Ru anomalies
352 in mantle-normalized PGE patterns from more fractionated magma compositions
353 (Park et al., 2017). Our results support this interpretation, as some of the IPGE-
354 depleted mantle-normalized patterns display negative Ru anomalies (Fig. 6).
355 Therefore, we argue that the decoupling of PPGE and IPGE in the rocks from the PMP
356 was likely caused by the crystallization of chromite, and olivine, during early
357 fractionation of the PMP magma types. This interpretation is also complimentary to

358 the conclusions drawn by Day et al. (2013). The authors argued that fractionated
359 mantle-normalized PGE patterns like those found at the PMP were caused by the
360 crystallization of IPGE-bearing alloys during magma fractionation.

361
362 **6.2 -The effect of sulfide segregation**
363

364 Given the chalcophile nature of TABS+ and PGE, sulfide segregation or
365 accumulation are among the most important processes affecting their distribution
366 (Patten et al., 2013; Mungall and Brenan, 2014; Liu and Brenan, 2015; Mansur et al.,
367 2020b). Although S contents are below detection limits for most samples, there is a
368 positive correlation between S and Se for samples with contents above detection limits.
369 These samples are mainly from dykes and sills of the Ponta Grossa and Florianópolis
370 dyke swarms, and also dykes from the Etendeka Province (Fig. 10a). For these
371 samples, S/Se ratios vary from 1200 to 3400 (one sample yields 4600), which are close
372 to or below mantle values (i.e. approximately 3050; Lyubetskaya and Korenaga,
373 2007). Sulfur/Se ratios greater than mantle values are argued to result from S
374 assimilation from country rocks, whereas ratios lower than mantle values are indicative
375 of S loss (Queffurus and Barnes, 2015 and references therein). In the case of the PMP,
376 the fact that higher concentrations are found in subvolcanic rocks suggests loss by
377 degassing upon eruption, which display no detectable contents of both S and Se. The
378 effect of magma degassing on the distribution of chalcophile elements will be
379 considered in the next section below. Because the precision of the Se analyses (Mansur
380 et al., 2020b) is higher than of S analyses (Bédard et al., 2008), Se is used as a proxy
381 for S.

382 There is a broad positive correlation between Pd and Se for most samples, apart
383 from the high-Ti flows of Pitanga type (Fig. 10). Also, the Se and Pd concentration

384 ranges for samples from the PMP are similar to those observed in other LIPs. There is
385 also a broad positive correlation between Te and Bi and Se for the dyke samples from
386 the Ponta Grossa dyke swarm. The positive correlation between Pd, Te, Bi and Se is
387 in agreement with these elements being moderately to highly chalcophile, and thus
388 their distribution being controlled by the presence of sulfide minerals (Mansur and
389 Barnes, 2020a, 2020b). In contrast, As and Sb concentrations vary regardless of Se
390 contents, which is also in agreement with these elements being only slightly
391 chalcophile, and their distribution not being significantly affected by the presence of
392 sulfide minerals (Maciag and Brenan, 2020; Mansur and Barnes, 2020b). Also of note
393 is that the As and Sb concentrations in samples from the PMP are amongst the lowest
394 values found in other LIP. The concentrations of these elements in mafic magmas are
395 normally upgraded in response to crustal assimilation (Barnes and Mansur, *in press*),
396 which is counterintuitive given that crustal assimilation is supported by mantle-
397 normalized patterns for at least part of the rocks from the PMP (Fig. 8). One possibility
398 is that the contaminants at the PMP were not particularly enriched in As and Sb, but
399 further investigations on this topic are still necessary prior to more conclusive
400 interpretations.

401 Variations in Cu and PGE contents in some of the PMP rocks suggest that
402 sulfide segregation occurred during the fractionation of the magmas (Fig. 3f and 4). A
403 possible manner of tracking sulfide removal is by investigating Cu/Pd ratios in
404 different samples (Barnes et al., 1993; Maier et al., 1996). Both Cu and Pd are
405 considered incompatible elements during fractional crystallization in the absence of a
406 sulfide liquid, but once an immiscible sulfide liquid segregates both elements are
407 sequestered from the silicate magma. However, the partition coefficients between
408 sulfide and silicate liquid are higher for Pd than for Cu (Barnes et al., 1993; Maier et

409 al., 1996). Therefore, the depletion of Pd relative to Cu can be used as a tool for
410 estimating the degree of fractional crystallization under sulfide-saturated conditions
411 that a given magma underwent.

412 Figure 11 shows a binary plot of Cu/Pd versus Pd for the studied samples. We
413 have calculated a curve representing the crystallization under sulfide-saturated
414 conditions in cotectic proportions to estimate the amount of sulfide removal for
415 different samples. The used partition coefficients for Pd and Cu between silicate and
416 sulfide liquids were of 400,000 and 1,500, respectively (Mungall and Brenan, 2014).
417 The concentrations of both elements in the fractionated liquid (C_f) was calculated
418 using to the following equation:

419 Rayleigh fractionation:

$$420 \quad C_f = C_i F^{[(D_{Sul} F_{Sul}) - 1]}$$

421 where C_i is the concentration of the element in the initial liquid, F is the weight fraction
422 of remaining liquid, D_{Sul} is the partition coefficient of the element between silicate and
423 sulfide liquids, and F_{Sul} is the weight fraction of sulfide liquid in cotectic proportions.
424 We assumed the S content at sulfide saturation at about 1,500 ppm (Li and Ripley,
425 2005). This value and the average S content in magmatic sulfides (approximately 35
426 wt. %) were used to estimate the weight fraction of sulfide liquid in cotectic
427 proportions (approximately 0.004). Concentrations of Pd and Cu in the transporting
428 melt at sulfide saturation were assumed at 15 ppb and 160 ppm, respectively.

429 Most samples from the PMP show evidence for sulfide removal with Cu/Pd
430 ratios higher and Pd values lower than initial liquids, and mantle-derived magmas (Fig.
431 11). The Cu/Pd ratios and Pd values observed for most samples can be modelled by up

432 to 0.5% crystallization under with sulfide segregation in co-tectic proportions. Four
433 high-Ti samples do require a slightly more extensive sulfide removal, and
434 approximately 2.5% bulk crystallization under sulfide-saturated conditions in order to
435 reproduce their Cu/Pd and Pd values. This understanding is critical for exploration
436 programs given that PGE- and base metal-depleted volcanic rocks are associated to
437 large magmatic sulfide deposits in other LIP (Lightfoot et al., 1994; Jowitt and Ernst,
438 2013). Thus, our results support the sulfide removal at cotectic proportions during the
439 fractionation of most the PMP magma types, but only locally more extensive sulfide
440 removal has occurred. In fact, this was observed specifically in the Pitanga type high-
441 Ti magma compositions.

442 Another observation that is possibly explained by the sulfide removal from
443 rocks of the PMP is the negative slope found in mantle-normalized Th, Nb, Cu,
444 TABS+, Pt, Pd spidergrams (Fig. 8). In these plots, the elements are organized from
445 left to right following progressively higher partition coefficients into an immiscible
446 sulfide liquid, following Barnes and Mansur (*in press*). Therefore, given the cotectic
447 sulfide removal from PMP, the negative slopes of the patterns probably reflect the
448 preferential removal of highly chalcophile elements relative to slightly chalcophile
449 ones. However, although the negative slopes may be explained by sulfide removal, As,
450 Se and Te negative anomalies cannot be explained by different partition coefficients
451 of these elements into an immiscible sulfide liquid. For instance, Se and Te have lower
452 partition coefficients into a sulfide liquid than both Pt and Pd, thus their negative
453 anomalies cannot be explained by sulfide removal. Therefore, some additional process
454 is needed to explain these recurrent negative anomalies found in rocks from the PMP.

455

456 **6.3 -The effect of low-pressure degassing upon eruption**

457
458 Understanding the distribution of volatile elements in subaerial lavas is not
459 straightforward given that degassing under lower pressure eruption may diminish their
460 original concentrations. Because TABS+ have relatively low condensation
461 temperatures during the formation of the solar system they are classified as volatile
462 elements (Lodders, 2003). Moreover, the presence of TABS+, together with S, in
463 volcanic gases is also supportive of their loss upon eruption (Self et al., 2005; Zelenski
464 et al., 2014; Forrest et al., 2017; Wieser et al., 2020). The presence of negative As, Se
465 and Te anomalies in mantle-normalized patterns could result from the loss of these
466 elements upon degassing. The size of the anomalies based on neighbour elements in
467 mantle-normalized plots suggests that more than 80% of these elements were
468 potentially lost (relative difference from dashed black line in figure 8). An interesting
469 observation is that subvolcanic rocks from the Ponta Grossa dyke swarm show less
470 pronounced negative anomalies, which is suggestive that degassing was less intense in
471 these rocks.

472 Barnes and Mansur (*in press*) have compared the distribution of TABS+ in
473 subaerial rocks from different LIPs and in subaqueous rocks, such as MORB. Similar
474 negative As, Se and Te anomalies were observed in mantle-normalized plots for rocks
475 from various LIPs but were not present in MORB. The authors used TABS+/Cu ratios
476 to estimate the effect of degassing because Cu is less volatile than TABS+, and thus
477 lower ratios are suggestive of degassing. This assumption is supported by the presence
478 of native Cu in some subaerial lavas argued to have experienced S loss (Harris, 1987;
479 Arguin et al., 2016; Li and Boudreau, 2017). Moreover, previous studies also highlight
480 the presence of native copper in rocks from the PMP, thus supporting a similar process
481 (Pinto et al., 2011b; Baggio et al., 2018).

482 The Se/Cu and Te/Cu ratios in the PMP samples are mostly lower than mantle
483 values (Fig. 12a and 12b), which supports that Se and Te have been lost by degassing.
484 Also, the values are in the same range of those found at other subaerial lavas from
485 different LIP. It is also noteworthy that Se/Cu ratios in these samples are significantly
486 lower than those found in MORB, for which degassing is diminished by greater
487 eruption depths (Mathez, 1976; Lissner et al., 2014). Tellurium/Cu ratios in MORB
488 are also lower than mantle values, but this does not necessarily reflect Te loss by
489 degassing (Fig. 12b). Tellurium has a greater partition coefficient into sulfide liquid
490 than Cu (Patten et al., 2013; Liu and Brenan, 2015), and therefore sulfide removal
491 could lead to lower Te/Cu ratios. In the case of the PMP, however, sulfide removal
492 would not account for lower Te/Cu ratios in most samples because Te/Pd ratios are
493 lower than mantle. Given that Te has a lower partition coefficient into the sulfide liquid
494 relative to Pd, sulfide removal should yield Te/Pd ratios greater than mantle values.
495 Therefore, Te/Cu ratios lower than mantle in the PMP rocks support Te loss upon
496 degassing.

497 The Bi/Cu ratios show a wide range, but most values for the PMP are close to,
498 or slightly greater than mantle values (Fig. 12c). Barnes and Mansur (*in press*) also
499 report a wide variation in Bi/Cu ratios in various LIP. The authors argue that Bi
500 distribution reflects the overprinted effect of crustal assimilation, sulfide removal or
501 accumulation and potentially loss by degassing. Consequently, Bi/Cu ratios may not
502 directly record the effect of degassing. High- and low-Ti rock varieties show different
503 degrees of crustal assimilation (Peate et al., 1992; Rocha-Júnior et al., 2013; Rosseti
504 et al., 2018; Beccaluva et al., 2020), however similar Bi/Cu ratios suggest no major
505 influence from these processes in Bi concentrations at the PMP. The reasons for the

506 scattering in Bi/Cu ratios observed at the PMP, and also other LIPs, are not fully
507 understood at this time.

508 Because As and Sb are only slightly chalcophile elements and show an
509 incompatible behaviour, we used As/Th and Sb/Th ratios to consider the effect of
510 degassing on these elements. Moreover, Barnes and Mansur (*in press*) demonstrated
511 that As and Sb concentrations, and Th, may be variably upgraded by assimilation of
512 crustal material. Thus, lower ratios could reflect As and Sb loss by degassing, whereas
513 higher ratios could reflect the assimilation of different crustal components (e.g.
514 average upper crust, black shale). Samples from the PMP have As/Th and Sb/Th ratios
515 lower than mantle values (Fig. 12d and 12e). These values are lower than most results
516 found at other LIP, for which As/Th and Sb/Th ratios are mainly close to mantle values.
517 However, at other LIPs the values close to mantle are interpreted to reflect the
518 combined effect of As and Sb upgrading by crustal assimilation (Barnes and Mansur,
519 *in press*) and fractional crystallization (Maciag and Brennan, 2020), and their loss upon
520 degassing. An upgrade of As and Sb due to crustal assimilation in samples from the
521 PMP is not supported by our results. For instance, low-Ti volcanoc rocks have positive
522 Th mantle-normalized anomalies (Fig. 8b), supporting crustal assimilation, but no
523 upgrade in As or Sb concentrations. Thus, in the case of the PMP, ratios lower than
524 mantle values suggest As and Sb loss upon degassing.

525

526 **6.4 – Potential relationship between LIPs emplacement and mass extinctions**

527 Understanding the input of volatile elements into the atmosphere caused by the
528 emplacement of LIPs is a long-debated issue. One of the main motivations for these
529 studies is to assess either the volatile input may contribute to mass extinctions, which
530 were frequently coeval to the formation of various LIPs (e.g. Wingall, 2001).

531 However, although the association of LIP emplacement and mass extinction events is
532 common, no such correlation has been documented for the PMP yet. For instance,
533 Callegaro et al. (2014) argued that lower magmatic S contents in the PMP magmas
534 could explain the lower environmental impact relative to other LIPs associate to mass
535 extinctions (Wignall, 2001; Callegaro et al., 2014; Heimdal et al., 2019). Therefore,
536 understanding how the environmental impacts caused by the emplacement of the PMP
537 differ from those caused by the emplacement of other LIPs is critical to better
538 understand the link between LIPs and mass extinctions.

539 Our results do support the loss of S, Te, Se, As and Sb during the emplacement
540 of the PMP. These have been considered as toxic elements, and their upgrade in the
541 atmosphere may have an environmental impact (e.g. Vandenbroucke et al., 2015).
542 However, quantifying the amount of these elements released in the atmosphere is not
543 straightforward as initial concentrations are not readily assessed. Also, an additional
544 problem is that metal input during LIP emplacement may also derive from heating and
545 degassing of the host rocks (Ganino and Arndt, 2009; Heimdal et al., 2019), and not
546 only from magma degassing. One possibility for assessing initial concentrations of
547 TABS+ in volcanic rocks from different LIPs would be by investigating their
548 concentrations in melt inclusions (Wallace and Edmonds, 2011). We suggest that
549 future contributions in this domain would further help refining our understanding on
550 the link between magma degassing upon LIP emplacement and mass extinctions.

551

552 **7 – Conclusions**

553 We investigated the distribution of PGE and TABS+ in volcanic and
554 subvolcanic rocks from the Paraná Magmatic Province (PMP). Our main findings are
555 summarized as follows:

556 1 – Distributions of PGE and TABS+ in the PMP rocks allows assessing the effects of
557 fractional crystallization, sulfide segregation and magma degassing.

558 2 - Crystallization of olivine and chromite at early stages of magma fractionation led
559 to a depletion of Rh and IPGE in samples from the PMP. The effect of chromite
560 removal is also recorded by negative Ru anomalies in mantle-normalized PGE
561 patterns. Platinum and Pd are not affected by olivine and chromite crystallization, and
562 show an incompatible behaviour, thus becoming more concentrated during progressive
563 fractionation.

564 3 – Selenium, Te and Bi do not record the effect of fractional crystallization as their
565 distribution is overprinted by other processes. On the other hand, As and Sb are
566 incompatible, and their concentrations increase in progressively more fractionated
567 rocks.

568 4 – Most magma varieties from the PMP rocks display Cu/Pd ratios higher and Pd
569 contents lower than mantle values, which support up to 0.5% bulk crystallization under
570 sulfide-saturated conditions. However, a few high-Ti samples do require around 2.5%
571 crystallization under sulfide-saturated conditions to reproduce their Cu/Pd ratios and
572 Pd contents.

573 5 – Most lavas from the PMP display negative As, Se and Te anomalies on mantle-
574 normalized patterns, which coupled with Te/Cu, Se/Cu and As/Th and Sb/Th ratios
575 lower than mantle values, support their loss upon degassing. The effects of magma
576 degassing for these elements is less pronounced in subvolcanic rocks from the Ponta
577 Grossa and Florianópolis dyke swarms relative to subaerial lava flows.

