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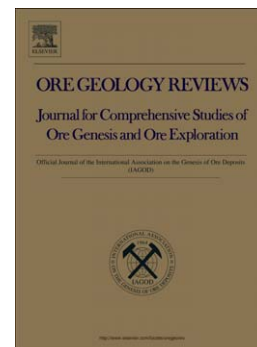
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The role of black shales as a source of sulfur and semimetals in magmatic nickel-copper deposits: Example from the Partridge River Intrusion, Duluth Complex, Minnesota, USA.

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Abstract

The basal unit of the Duluth Complex (Minnesota, USA) contains Ni-Cu sulfide deposits. The S in these is thought to be derived from a sulfide-rich black shale unit known as the Bedded Pyrrhotite Unit, a stratigraphic unit within the Virginia Formation host rocks. However, the mechanism of S transfer has not been clearly established. In order to understand how this transfer occurs we have undertaken petrography and whole rock geochemistry of the rocks from the contact aureole and the basal unit.

In the contact aureole, the Bedded Pyrrhotite Unit consists of a very fine-grained graphitic shales with thin beds of sulfides consisting of pyrrhotite with minor chalcopyrite (<1%). The basal unit contains numerous Bedded Pyrrhotite Unit xenoliths surrounded by norites. The Bedded Pyrrhotite Unit xenoliths are partially melted and the sulfide beds are disrupted. Leucosomes are present and these contain blebs of sulfides consisting of pyrrhotite, pentlandite, cubanite and chalcopyrite. In the mafic rocks surrounding the xenoliths small patches of sulfide-bearing leucosome are found. In addition to being rich in S the Bedded Pyrrhotite Unit is rich in As 38 ppm, Sb 4.1 ppm and Bi 0.6 ppm and Te 0.4 ppm and has high $\delta^{34}\text{S}$ values. The $\delta^{34}\text{S}$, As/S, Bi/S and Sb/S decrease with distance from the xenoliths. Similarly, the Ni/S, Cu/S, Se/S and (platinum-group elements)/S ratios are higher in the mafic rocks and increase with distance from the xenoliths.

Our model proposes that droplets of sulfide melt derived from the Bedded Pyrrhotite Unit xenoliths were entrained in the anatectic silicate melt of the xenoliths and transferred to the mafic magma. The sulfide droplets equilibrated with the mafic magma. Those close to the xenoliths did not have the opportunity to react with a large quantity of magma, and hence their composition is similar to the sulfides of the Bedded Pyrrhotite Unit, i.e., rich in semimetals and poor in Ni, Cu

and PGE. Farther away from the xenoliths, the sulfide droplets could have reacted with more magma, and the composition of these sulfides approach that of sulfides derived mainly from mafic magma.

Keywords. Duluth Complex; Proterozoic black shales; *in-situ* contamination; partial melting; diffusion; Ni deposits; magmatic sulfides.

Introduction

Much of the S in the world's magmatic Ni-Cu-platinum-group element (PGE) deposits is thought to be derived from external sources in the country rocks (Leshner et al., 1984; Ripley and Li, 2013). It is generally thought that S is derived from black shales (Thériault and Barnes, 1998; Leshner and Burnham, 2001). The world famous Noril'sk-Talnakh deposits may be an exception (Grinenko, 1985; Li et al., 2003). Sulfur is not the only element that could be derived from a black shale source. Black shales are enriched in As, Sb, Te and Bi by 1 to 3 orders of magnitude relative to Mid-Ocean Ridge Basalts, picrites and primitive mantle (Table 1). These elements are all chalcophile, and therefore contamination of a mafic magma with black shale could produce sulfide melts enriched in these elements compared to sulfide melts formed from mafic magmas. The semimetals may be important in concentrating PGEs within a Ni-Cu-PGE deposit because they are the anions required to form many platinum-group minerals (PGM).

A number of mechanisms for transfer of S from black shales to magma have been proposed: bulk melting of the country rock (Leshner and Burnham, 2001), transfer by gas or hydrothermal fluids (Baker et al., 2001; Ripley et al., 2007; Molnár et al., 2009; Benkó et al., 2015a, b), and transfer of sulfide droplets during partial melting of black shale xenoliths (Queffurus and Barnes, 2014). The Partridge River Intrusion of the Duluth Complex represents an ideal and well-documented intrusion for studying contamination processes because Ni-Cu sulfide deposits are found in the basal part of the intrusion close to the contact with S-rich black shales of the Virginia Formation (Mainwaring and Naldrett, 1977; Ripley, 1981; Andrews and Ripley, 1989; Thériault et al., 1997; Thériault and Barnes, 1998; Ripley et al., 2007; Severson and Hauck, 2008; Queffurus and Barnes, 2014; Robertson et al., 2015, Fig. 3), and detailed sampling is possible through numerous boreholes across the contact between the basal unit and the country rocks.