578
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580
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Figure Captions

589 **Figure 1** – Simplified geological map illustrating the distribution of the main lavas in
590 the Paraná Magmatic Province. The black stars represent the location of the samples
591 from this study. Modified from Peate et al. (1992) and Janasi et al. (2011).
592 Abbreviations – Par: Paraná, Ete: Etendeka, Tr: Trindade, TC: Tristán da Cunha.

593 **Figure 2** – Binary plots of (a) Mg# vs Ti (ppm), (b) Mg# vs Ti/Y, (c) La/Yb vs Ti/Y
594 and (d) Sm/Yb vs Ti/Y for various samples from the PMP. Note the general
595 geochemical separation of the samples with higher Ti contents and Ti/Y ratios from
596 those with lower values. Whole-rock results are given in electronic materials (ESM,
597 Table 1).

598 **Figure 3** - Binary plots of (a) Al₂O₃, (b) Fe₂O₃, (c) CaO, (d) Ni, (e) Cr and (f) Cu
599 versus Mg# for various samples from the PMP. The fractionation trends are compatible
600 with the progressive crystallization of olivine, pyroxene, plagioclase and chromite.
601 The increase in Cu concentrations following fractionation is in agreement with its
602 incompatible behaviour, whereas lower concentrations support sulfide removal.
603 Whole-rock results are given in electronic materials (ESM, Table 1).

604 **Figure 4** - Binary plots of (a) Pt, (b) Pd, (c) Rh, (d) Ru, (e) Ir and (f) Os versus Mg#
605 for various samples from the PMP. The back arrows indicate the processes variably
606 affecting the distribution of PGE in the samples. Please see text for further explanation.
607 Dashed lines indicate the detection limits. Whole-rock results are given in electronic
608 materials (ESM, Table 1).

609 **Figure 5** - Binary plots of (a) Se, (b) Te, (c) Bi, (d) As and (e) Sb versus Mg# for
610 various samples from the PMP. Compositional fields of samples from other LIP and
611 MORB from Barnes and Mansur (*in press*) are shown for reference. There is no clear
612 correlation between Se, Te and Bi with Mg#, whereas Sb, and to a lesser extent As,
613 show a negative correlation with Mg#, thus supporting their incompatible behaviour
614 during fractionation. Dashed lines indicate the detection limits. Whole-rock results are
615 given in electronic materials (ESM, Table 1).

616 **Figure 6** - Primitive mantle-normalized Ni–PGE–Au–Cu patterns of samples from (a)
617 high-Ti flows, (b) low-Ti flows, (c) Ponta Grossa dyke swarm and (d) Florianópolis
618 dyke swarm. Compositional fields for results from Mincato (2000) and Rocha-Júnior
619 et al. (2012) are shown for reference. Primitive mantle values from Lyubetskaya and
620 Korenaga (2007). Dashed lines indicate the detection limits. Whole-rock results are
621 given in electronic materials (ESM, Table 1).

622 **Figure 7** - Binary plots of (a) Pt vs Pd, (b) Pd vs Rh, (c) Pd vs Ir, (d) Ru vs Rh, (e) Os
623 vs Ru and (f) Os vs Ir for various samples from the PMP. Black lines indicate different
624 Pt/Pd and Pd/Ir ratios on (a) and (c), respectively. Dashed lines indicate the detection
625 limits. Note that some previous analyses from literature have lower detection limits
626 than this study, and thus plot below dashed lines. Whole-rock results are given in
627 electronic materials (ESM, Table 1).

628 **Figure 8** - Primitive mantle-normalized Th, Nb, Cu, TABS+, Pd and Pt patterns of
629 samples from (a) high-Ti flows, (b) low-Ti flows, (c) Ponta Grossa dyke swarm and
630 (d) Florianópolis dyke swarm. Primitive mantle values from Lyubetskaya and
631 Korenaga (2007). Black dashed lines indicate the detection limits, whereas black full

632 lines indicate the connection between Nb and Pd. Whole-rock results are given in
633 electronic materials (ESM, Table 1).

634 **Figure 9** – Binary plots of (a) Ir versus Cr and (b) Pd/Ir versus Ir for samples from the
635 PMP. The gray X symbols are the results from Emeishan, Karoo, Siberian Traps from
636 Barnes and Mansur (*in press*). The black line indicates the fractionation trend by the
637 progressive crystallization of olivine and chromite leading to lower Ir contents and
638 higher Pd/Ir ratios accompanied by lower Cr contents. Primitive mantle values (blue
639 diamond) from Lyubetskaya and Korenaga (2007). Dashed line indicates the detection
640 limit for Ir.

641 **Figure 10** - Binary plots of (a) S vs Se, (b) Pd, (c) Te, (d) Bi and (e) Sb versus Se and
642 (f) As vs Sb for various samples from the PMP. Compositional fields of samples from
643 other LIP and MORB from Barnes and Mansur (*in press*) are shown for reference.
644 Black line represents primitive mantle ratio (Lyubetskaya and Korenaga, 2007) and
645 dashed lines indicate the detection limits. Whole-rock results are given in electronic
646 materials (ESM, Table 1).

647 **Figure 11** - Binary plot of Cu/Pd versus Pd for samples from the PMP. The black line
648 corresponds to the product of bulk crystallization with sulfide liquid removal in
649 cotectic proportions from an initial silicate liquid containing 15 ppb Pd and 160 ppm
650 Cu. We used the partition coefficients for Pd and Cu between silicate and sulfide liquid
651 of 400000 and 1500, respectively (Mungall and Brenan, 2014). Please see text for
652 further explanation.

653 **Figure 12** - Binary plots of (a) Se, (b) Te, (c) Bi versus Cu and (d) As and (e) Sb versus
654 Th for various samples from the PMP. Compositional fields of samples from other LIP
655 and MORB from Barnes and Mansur (*in press*) are shown for reference. Black lines

656 represent primitive mantle ratio (Lyubetskaya and Korenaga, 2007) and dashed lines
657 indicate the detection limits. Please see text for further explanation on the meaning of
658 the ratios.

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Electronic supplementary materials – Data availability

Table 1 - Whole-rock results obtained in this study.

663 **Table 2** - Analyses of reference materials used to monitor the data quality of HG-AFS
664 analyses.

665 **Table 3** - Analyses of reference materials used to monitor the data quality of PGE, Au
666 and S analyses.

667 **Table 4** - Analyses of reference materials used to monitor the data quality of whole-
668 rock analyses of major oxides and trace elements.

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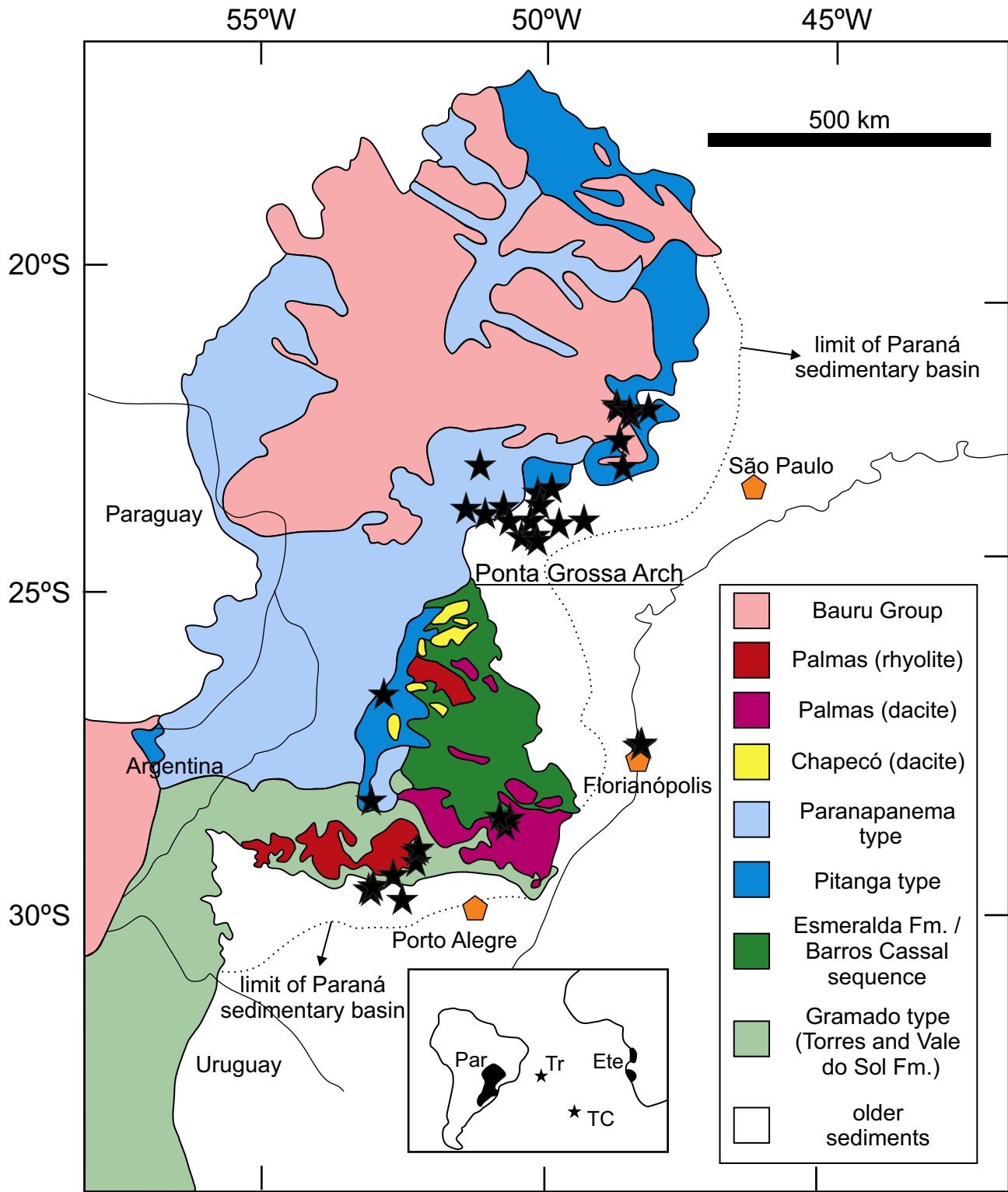
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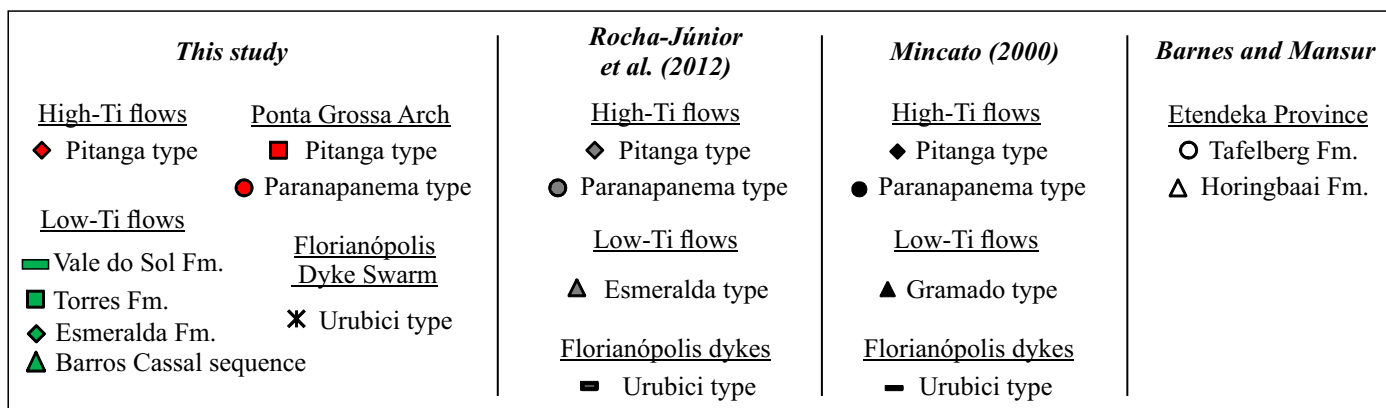
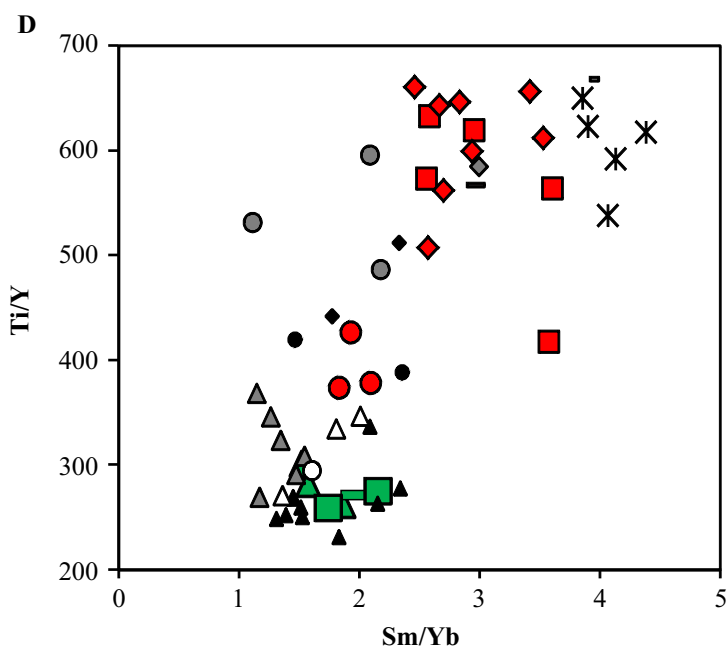
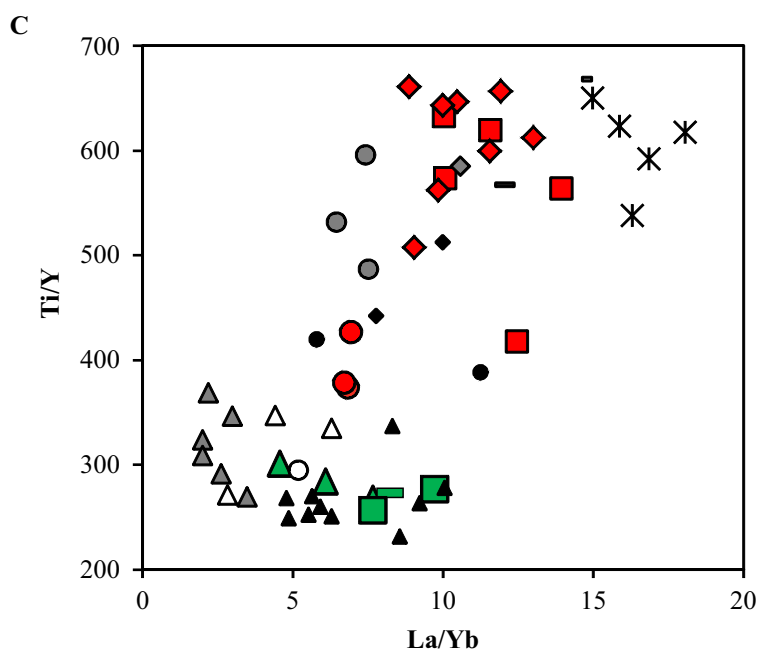
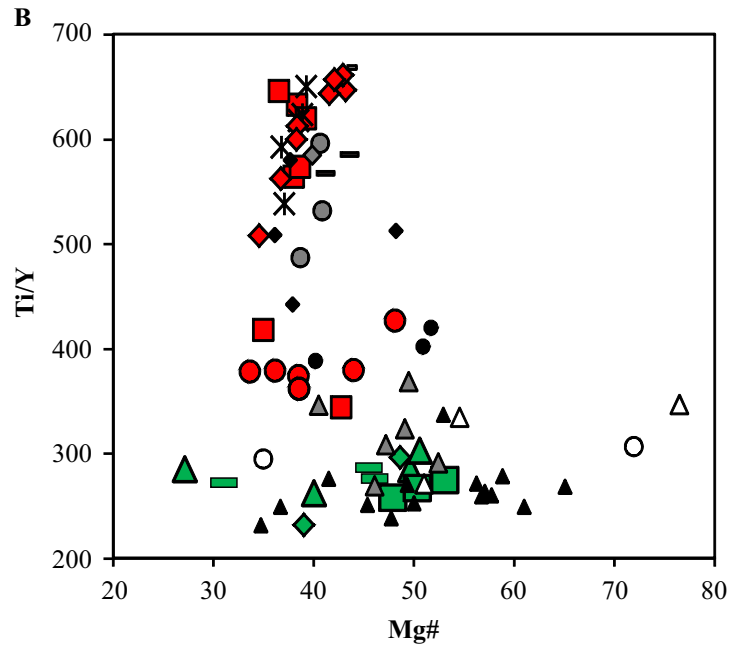
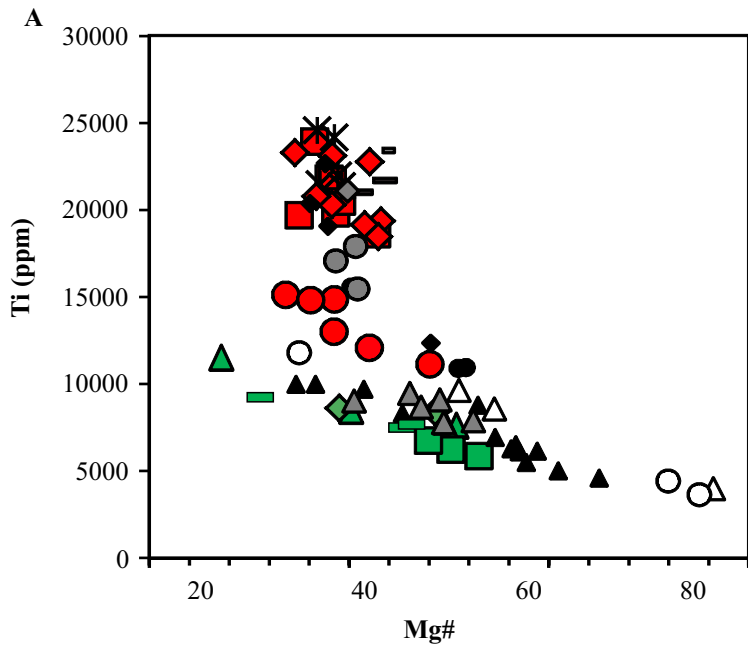
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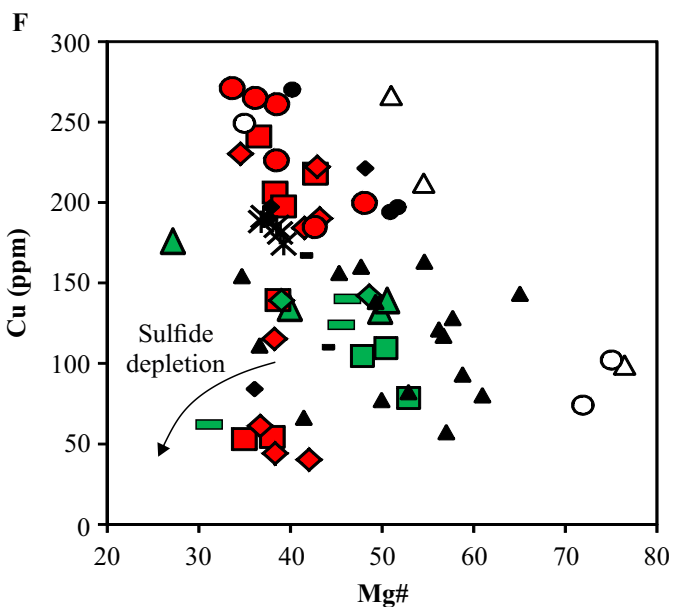
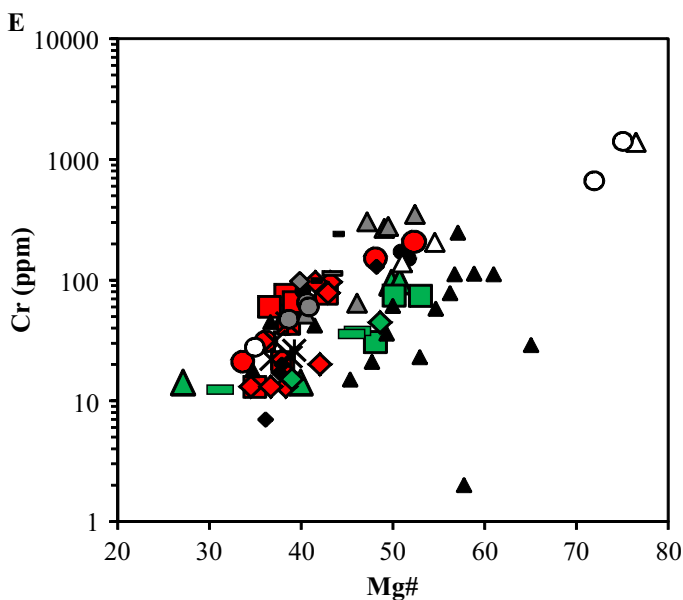
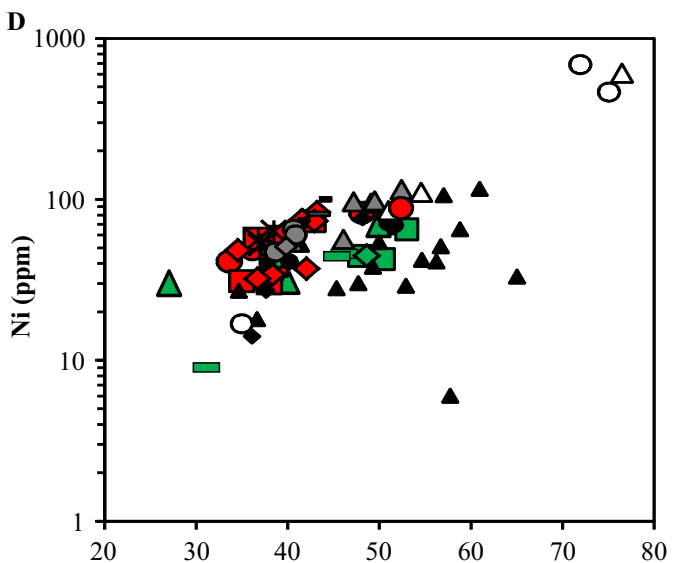
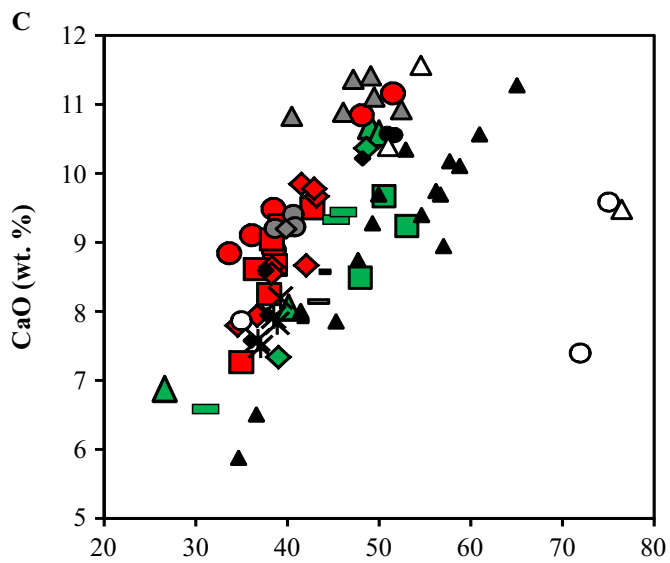
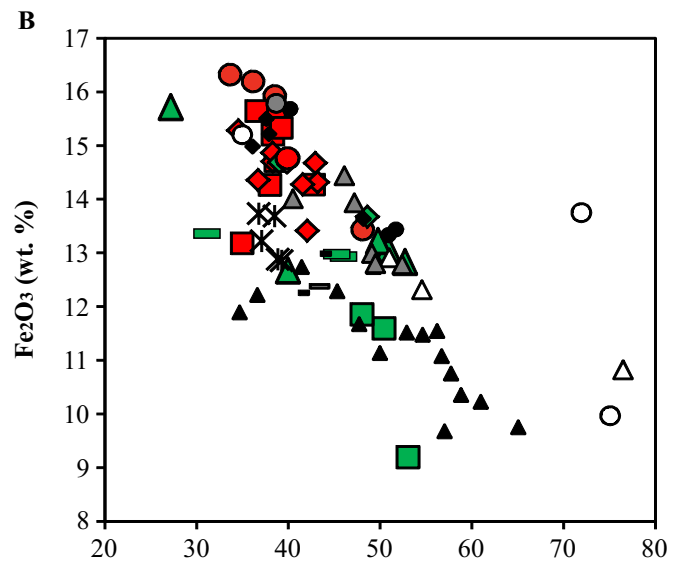
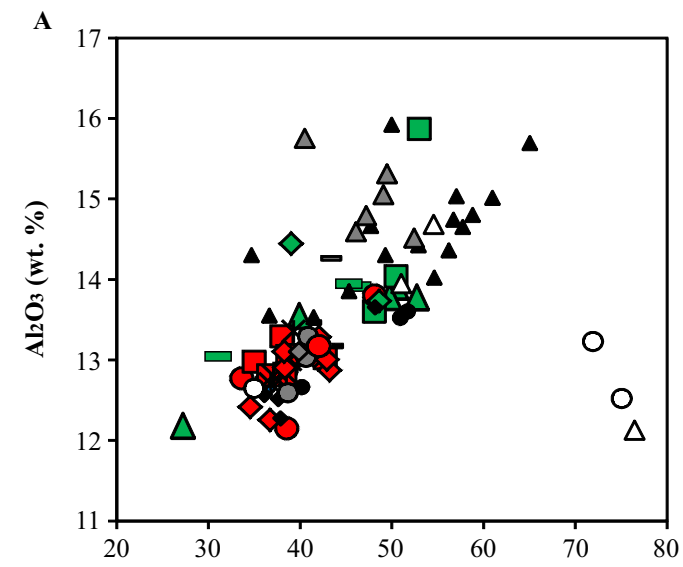
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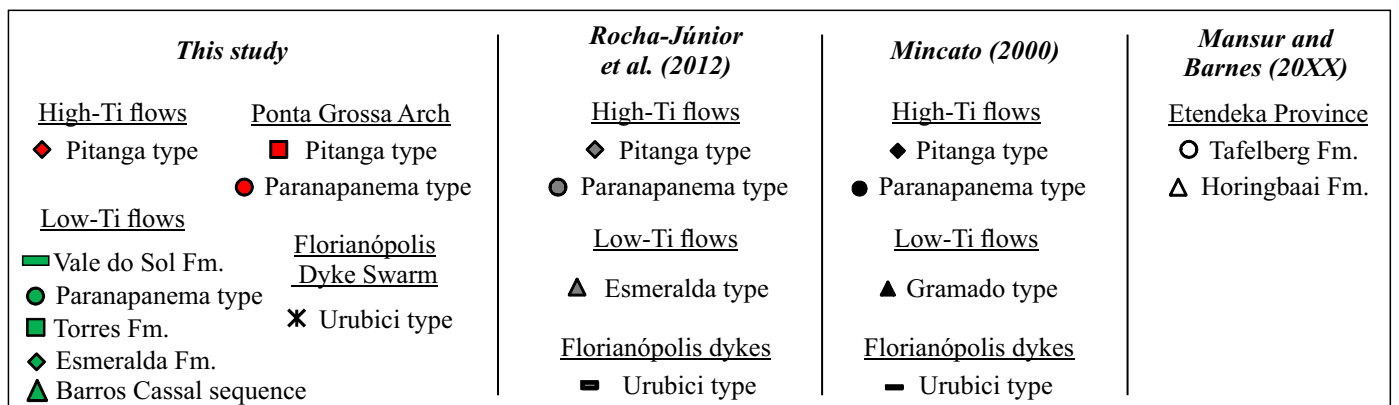
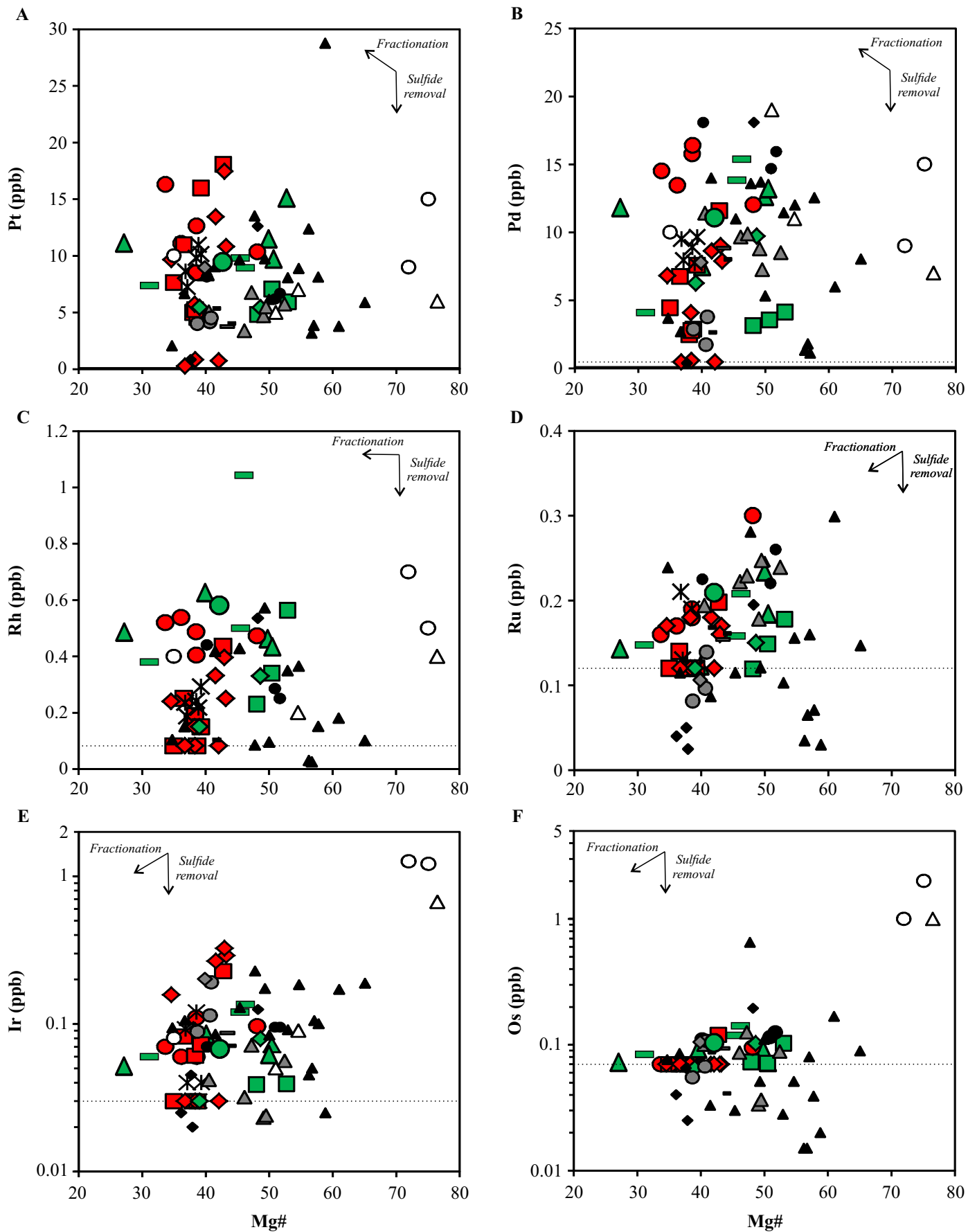
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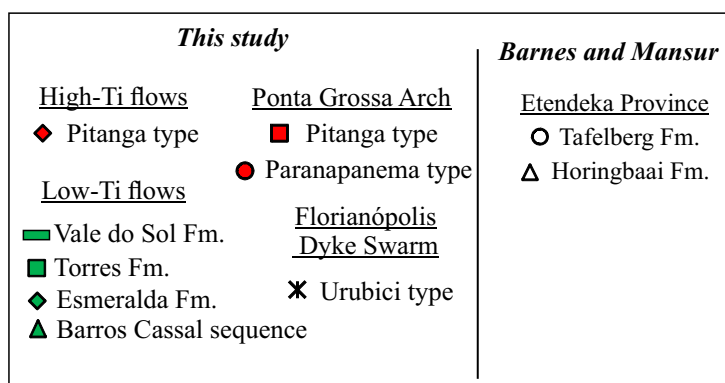
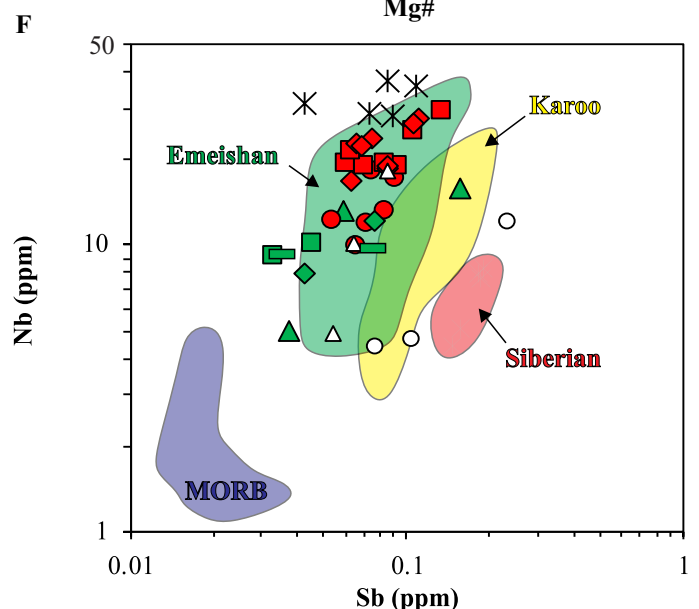
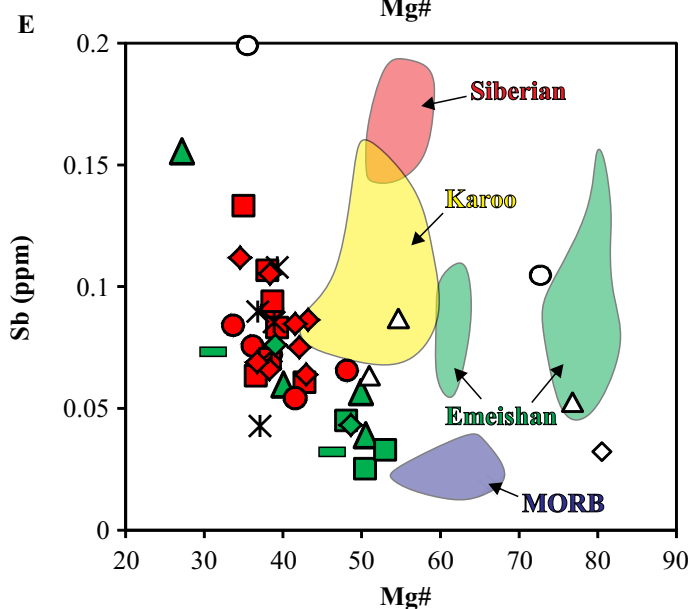
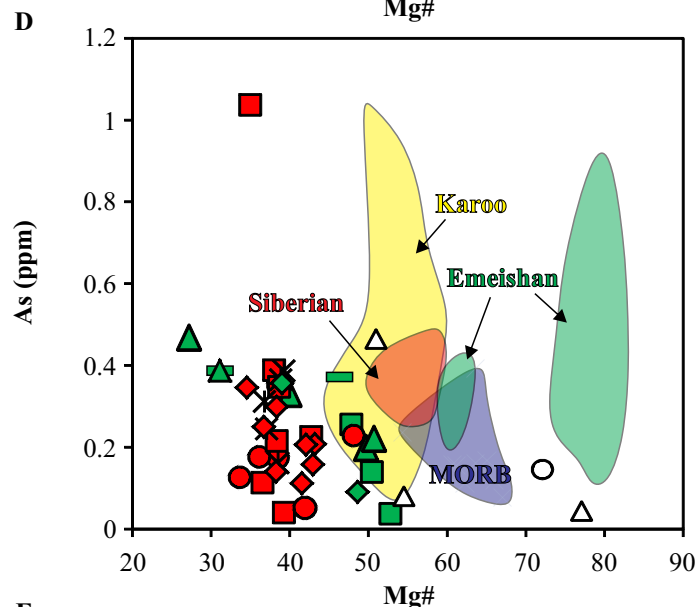
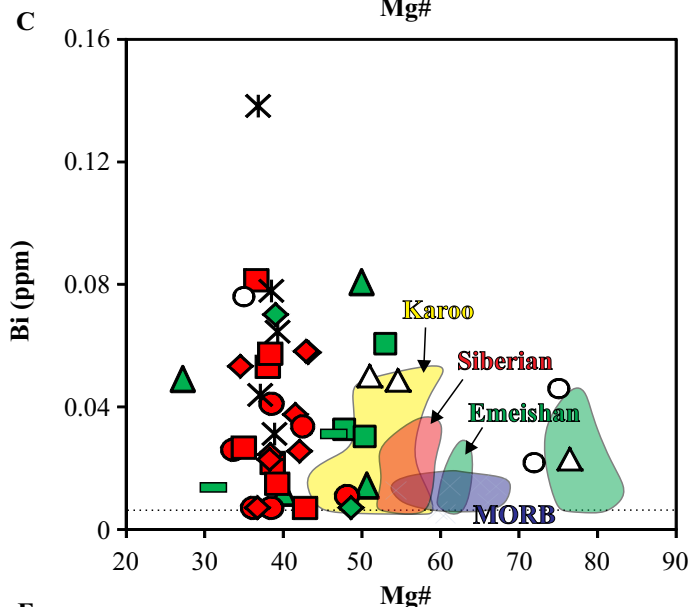
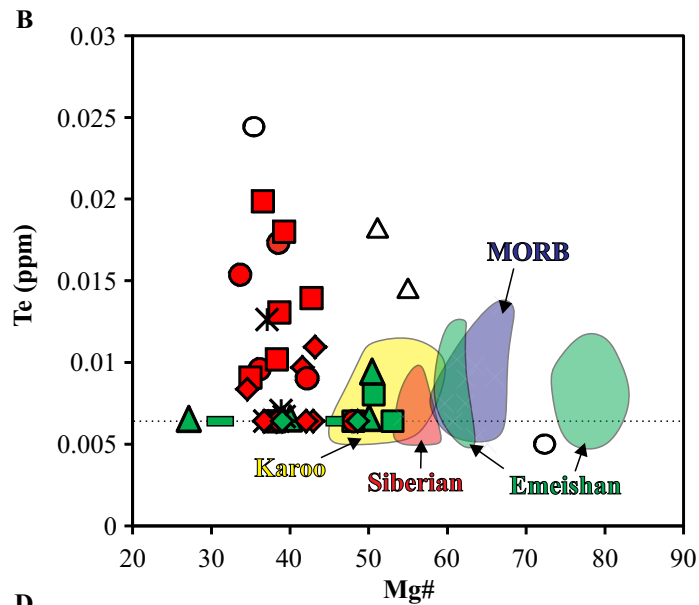
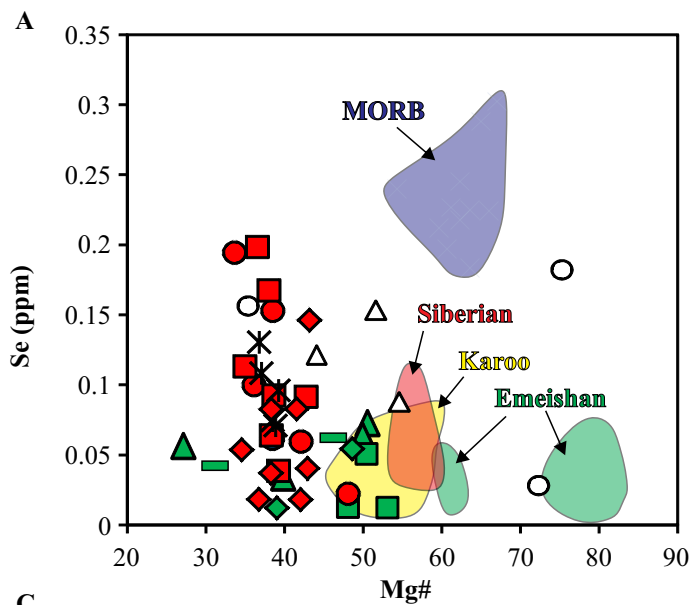