The sulfide-rich black shale unit is known as the Bedded Pyrrhotite Unit (Severson, 1994). Based on changes in $\delta^{34}\text{S}$ and S/Se values, the Bedded Pyrrhotite Unit has been identified as the source of the sulfur that contaminated the mafic magma (Zanko et al., 1994; Queffurus and Barnes, 2014). We have undertaken a petrographic, mineralogical and whole-rock geochemical study of the Bedded Pyrrhotite Unit country rocks, the Bedded Pyrrhotite Unit xenoliths and the enclosing mafic magma with the aim of investigating S and the behaviour of the semimetals during the interaction between the black shales and the mafic magma.

1. Geological context

The Duluth Complex is a Mesoproterozoic (1100 Ma) mafic complex located in Minnesota, USA. It consists of a number of mafic intrusions (Fig. 1) that are associated with the Midcontinent Rift System, and it is related to overlying the Keweenawan flood basalts (Severson and Hauck, 1997; Ojakangas et al., 2001; Miller and Severson, 2002). The country rocks range from Paleoproterozoic sedimentary rocks of the Animikie Group in the south to the Archean granite-greenstone in the north (Fig. 1).

Magmatic Ni-Cu deposits occur at the base of two of the intrusions, the Partridge River and the South Kawishiwi. Our study focused on the deposits of the Partridge River Intrusion (Fig. 1). The basal unit is composed of the following lithologies: norite, gabbro-norite, troctolite and ultramafic rocks (Hauck et al., 1997; Thériault et al., 1997; Miller and Severson, 2002; Severson and Hauck, 2008). Norites correspond to the contaminated part of the magma and is localised in the vicinity of xenoliths in the basal part of the intrusion (Thériault et al., 1997; Queffurus and Barnes, 2014).

The deposits consist of disseminated to massive sulfides. The main minerals present are pyrrhotite, cubanite, chalcopyrite and pentlandite (Thériault and Barnes, 1998; Ripley, 2014). Massive sulfides are mainly found surrounding the country rocks xenoliths. Anastomosing veins and veinlets of massive sulfides are also found hosted by the gabbro-norites, and in the country rocks close to the contact with the intrusion (Ripley and Alawi, 1986).

The country rocks of the Partridge River intrusion are Virginia Formation sedimentary rocks of the Animikie Group. The Virginia Formation is composed of carbonate, greywacke, pelite, black shale, and siltstone (Lucente and Morey, 1983). Away from the intrusion, the sedimentary rocks are essentially unmetamorphosed and the sulfide present is pyrite (Bonnichsen, 1972; Lucente and Morey, 1983, Fig. 2A; Queffurus and Barnes, 2014). However, close to the intrusion, the sedimentary rocks of the Virginia Formation have undergone contact metamorphism at temperatures hotter than 800°C, and formed diatexite migmatites close to the contact with the mafic magma (Labotka et al., 1981; Tracy and Frost, 1991; Sawyer, 2014).

One unit of particular interest is the Bedded Pyrrhotite Unit consisting of sulfide-rich black shales in the Virginia Formation and believed to have been deposited in restricted anoxic basins (Hauck et al., 1997). The Bedded Pyrrhotite Unit is mostly present close to the contact with the Duluth Complex (Severson and Hauck, 2008). This unit is approximately 200 m thick, but has a sporadic distribution. The basal unit of the Partridge River Intrusion (Unit I) contains numerous xenoliths of the Bedded Pyrrhotite Unit and the Virginia Formation (Ripley and Alawi, 1988; Thériault et al., 2000; Severson and Hauck, 2008; Queffurus and Barnes, 2014).

2. Methodology

Thirty-five samples were selected from diamond drill-hole cores that intersected the contact between the Virginia Formation and Partridge River Intrusion at the NorthMet, Mesaba, and Wetlegs deposits (Fig. 1 and Appendix A). Bedded Pyrrhotite Unit samples from the contact aureole come from boreholes AA-18 and A4-15. Xenoliths of the Bedded Pyrrhotite Unit, plus norites and gabbronorites, were collected from the NorthMet and Mesaba deposits (Fig. 1 and Appendix A). Data were combined with results from previous studies for a total database of 126 samples (Thériault et al., 1997; Thériault and Barnes, 1998; Thériault et al., 2000; Duchesne, 2004; Queffurus and Barnes, 2014).

Polished sections of each sample were obtained and petrographic observations were made, taking particular note of the sulfide mineralogy and textural changes between samples from the contact aureole, the xenoliths, and the norites.

Sulfur, Se, PGEs, and semimetals were determined at LabMaTer, Université du Québec à Chicoutimi (UQAC). Sulfur concentrations were determined by HORIBA EMIA-220V infrared and combustion S and C analyser using the method of Bédard et al., (2008). Platinum-group elements were determined by Ni-fire assay Te-co-precipitations and ICP-MS analysis and in addition, five samples were analysed by isotope dilution using the method of Savard et al., (2010). The semimetals in the black shales were determined using a new analytical protocol specially designed for black shales using an Agilent 7700X Series ICP-MS (Henrique-Pinto et al., 2016).

Major oxides and trace elements were determined at Activation Laboratories Ltd (Actlabs), Ontario, Canada by Fusion ICP-MS (WRA42B). Sulfur isotopes were determined at the Environmental Isotope Laboratory (University of Waterloo, ON, Canada) using Elemental Analyzer-Isotope Ratio Mass Spectrometer (EA-IRMS) with V-CDT as internal standard.