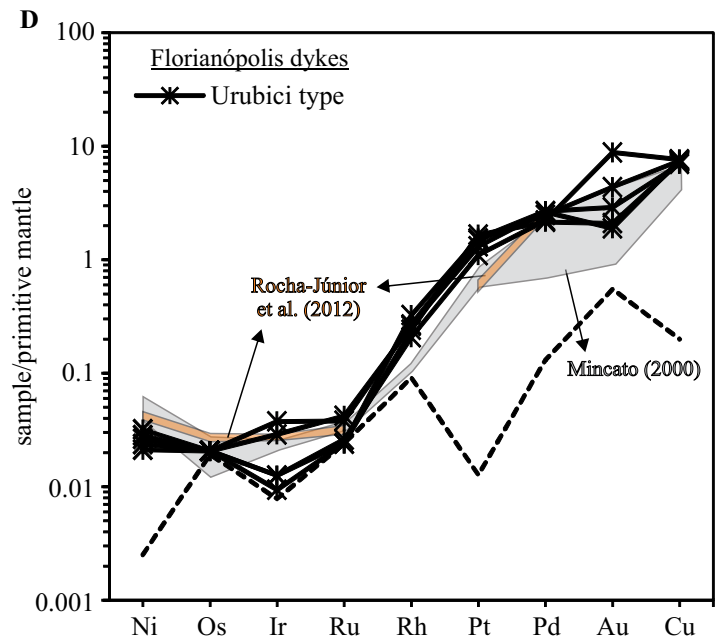
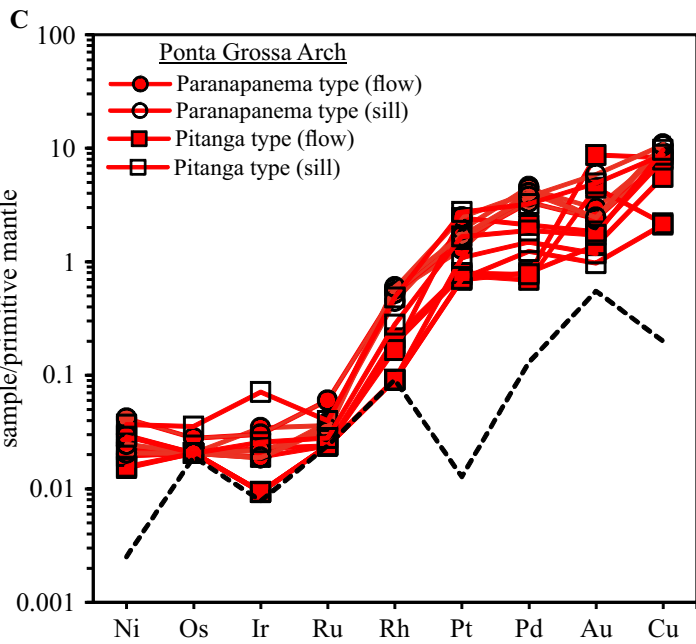
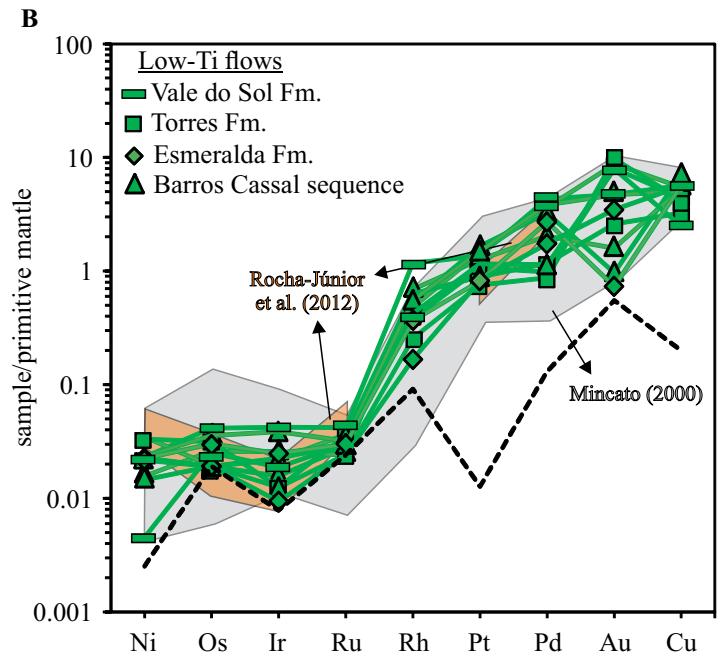
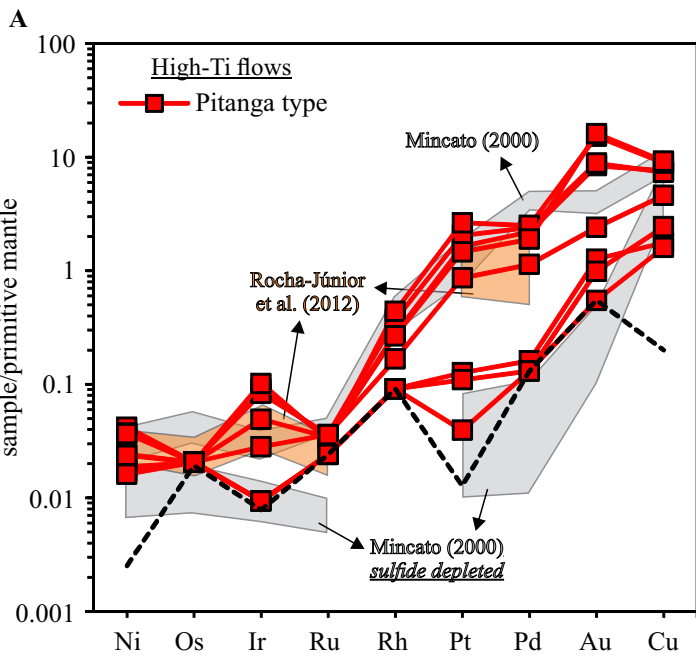


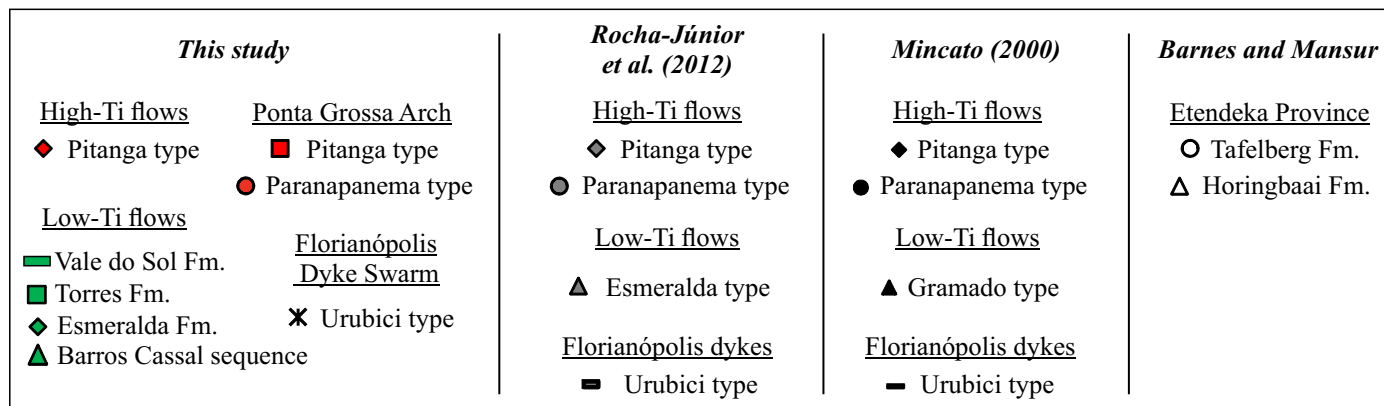
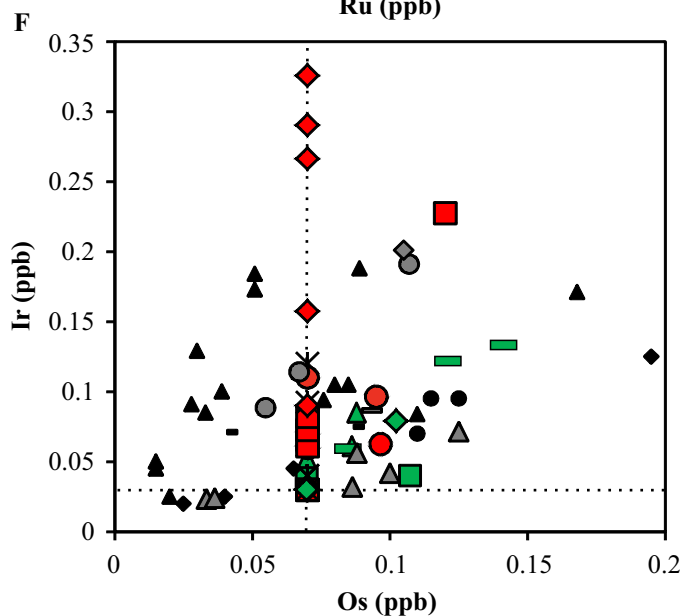
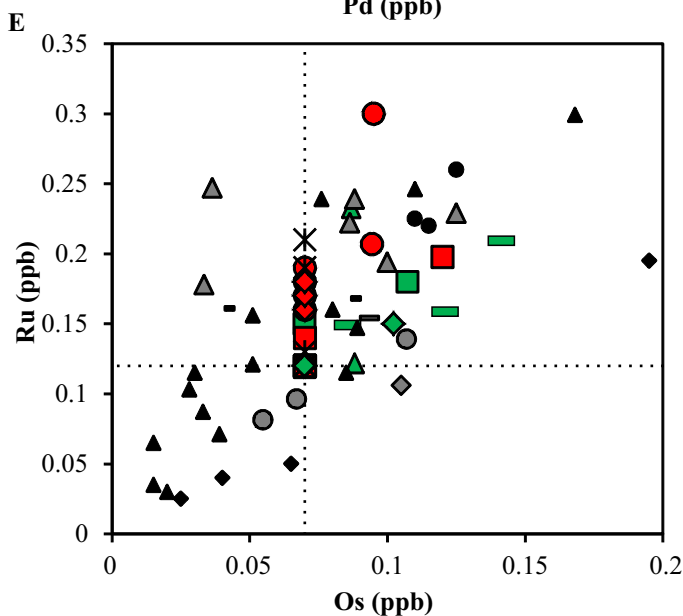
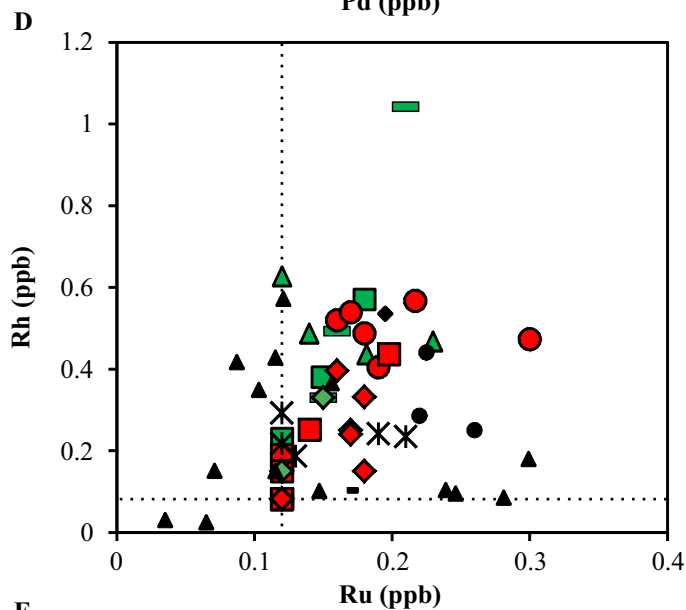
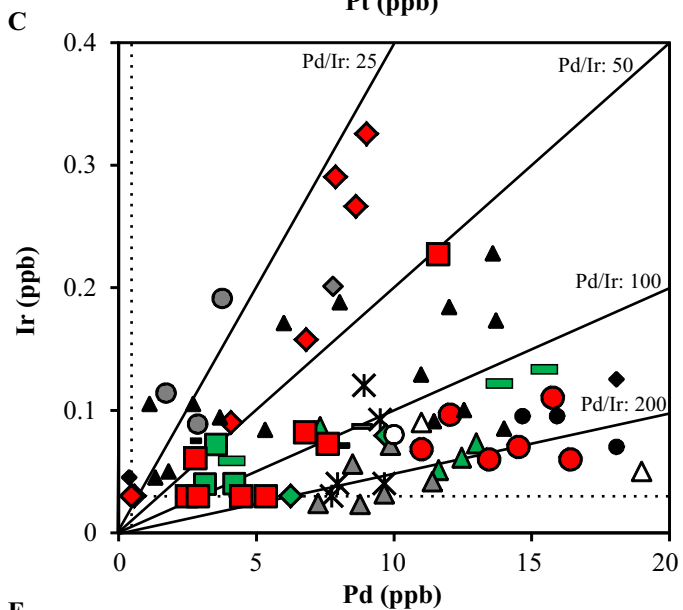
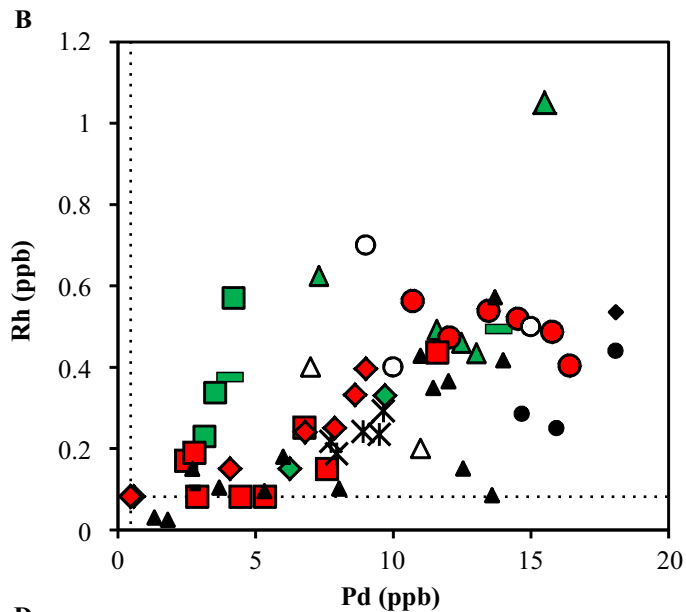
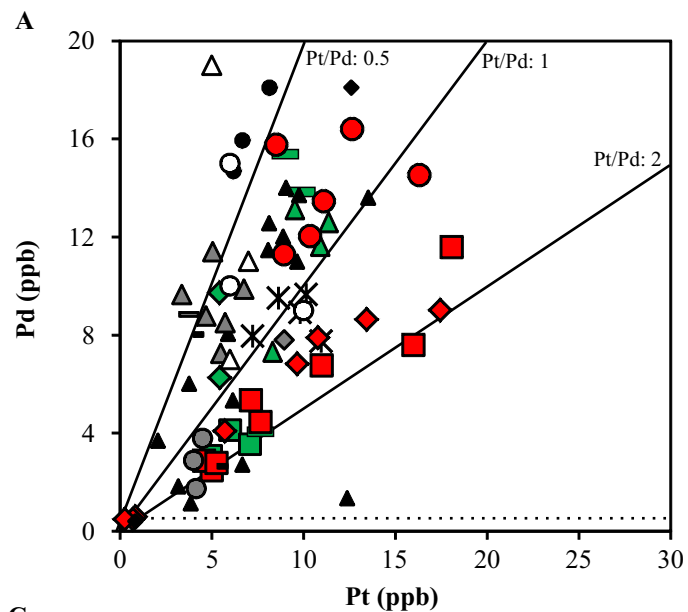


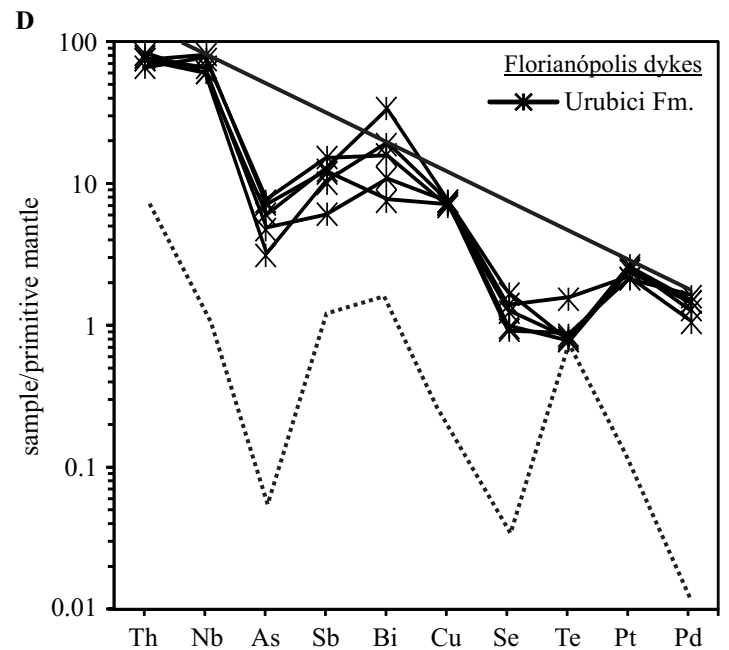
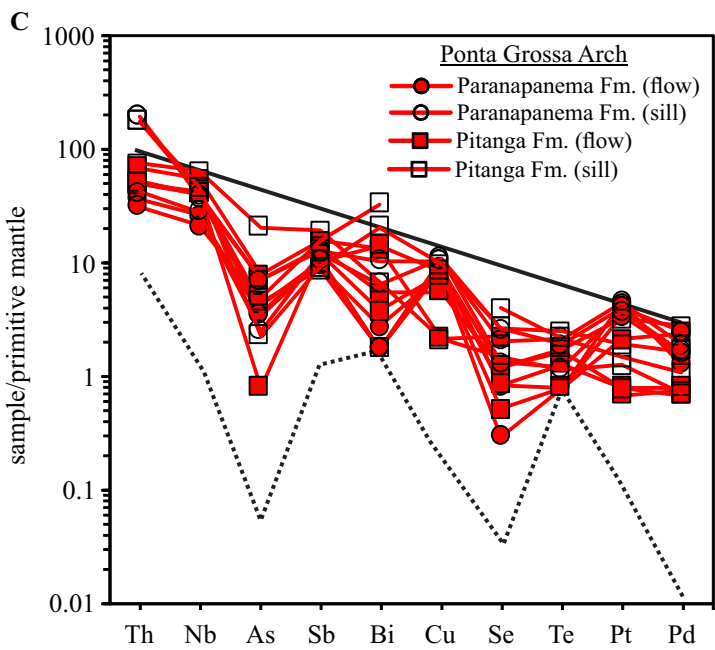
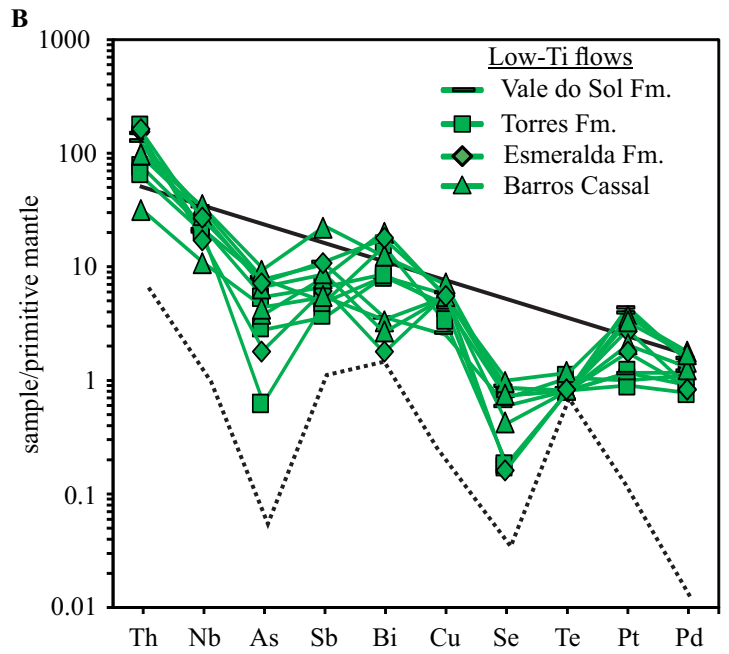
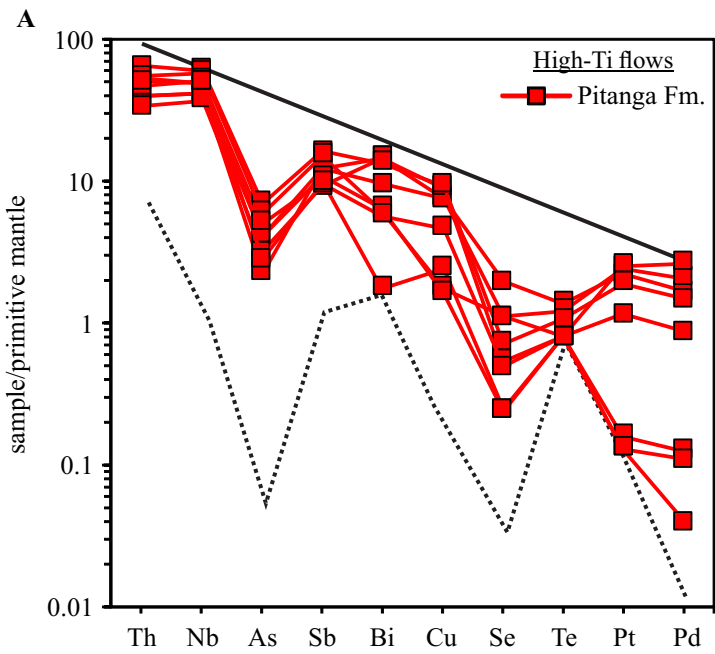
<i>This study</i>		<i>Rocha-Júnior et al. (2012)</i>	<i>Mincato (2000)</i>	<i>Barnes and Mansur</i>
<u>High-Ti flows</u>		<u>High-Ti flows</u>	<u>High-Ti flows</u>	<u>Etendeka Province</u>
◆ Pitanga type	■ Pitanga type	◆ Pitanga type	◆ Pitanga type	○ Tafelberg Fm.
	● Paranapanema type	● Paranapanema type	● Paranapanema type	△ Horingbaai Fm.
<u>Low-Ti flows</u>		<u>Low-Ti flows</u>	<u>Low-Ti flows</u>	
■ Vale do Sol Fm.	■ Florianópolis Dyke Swarm	▲ Esmeralda type	▲ Gramado type	
■ Torres Fm.	✱ Urubici type			
◆ Esmeralda Fm.		<u>Florianópolis dykes</u>	<u>Florianópolis dykes</u>	
▲ Barros Cassal sequence		■ Urubici type	■ Urubici type	

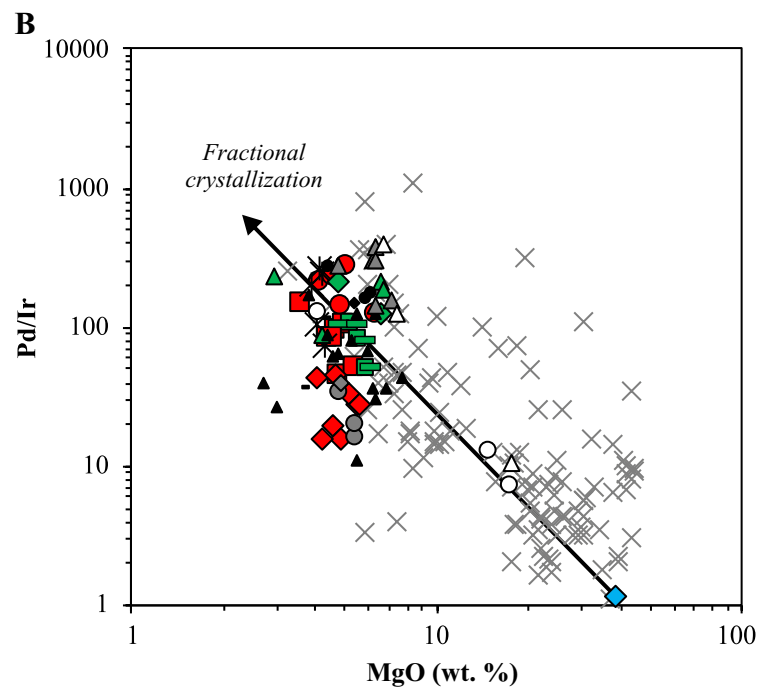
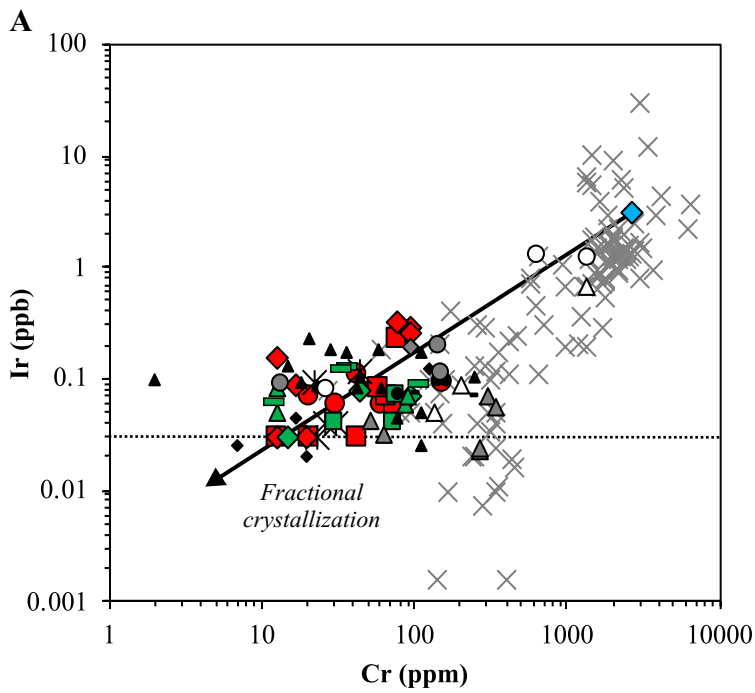




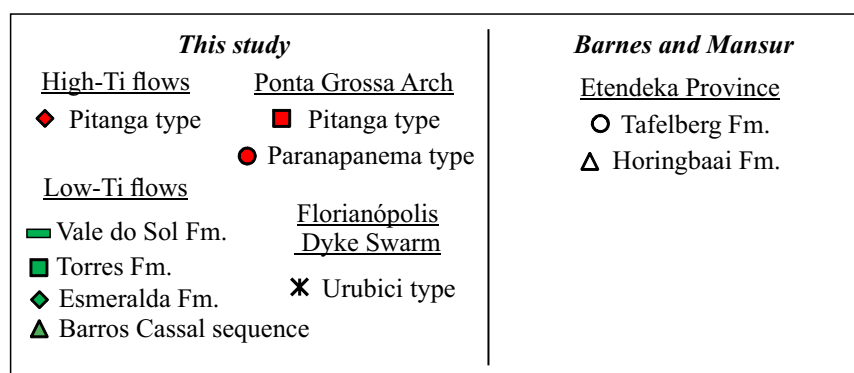
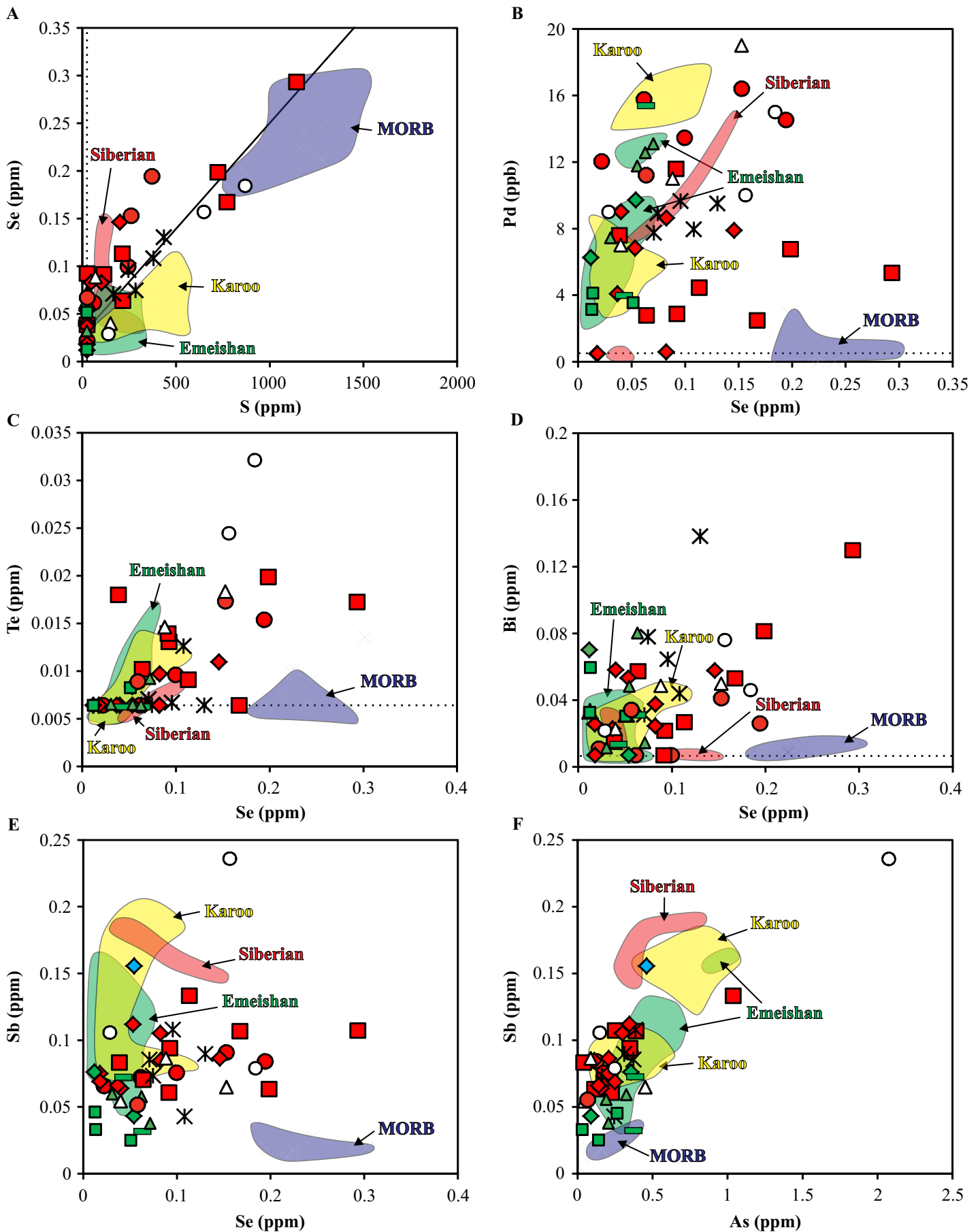


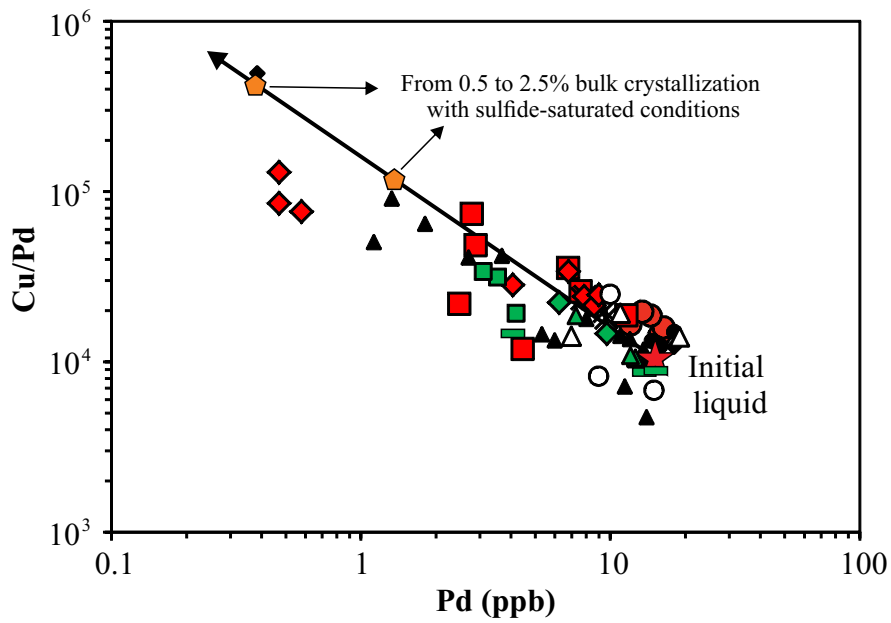




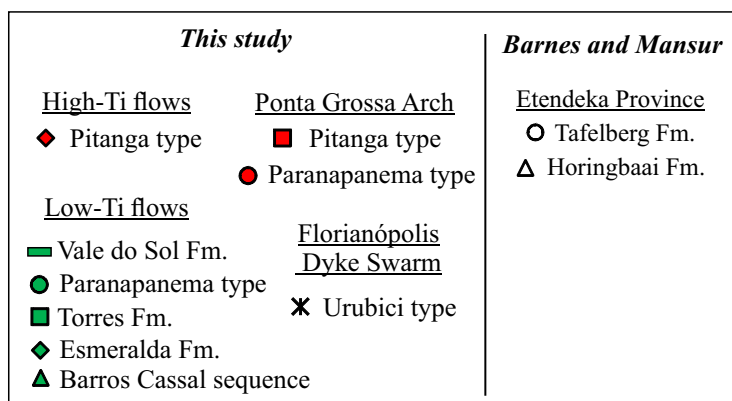
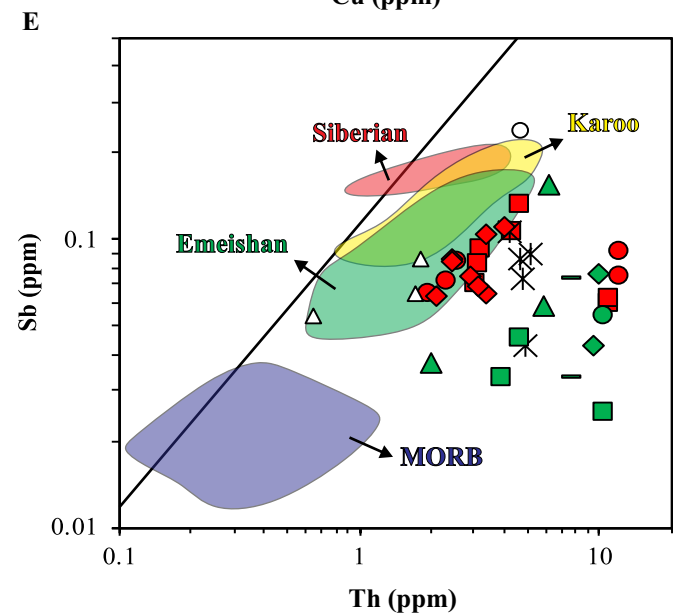
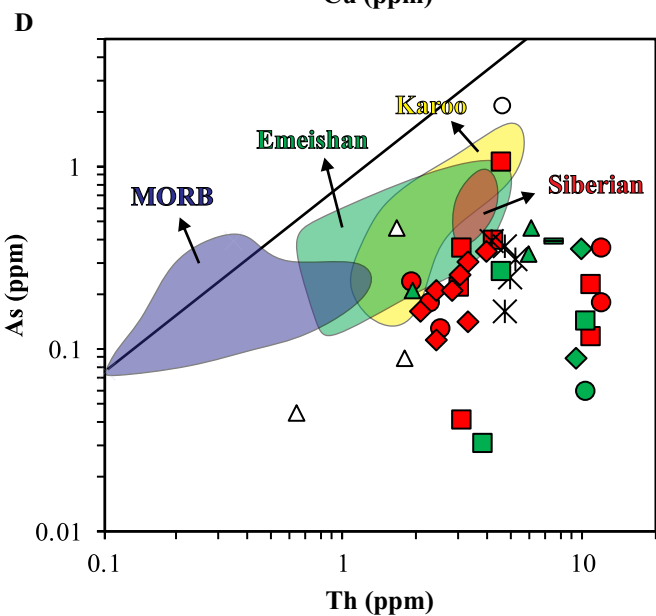
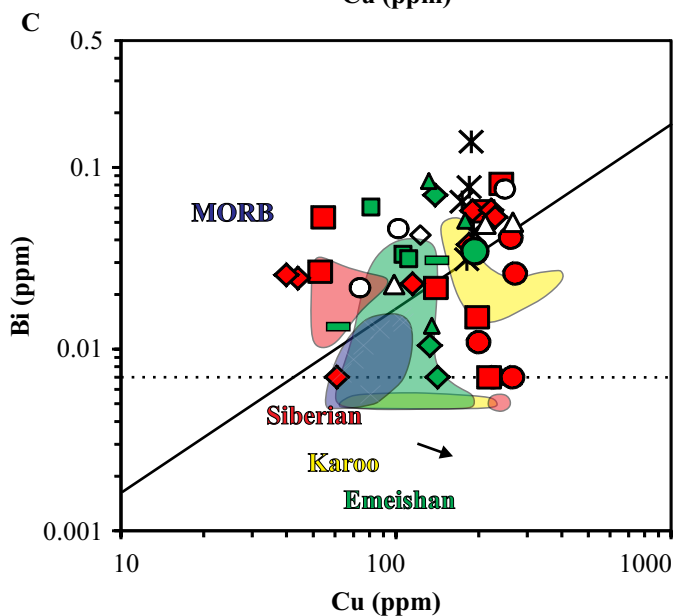
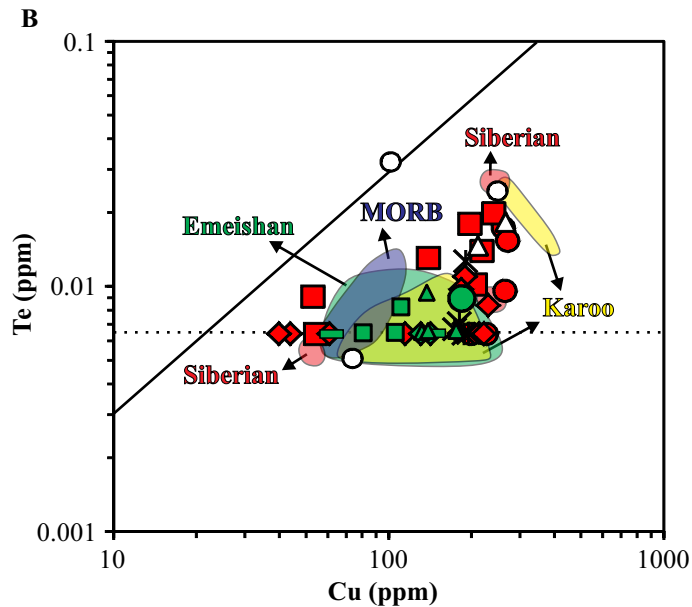
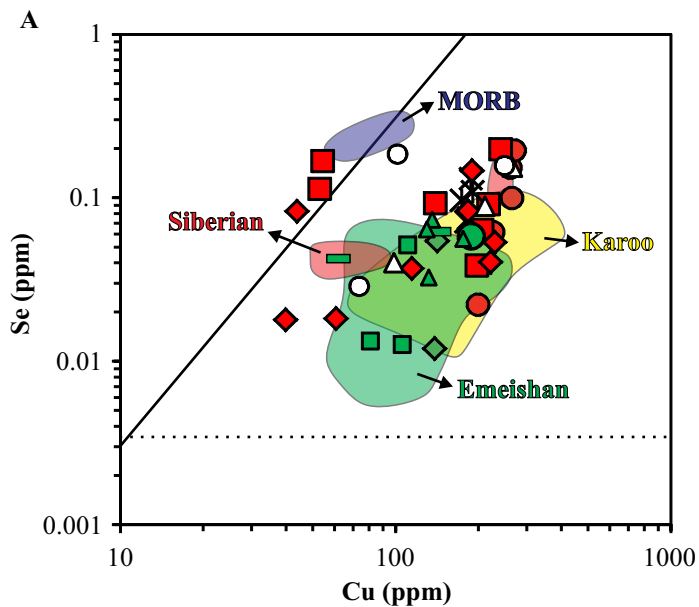


<i>This study</i>		<i>Rocha-Júnior et al. (2012)</i>	<i>Mincato (2000)</i>	<i>Barnes and Mansur in press</i>
<u>High-Ti flows</u>	<u>Ponta Grossa Arch</u>	<u>High-Ti flows</u>	<u>High-Ti flows</u>	<u>Etendeka Province</u>
◆ Pitanga type	■ Pitanga type	◆ Pitanga type	◆ Pitanga type	○ Tafelberg Fm.
	● Paranapanema type	● Paranapanema type	● Paranapanema type	△ Horingbaai Fm.
<u>Low-Ti flows</u>	<u>Florianópolis Dyke Swarm</u>	<u>Low-Ti flows</u>	<u>Low-Ti flows</u>	× Karroo, Emeishan Siberia, MORB
■ Vale do Sol Fm.		△ Esmeralda type	▲ Gramado type	
■ Torres Fm.				
◆ Esmeralda Fm.	× Urubici type	<u>Florianópolis dykes</u>	<u>Florianópolis dykes</u>	
▲ Barros Cassal sequence		■ Urubici type	■ Urubici type	





<i>This study</i>		<i>Mincato (2000)</i>	<i>Barnes and Mansur</i>
<u>High-Ti flows</u>		<u>High-Ti flows</u>	<u>Etendeka Province</u>
◆ Pitanga type	■ Pitanga type	◆ Pitanga type	○ Tafelberg Fm.
	● Paranapanema type	● Paranapanema type	△ Horingbaai Fm.
<u>Low-Ti flows</u>		<u>Low-Ti flows</u>	
■ Vale do Sol Fm.	<u>Florianópolis Dyke Swarm</u>	▲ Gramado type	
■ Torres Fm.	✱ Urubici type	<u>Florianópolis dykes</u>	
◆ Esmeralda Fm.		— Urubici type	
▲ Barros Cassal sequence			



READ ME

Online Resource (ESM) for

The distribution of platinum-group elements and Te, As, Bi, Sb and Se (TABS+) in the Paraná Magmat fractionation, sulfide segregation and magma degassing

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ic Province: Effects of crystal

ha Sarde Marteleto²

Table 1 - Whole-rock results used in this study.

Sample	Occurrence	Magma-type/Formation	Ru ppb	Rh ppb	Pd ppb
			Data Source	This work	
Ponta Grossa Dyke Swarm - High-Ti					
as021	basaltic flow	Paranapanema type	0.18	0.49	15.77
as097a	basaltic flow	Pitanga type	0.12	0.17	2.48
as150	basaltic flow	Pitanga type	0.12	0.08	2.88
as202b	basaltic flow	Pitanga type	0.12	0.19	2.78
as296	basaltic flow	Paranapanema type	0.30	0.47	12.04
as552	basaltic flow	Pitanga type	0.12	0.15	7.59
gpl115a	dyke	Paranapanema type	0.19	0.40	16.41
gpl132a	sill	Pitanga type	0.12	0.08	4.45
gpl168	dyke	Paranapanema type	0.16	0.52	14.52
gpl179	sill	Pitanga type	0.12	0.08	5.34
gpl188	dyke	Paranapanema type	0.17	0.54	13.46
gpl190	dyke	Pitanga type	0.20	0.44	11.59
gpl192	sill	Pitanga type	0.14	0.25	6.77
Florianopolis Dyke Swarm					
CSC-01F	dyke	Urubici type	0.12	0.22	7.74
CSC-01G	dyke	Urubici type	0.12	0.29	9.65
VLF-31B	dyke	Urubici type	0.21	0.23	9.50
E-01F	dyke	Urubici type	0.19	0.24	8.91
E-24A	dyke	Urubici type	0.13	0.19	7.95
Northern high-Ti Basaltic Flows					
AQ-23	basaltic flow	Pitanga type	0.17	0.25	7.88
AQ-24A	basaltic flow	Pitanga type	0.18	0.33	8.62
AQ-25	basaltic flow	Pitanga type	0.16	0.40	9.01
AQ-35	basaltic flow	Pitanga type	0.17	0.24	6.81
IGN-83AA	basaltic flow	Pitanga type	0.12	0.08	0.58
IGN-83AC	basaltic flow	Pitanga type	0.12	0.08	0.47
AQ-12	basaltic flow	Pitanga type	0.18	0.15	4.08
AQ-42	basaltic flow	Pitanga type	0.12	0.08	0.47
Southern low-Ti Basaltic Flows					
RS-46	basaltic flow	Vale do Sol Fm.	0.21	1.05	15.50
RS-47	basaltic flow	Vale do Sol Fm.	0.16	0.50	13.83
RS - 71	basaltic flow	Esmeralda Fm.	0.15	0.33	9.71
RS 101	basaltic flow	Torres Fm.	0.15	0.34	3.54
RS 114	basaltic flow	Paranapanema type (south)	0.21	0.58	11.05
RS 121	basaltic flow	uncertain	0.25	0.39	13.03
GX-360B	dyke	Esmeralda Fm.	0.12	0.15	6.25
GX479	basaltic flow	Torres Fm.	0.12	0.23	3.14
GX486	basaltic flow	Barros Cassal sequence	0.23	0.46	12.48

GX489	basaltic flow	Barros Cassal sequence	0.18	0.43	13.01
XG-25	basaltic flow	Torres Fm.	0.18	0.57	4.20
XG-26	basaltic flow	Vale do Sol Fm.	0.15	0.38	4.10
XG-28	basaltic flow	Barros Cassal sequence	0.12	0.62	7.31
GX-417	basaltic flow	Barros Cassal sequence	0.14	0.48	11.63

PGE by Ni-fire assay and Te Co-precipitation, S by IR, As, Bi, Sb, Se, Te by HG-AFS, Bold for
Sources of analyses: all trace-elements by ICPMS (values in italics) and major oxides and trace
Ponta Grossa Dyke swarm: Natasha Marteleto, Francisco Negri and Júlia Guerra
Florianópolis Dike Swarm: Natasha Marteleto and Luana Florisbal
Northern Basalt Flows: João Paulo Gusão
Southern Basalt Flows: Liza Polo

Os	Ir	Pt	Au	S	As	Sb
ppb	ppb	ppb	ppb	ppm	ppm	ppm
This work						
0.07	0.11	8.51	2.58	61	0.174	0.072
0.07	0.03	4.99	3.93	771	0.389	0.107
0.07	0.03	4.57	1.23	25	0.348	0.094
0.07	0.06	5.25	7.65	215	0.216	0.070
0.10	0.10	10.36	2.09	25	0.228	0.066
0.07	0.07	15.99	1.62	25	0.040	0.083
0.07	0.06	12.64	2.08	260	0.352	0.091
0.07	0.03	4.64	0.85	213	1.037	0.133
0.07	0.07	16.30	2.15	371	0.127	0.084
0.07	0.03	7.12	1.04	1142	0.250	0.107
0.07	0.06	11.10	5.15	244	0.176	0.076
0.12	0.23	18.07	4.31	115	0.225	0.061
0.07	0.08	10.97	1.52	723	0.113	0.063
0.07	0.03	10.96	1.84	166	0.364	0.085
0.07	0.04	10.14	2.54	247	0.386	0.108
0.07	0.09	8.63	1.66	435	0.311	0.090
0.07	0.12	9.81	3.84	284	0.161	0.073
0.07	0.04	7.23	7.75	380	0.244	0.043
0.07	0.29	10.80	7.39	201	0.207	0.086
0.07	0.27	13.45	7.79	47	0.111	0.085
0.07	0.33	17.45	13.39	25	0.158	0.064
0.07	0.16	9.66	14.15	25	0.345	0.112
0.07	0.03	0.83	1.11	103	0.298	0.105
0.07	0.03	0.72	0.48	25	0.206	0.075
0.07	0.09	5.71	2.12	25	0.140	0.066
0.07	0.03	0.26	0.87	25	0.249	0.069
0.14	0.14	9.03	4.16	25	0.377	0.033
0.12	0.12	9.90	4.14	25		
0.10	0.08	5.42	0.64	25	0.090	0.043
0.07	0.07	7.07	7.01	25	0.139	0.025
0.10	0.07	9.04	3.87	25	0.057	0.054
0.09	0.11	14.92	5.82	25	0.112	0.023
0.07	0.03	5.44	3.04	25	0.356	0.076
0.07	0.04	4.96	8.55	25	0.262	0.046
0.09	0.06	11.41	3.61	25	0.189	0.056

0.07	0.07	9.55	7.68	25	0.211	0.038
0.11	0.04	5.89	2.29	25	0.030	0.033
0.09	0.06	7.62	6.96	25	0.387	0.073
0.09	0.08	8.31	1.39	25	0.324	0.059
0.07	0.05	10.92	0.85	25	0.460	0.155

ont detection limits. Majors by XRF, trace elements by XFR (normal font) or ICP-MS italics
-elements by XFR from works supervised by Valdecir Janasi, with the participation of the following

Bi	Te	Se	SiO2	TiO2	Ti	Al2O3	Fe2O3
ppm	ppm	ppm	%	%	%	%	%
			??				
0.007	0.006	0.062	50.86	2.14	1.30	12.72	15.38
0.053	0.006	0.167	50.34	3.58	2.18	13.12	14.08
0.022	0.013	0.092	50.43	3.28	1.98	12.96	14.61
0.057	0.010	0.064	49.61	3.59	2.16	12.76	15.11
0.011	0.006	0.022	49.85	1.82	1.11	13.58	13.21
0.015	0.018	0.039	48.77	3.34	2.04	12.81	15.03
0.041	0.017	0.153	50.74	2.47	1.48	12.15	15.92
0.027	0.009	0.113	53.67	3.28	1.97	12.98	13.18
0.026	0.015	0.194	51.56	2.52	1.51	12.77	16.32
0.130	0.017	0.293	48.67	4.37		12.17	15.36
0.007	0.010	0.100	51.26	2.47	1.48	12.70	16.19
0.007	0.014	0.091	49.06	3.10	1.86	13.02	14.27
0.081	0.020	0.198	48.33	3.99	2.39	12.81	15.64
0.031	0.007	0.071	52.05	3.66	2.19	13.00	12.88
0.064	0.007	0.096	51.59	3.56	2.13	13.35	12.85
0.138	0.006	0.130	50.28	4.09	2.45	12.70	13.73
0.078	0.006	0.075	50.44	4.02	2.41	12.83	13.68
0.044	0.013	0.108	52.13	3.58	2.14	12.80	13.22
0.058	0.011	0.146	48.70	3.22	1.93	12.87	14.31
0.038	0.010	0.083	48.76	3.19	1.91	13.04	14.27
0.058	0.006	0.040	49.19	3.07	1.84	13.00	14.67
0.053	0.008	0.053	50.63	3.88	2.32	12.41	15.28
0.024	0.006	0.083	48.90	3.85	2.31	12.90	14.70
0.026	0.006	0.018	49.74	3.79	2.27	13.28	13.41
0.023	0.006	0.037	50.42	3.38	2.03	13.10	14.86
0.007	0.006	0.018	52.26	3.46	2.07	12.25	14.35
0.032	0.006	0.063	52.11	1.30	0.78	13.94	12.99
			52.15	1.30	0.78	13.98	13.02
0.007	0.006	0.054	49.80	1.39	0.84	13.73	13.67
0.031	0.008	0.051	53.59	1.07	0.64	14.08	11.64
0.035	0.009	0.061	50.62	2.01	1.20	13.19	14.68
0.007	0.011	0.154	49.41	1.26	0.76	13.74	12.73
0.070	0.006	0.012	51.12	1.43	0.86	14.44	14.68
0.033	0.006	0.013	52.18	1.16	0.69	13.64	11.91
0.080	0.006	0.063	49.71	1.21	0.72	13.73	13.14

0.013	0.009	0.072	50.02	1.22	0.73	13.86	13.01
0.061	0.006	0.013	52.75	1.00	0.60	15.87	9.23
0.014	0.006	0.043	57.05	1.57	0.94	13.07	13.36
0.010	0.006	0.031	54.10	1.39	0.83	13.50	12.62
0.048	0.006	0.055	54.17	1.88	1.12	12.15	15.65

: researchers:

MnO	MgO	CaO	Na2O	K2O	P2O5	P	Loi	Total	Ba
%	%	%	%	%	%	%	%	%	ppm
0.22	4.86	9.37	2.34	0.68	0.24	0.10	0.85	99.66	285
0.20	4.36	8.14	2.90	1.48	0.51	0.22	1.11	99.82	551
0.25	4.65	8.61	2.78	1.39	0.37	0.16	0.01	99.34	433
0.25	4.74	8.99	2.64	1.30	0.43	0.19	0.16	99.58	415
0.20	6.19	10.66	2.19	0.48	0.20	0.09	0.76	99.13	270
0.22	4.90	9.07	2.52	1.00	0.39	0.18	0.36	98.41	433
0.23	5.04	8.88	2.30	0.96	0.26	0.11	0.00	98.95	254
0.19	3.58	7.26	2.95	2.04	0.56	0.25	0.20	99.89	608
0.23	4.18	8.84	2.48	0.99	0.30	0.13	0.00	100.18	287
0.22	4.37	7.61	3.15	1.40	0.48		1.09	98.88	< 37
0.23	4.63	9.10	2.27	1.02	0.28	0.12	0.00	100.14	330
0.21	5.38	9.49	2.27	0.93	0.34	0.15	1.48	99.55	484
0.19	4.55	8.61	2.27	1.08	0.44	0.19	1.27	99.18	250
0.16	4.14	7.82	3.01	1.84	0.56	0.24	< 0.01	99.12	609
0.16	4.20	8.19	2.75	1.87	0.53	0.23	0.04	99.09	549
0.17	4.04	7.58	2.73	2.19	0.60	0.26	0.66	98.77	648
0.19	4.33	7.91	2.59	1.97	0.57	0.25	0.68	99.21	619
0.17	3.94	7.48	2.98	2.06	0.58	0.25	0.61	99.54	629
0.19	5.50	9.66	2.25	0.87	0.35	0.15	1.43	99.36	585
0.18	5.13	9.84	2.20	0.84	0.36	0.16	1.95	99.76	304
0.20	5.58	9.77	2.32	1.06	0.32	0.14	0.56	99.74	315
0.23	4.08	7.79	2.64	1.76	0.64	0.28	0.17	99.51	537
0.21	4.62	8.64	2.51	1.18	0.57	0.25	1.23	99.31	480
0.17	4.92	8.66	2.57	1.50	0.50	0.22	1.07	99.60	449
0.21	4.66	8.54	2.58	1.53	0.41	0.18	0.42	100.11	458
0.21	4.21	7.93	2.20	1.25	0.48	0.21	1.07	99.66	533
0.20	5.64	9.47	2.54	0.99	0.18	0.08	0.20	99.55	273
0.20	5.49	9.36	2.78	1.22	0.18	0.08	0.32	100.00	297
0.19	6.54	10.36	2.30	0.58	0.16	0.07	1.23	99.96	284
0.19	5.98	9.69	2.23	1.00	0.12	0.05	0.30	99.89	216
0.20	5.38	9.35	2.44	1.06	0.23	0.10	0.66	99.82	295
0.20	7.18	11.18	2.48	0.32	0.11	0.05	1.46	100.07	77
0.17	4.75	7.33	2.24	0.70	0.18	0.08	2.50	99.54	296
0.18	5.57	8.51	3.25	1.62	0.17	0.07	0.93	99.12	369
0.18	6.62	10.60	2.17	0.52	0.13	0.06	0.52	98.53	156

0.19	6.72	10.55	2.18	0.74	0.13	0.06	0.48	99.10	<i>146</i>
0.15	5.26	9.26	1.96	1.05	0.18	0.08	3.22	99.94	<i>399</i>
0.19	3.05	6.60	2.57	2.33	0.21	0.09	0.42	100.42	<i>407</i>
0.17	4.25	8.03	2.60	1.74	0.20	0.09	0.53	99.13	<i>333</i>
0.21	2.95	6.90	2.57	2.08	0.28	0.12	0.44	99.27	<i>352</i>

Ce ppm	Co ppm	Cr ppm	Cs ppm	Cu ppm	Dy ppm	Er ppm	Eu ppm	Ga ppm	Gd ppm
44	40	43	0.20	226	5.80	3.89	1.96	20.4	6.33
85	31	21	0.45	54	7.66	4.04	3.46	24.2	9.31
64	39	44	0.18	139	6.50	3.56	2.63	22.6	7.45
68	36	76	0.18	206	6.69	3.33	2.44	22.8	7.40
37	44	152	0.04	200	4.60	2.70	1.52	18.5	4.78
67	41	66	0.25	198	6.15	3.49	2.60	20.9	7.99
44	42	61		261				21.0	
97	29	13	0.62	53	8.77	4.05	3.32	24.0	9.79
51	41	21	0.22	271	6.73	3.92	2.05	20.0	6.60
72	42	85		232				19.0	
48	47	31		265				21.0	
35	48	77		218				23.0	
37	41	60		241				21.0	
86	34	23	2.02	181	6.86	3.35	3.11	26.0	9.11
78	33	26	2.97	174	6.37	3.17	3.01	25.0	8.47
112	38	22	1.75	188	8.69	4.03	3.85	26.0	11.08
102	38	44	1.77	185	7.83	3.58	3.57	25.0	10.19
101	32	31	0.95	191	8.01	3.75	3.51	25.0	10.19
62	43	96	0.60	190	5.95	2.88	2.24	21.0	6.87
68	41	96	0.52	184	5.80	2.88	2.17	21.0	7.20
74	42	78	0.09	222	5.55	2.58	1.99	21.0	6.29
92	38	< 13	0.32	230	9.31	4.25	3.06	24.0	9.68
57	32	< 13	0.22	44	7.32	3.68	3.17	25.0	9.50
65	36	20	0.15	40	7.14	3.08	2.78	24.0	8.38
66	37	17	0.17	115	6.93	3.15	2.56	23.0	7.89
61	35	< 13	0.53	61	7.48	3.50	2.82	23.0	9.17
67	43	40		142				19.9	
50	41	37		126				19.1	
53	45	44		142				20.1	
< 35	40	78		111				20.2	
74	39	94		185				20.4	
64	41	218		98				19.4	
49	45	15		139				24.0	
47	42	31	0.61	106	4.74	2.88	1.39	20.0	5.02
< 35	43	89		131				18.0	

23	42	91	3.22	136	4.01	2.60	1.17	19.0	4.07
42	33	74	0.86	81	3.96	2.31	1.27	19.0	4.17
58	36	< 13	2.35	62	6.19	3.72	1.69	22.0	6.26
48	34	< 13	2.54	133	5.63	3.40	1.54	20.0	5.54
42	42	< 13	3.72	175	6.32	3.71	1.68	22.0	6.23

Hf	Ho	La	Lu	Mo	Nb	Nd	Ni	Pb	Pr
ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm
4.61	1.37	21.3	0.52	1.44	11.8	21.8	39.2	3.01	4.89
8.87	1.58	42.5	0.48	1.66	25.1	46.8	30.1	5.04	10.30
6.15	1.36	31.0	0.47	1.16	18.8	34.4	38.6	4.11	7.94
5.88	1.28	30.8	0.45	1.26	18.6	37.6	44.1	4.17	8.86
3.39	0.96	16.8	0.39	1.01	9.8	20.2	82.5	2.87	4.64
5.94	1.27	31.0	0.43	1.36	19.1	34.1	58.7	3.92	8.08
					17.0	30.0	58.0		
9.32	1.65	44.6	0.53		29.1	58.5	31.0	5.99	12.80
5.31	1.47	22.7	0.53		13.0	27.9	41.0	3.29	6.29
					18.0	34.0	66.0		
					18.0	24.0	49.0		
					19.0	33.0	73.0		
					21.0	29.0	57.0		
7.90	1.28	39.9	0.35		37.5	45.1	42.0	7.13	12.31
7.50	1.21	35.7	0.33		35.6	41.8	47.0	10.13	11.24
9.20	1.70	51.7	0.44		28.2	60.7	56.0	7.18	14.50
8.23	1.52	47.6	0.39		28.4	54.8	63.0	7.19	13.19
8.31	1.59	47.0	0.42		30.9	54.5	52.0	6.11	12.94
5.19	1.13	26.3	0.36		18.5	32.5	83.0	3.33	7.27
5.10	1.11	26.1	0.34		18.6	32.9	74.0	3.43	7.87
4.62	1.01	22.8	0.32		16.8	32.2	73.0	3.20	7.20
7.95	1.77	40.1	0.66		27.3	54.6	48.0	5.28	12.00
7.33	1.47	37.2	0.41		26.5	46.1	34.0	4.40	10.10
6.50	1.38	32.3	0.40		23.4	43.9	37.0	3.87	9.43
6.36	1.29	34.0	0.42		22.6	43.4	47.0	4.64	9.66
6.68	1.36	32.8	0.42		22.2	44.6	32.0	4.94	10.50
					9.6	35.7	45.5		
					9.1	48.7	46.1		
					7.9	42.1	44.6		
					< 9	42.0	44.4		
					12.3	44.9	57.3		
					< 9	24.8	86.1		
					12.0	30.0	43.0		
3.82	1.00	21.3	0.41		10.2	23.5	45.0	6.88	5.99
					< 9	< 14	68.0		

<i>2.53</i>	<i>0.93</i>	<i>10.2</i>	<i>0.35</i>	<i>5.0</i>	<i>13.2</i>	<i>69.0</i>	<i>3.12</i>	<i>2.97</i>
<i>3.30</i>	<i>0.83</i>	<i>19.7</i>	<i>0.32</i>	<i>9.0</i>	<i>21.1</i>	<i>66.0</i>	<i>8.45</i>	<i>5.72</i>
<i>5.15</i>	<i>1.33</i>	<i>27.1</i>	<i>0.53</i>	<i>14.8</i>	<i>29.5</i>	<i>9.0</i>	<i>12.53</i>	<i>8.00</i>
<i>4.11</i>	<i>1.23</i>	<i>22.9</i>	<i>0.47</i>	<i>13.0</i>	<i>24.8</i>	<i>31.0</i>	<i>9.75</i>	<i>6.64</i>
<i>4.00</i>	<i>1.34</i>	<i>22.0</i>	<i>0.52</i>	<i>15.6</i>	<i>25.8</i>	<i>29.0</i>	<i>9.00</i>	<i>6.02</i>

Rb	Sc	Sm	Sr	Ta	Tb	Th	Tm	U	V
ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm
18.6	41.2	5.72	318	0.86	0.92	2.31	0.51	0.41	457
37.3	24.3	11.00	463	1.83	1.32	4.32	0.48	0.82	402
37.3	29.7	7.88	412	1.38	1.14	3.19	0.48	0.61	496
37.3	30.9	7.93	434	1.44	1.21	3.08	0.49	0.61	435
37.3	39.2	4.67	307	0.68	0.78	1.96	0.39	0.34	372
37.3	30.7	7.92	390	1.33	1.10	3.15	0.46	0.62	520
29.0	40.0		215						452
43.9	24.2	12.80	502		1.51	4.64	0.55	0.95	323
21.4	40.9	7.08	241		1.09	2.57	0.53	0.46	401
18.0	36.0		358						320
17.0	37.0		263						460
12.0	32.0		476						417
38.0	33.0		338						386
38.4	25.0	9.80	677		1.31	4.73	0.42	1.03	323
40.4	25.0	9.19	696		1.20	4.17	0.39	0.93	315
139.7	27.0	12.68	756		1.61	5.22	0.55	1.10	350
73.3	26.0	11.57	738		1.45	4.74	0.48	0.99	343
45.5	25.0	11.72	680		1.48	4.97	0.50	0.97	297
24.3	30.4	7.12	474		1.03	2.44	0.40	0.53	422
23.4	30.1	6.96	487		1.09	2.43	0.44	0.55	436
18.7	30.2	6.33	494		1.03	2.09	0.40	0.46	432
36.5	27.5	11.40	482		1.57	3.99	0.66	0.83	401
23.7	27.0	10.10	564		1.30	3.37	0.48	0.71	403
26.7	27.0	9.27	560		1.25	2.89	0.43	0.63	409
29.3	28.7	8.64	499		1.21	3.34	0.46	0.68	460
43.6	25.8	9.00	523		1.45	3.11	0.55	0.72	389
30.4	35.4		221						346
38.4	38.0		210						529
7.8	38.4		196						379
32.4	36.3		209						293
23.6	37.3		255						384
10.1	41.0		151						339
35.0	40.0		210						382
53.5	35.1	4.84	232		0.85	4.68	0.44	0.76	266
14.0	36.0		183						328

<i>42.8</i>	<i>37.6</i>	<i>3.46</i>	<i>171</i>	<i>0.69</i>	<i>1.97</i>	<i>0.37</i>	<i>0.46</i>	<i>334</i>
<i>24.6</i>	<i>31.0</i>	<i>4.40</i>	<i>367</i>	<i>0.68</i>	<i>3.91</i>	<i>0.34</i>	<i>1.03</i>	<i>236</i>
<i>87.6</i>	<i>33.0</i>	<i>6.49</i>	<i>199</i>	<i>1.07</i>	<i>7.76</i>	<i>0.57</i>	<i>1.78</i>	<i>374</i>
<i>69.6</i>	<i>31.0</i>	<i>5.57</i>	<i>213</i>	<i>0.96</i>	<i>5.92</i>	<i>0.52</i>	<i>1.55</i>	<i>329</i>
<i>96.5</i>	<i>32.0</i>	<i>5.74</i>	<i>158</i>	<i>1.03</i>	<i>6.16</i>	<i>0.56</i>	<i>1.54</i>	<i>464</i>

Y	Yb	Zn	Zr
ppm	ppm	ppm	ppm
34.7	3.12	117	177
38.6	3.05	128	349
34.5	3.08	132	240
34.2	3.07	167	246
26.0	2.42	176	137
33.0	2.68	126	246
41.0		120	178
47.1	3.58	124	381
39.9	3.38	125	206
40.0		119	193
39.0		127	184
54.0		113	194
37.0		125	184
35.2	2.51	119	283
32.8	2.38	130	271
41.4	3.07	121	356
39.1	2.64	165	322
39.8	2.88	110	327
29.9	2.51	109	210
29.7	2.61	115	213
27.9	2.57	108	193
45.8	4.43	129	326
37.7	2.86	127	306
34.6	2.71	119	271
33.8	2.94	118	267
36.9	3.33	126	279
27.8		95	120
27.0		93	119
28.2		95	123
23.9		89	100
32.3		113	157
24.3		93	74
37.0		117	154
26.8	2.77	94	146
26.0		86	100

<i>24.7</i>	<i>2.22</i>	<i>86</i>	<i>91</i>
<i>21.7</i>	<i>2.03</i>	<i>73</i>	<i>129</i>
<i>34.3</i>	<i>3.32</i>	<i>112</i>	<i>193</i>
<i>32.0</i>	<i>2.99</i>	<i>94</i>	<i>160</i>
<i>40.1</i>	<i>3.62</i>	<i>124</i>	<i>178</i>

Table 2 - Analyses of reference materials used to monitor the data quality of HG-AFS analyses.

			As ppm	Bi ppm	Sb ppm	Se ppm
Detection limits			0.003	0.007	0.009	0.003
Name	CH 4	Certificate value*	8.80	0.60	0.77	2.10
Rock Type	Anorthosite	Stdev	0.60	0.20	0.40	0.20
Source	CANMET	This study	8.29	0.50	0.74	1.98
		Stdev (n= 6)	0.45	0.06	0.07	0.15
		Relative stdev %	5.44	11.09	9.74	7.72
Name	TDB-1 Diabase	Certificate value*	2.50	<i>0.063</i>	1.00	<i>0.38</i>
Rock Type	Diabase	Stdev	0.50	<i>0.019</i>	0.40	<i>0.03</i>
Source	CANMET	This study	1.95	0.090	0.90	0.35
		Stdev (n= 6)	0.15	0.019	0.08	0.05
		Relative stdev %	7.44	21.11	8.64	13.04
Name	OKUM	Certificate value*	<i>0.24</i>	<i>0.072</i>	<i>0.111</i>	<i>0.101</i>
Rock Type	Komatiite	Stdev	0.03	0.009	0.013	0.009
Source	IAG	This study	0.27	0.093	0.093	0.090
		Stdev (n= 6)	0.03	0.019	0.011	0.010
		Relative stdev %	12.23	20.53	11.84	11.14

Detection limit = 3*sigma of the blank

*Normal font certificate or assigned value, italics average of literature values and Mansur et al. (2000)
 Stdev = standard deviation; Relative stdev= relative standard deviation; n= number of individual analyses

Te
ppm
0.006
<i>0.36</i>
<i>0.07</i>
0.33
0.04
13.11
<i>0.04</i>
<i>0.01</i>
0.03
0.01
34.62
<i>0.053</i>
0.005
0.035
0.012
34.69

020a) for
analyses

Table 3 - Analyses of reference materials used to monitor the data quality of PGE, Au and S analyses.

		Ru ppb	Rh ppb	Pd ppb	Os ppb	Ir ppb
Ref Material	Detection limits	0.12	0.08	0.47	0.07	0.03
OKUM	Certificate value	4.25	1.40	11.70		0.99
Abitibi	Stdev	0.30	0.13	0.50		0.07
komatiite	This study	4.53	1.53	12.75	<dl	1.07
IAG	Stdev (n= 3)	0.16	0.03	0.48		0.02
	Relative stdev	3.48	1.85	3.75		1.63
LK-NIP-1	Certificate value	0.44	0.90	17.96		0.19
Diabase	Stdev					
Geolabs	This study	0.54	1.00	18.70	<dl	0.17
	Stdev (n= 3)	0.02	0.09	0.77		0.02
	Relative stdev	4.21	8.57	4.14		13.78
KPT-1	Certificate value	17.30	17.10	122.00	2.80	6.80
Sudbury	Stdev	2.00	1.10	17.00	0.60	1.50
Dolerite dyke	This study	18.27	17.29	143.17	3.28	6.63
IAG	Stdev (n= 3)	0.27	0.46	20.08	0.50	0.27
	Relative stdev	1.48	2.64	14.02	15.18	4.10
LDI	Certificate value	0.31	0.69	939.26	2.80	0.07
Lac des Iles	Stdev	0.03	0.04	35.41	0.60	0.01
Gabbronorite	This study	0.40	0.82	980.12	2.34	0.15
Geolabs	Stdev (n= 3)	0.04	0.20	13.98	0.86	0.08
	Relative stdev	10.08	23.90	1.43	36.71	53.33

Detection limit = 3*sigma of the blank

Stdev = standard deviation; Relative stdev= relative standard deviation; n= number of individual analy

Values for PGE and Au in ppb. Values for S in ppm.

Pt ppb	Au ppb	S ppm
0.08	0.48	25
11.00	1.49	240
0.60	0.16	
12.12	0.77	265
0.87	0.54	
7.16	70.17	
13.43	4.63	200
14.72	5.14	225
0.35	1.49	
2.39	29.01	
97.40	41.10	10293
20.10	21.30	341
114.52	40.94	
33.92	10.12	10188
29.62	24.73	
100.46	63.00	
10.87	17.91	
119.51	80.22	
5.64	1.82	
4.72	2.27	

yses

Table 4 - Analyses of reference materials used to monitor the data quality of whole-rock analyses of maj

		SiO₂	TiO₂	Al₂O₃	Fe₂O₃	MnO	MgO	CaO
JB 1a	This study	52.60	1.284	14.46	9.05	0.148	7.81	9.37
	Certificate value	52.160	1.3	14.51	9.10	0.15	7.75	9.23
JG 1a	This study	72.03	0.234	14.23	2.00	0.060	0.77	2.13
	Certificate value	72.19	0.25	14.22	2.05	0.06	0.69	2.13

or oxides and trace elements. Values for major oxides in percentage. Values for trace elements in ppm.

Na2O	K2O	P2O5	LoI	Total	Ba	Ce	Co	Cr
2.84	1.40	0.258	1.1	100.32	495	86	37	420
2.74	1.42	0.26	1.1	99.72	497	66.1	39.5	415
3.53	4.00	0.090	0.59	99.66	430	43	< 6	23
3.41	4.01	0.08	0.59	99.68	458	45.2	5.7	18.6

Cu	Ga	La	Nb	Nd	Ni	Pb	Rb	Sc	Sr	Th
56	17	40	25	30	144	5	39	30	437	10
55.5	18	38.1	27	25.5	140	7.2	41	27.9	443	8.8
< 5	16	< 28	11	21	< 5	38	169	< 14	170	15
1.3	17	21.8	12	21	6.4	27	180	6.31	185	12.1

U	V	Y	Zn	Zr	F
6	200	23	81	135	461
1.6	220	24	82	146	385
13	22	27	37	111	695
4.7	23	31.6	38.8	121	450