Results for the certified reference materials are presented in appendix B.

3. Results

3.1 Petrography

Sulfide mineralogy and the form of the sulfides found within the contact aureole differ from those in the xenoliths. In samples of the Bedded Pyrrhotite Unit from the contact aureole partial melting occurred (Duchesne, 2004; Sawyer, 2014), and the rocks close to the contact with the intrusion are now diatexite migmatites with thin pyrrhotite beds (~0.5 to 3 mm thick) disrupted by a silicate anatectic melt network (Fig. 2A) that is now pseudomorphed by quartz, plagioclase and K-feldspar. The beds consist almost exclusively of pyrrhotite with rare (<1%) chalcopyrite intergrowths (Fig. 2B), but no pentlandite or cubanite.

In contrast, in the xenoliths the sulfide assemblage contains less pyrrhotite (~70 modal%), and more chalcopyrite, cubanite, and pentlandite (Fig. 2C). Angular shaped sulfide patches with low interfacial angles with the silicate phases, together with sulfide microveinlets that filled space between silicate grains, are observed at the margins of sulfide beds (Fig. 2D). There is also a variation in the proportion of sulfides within the xenoliths. In meter-sized xenoliths, the sulfide assemblages at the cores consists largely of pyrrhotite (~90 modal%), with some chalcopyrite and cubanite (~10 modal%), and little or no pentlandite (<1 modal%). The sulfide assemblage at the margins contain a large proportion of chalcopyrite and cubanite (up to 30 modal%), and some pentlandite (~5 modal%).

In some cases pyrrhotite-rich massive sulfides occur mixed with xenolith anatectic melt in a narrow zone (~5cm) at the edge of the Bedded Pyrrhotite Unit xenoliths (Fig. 2E). The sulfide assemblage consists of approximately equal amounts of pentlandite and chalcopyrite (~ 5 to 10 modal% of each), with the balance consisting of pyrrhotite.

The silicate portions of the Bedded Pyrrhotite Unit xenoliths have undergone partial melting. The mineralogical assemblage in the silicate portion of the Bedded Pyrrhotite Unit xenoliths is orthopyroxene + cordierite + biotite + plagioclase + melt +/- K-feldspar, graphite and ilmenite (Fig. 2F). Former melt pockets are composed of large (~500 μ m) K-feldspar and quartz grains with plagioclase, orthopyroxene, and less commonly cordierite inclusions. The melt pockets occur as films and interconnected networks, i.e., microleucosomes, in the xenoliths of the Bedded Pyrrhotite Unit. The xenolith mineralogical assemblage is consistent with those of the migmatites from the contact aureole that recorded temperatures hotter than 800°C (Sawyer, 2014).

Sulfide patches, which consist of an intergrowth of the sulfide minerals, commonly occur within the patches of former anatectic melt in the Bedded Pyrrhotite Unit xenoliths (Fig. 3A, B). The rounded shape of sulfide patches in anatectic melt contrasts with the angular shape of sulfides found in the matrix of the Bedded Pyrrhotite Unit xenoliths.

In addition, the zones (~5cm) around the xenoliths contains large amounts of anatectic melt (up to 50 modal%), and in some samples pyrrhotite-rich massive sulfides. These zones consist of xenolith anatectic melt mixed with the mafic magma (Fig. 3C, E). Relics of former anatectic melt consists of quartz-rich patches (~mm) and films (less than 100 μ m). Sulfide droplets (~50-200 μ m) are present in the anatectic melt (Fig. 3D, F).

4.2 Geochemistry

4.2.1 Chalcophile elements

The S content of the Bedded Pyrrhotite Unit from the contact aureole and the xenoliths ranges from ~ 0.5 to 10 % S. The $\delta^{34}\text{S}$ values vary from 16 to 20‰ in the contact aureole, and from 9 to 18‰ in the xenoliths (Fig. 4A). The massive sulfides contain 16 to 35 % S, and the $\delta^{34}\text{S}$ values range from 8 to 16‰, similar to the xenoliths. Most norites and gabbro-norites are poorer in S,

containing ~0.01 to 8.35% S; the $\delta^{34}\text{S}$ values are more depleted, ranging from 1.6 to ~15‰ (Fig.4A).

The metals are plotted versus S in order to examine whether they are hosted by sulfide minerals. Taking all the rock types together there is a broad correlation between the metals and S (Figs. 4 and 5). The mafic rocks are the most enriched in metals for a given S content, the xenoliths have an intermediate content, and the contact aureole rocks record the lowest metal contents (Figs. 4B, C, D and 5). The massive sulfides are richer in Ni and Co than the other rock types, and they lie along the extension of the trends for the mafic rocks. In contrast the massive sulfide samples contain more Cu than most, but not all, samples from other rock types.

Despite having similar S values, the Bedded Pyrrhotite Unit from the contact aureole has lower concentrations of chalcophile elements than the Bedded Pyrrhotite Unit xenoliths (Figs. 4B, C, D and 5). For example, the Cu content of rocks from the contact aureole is ~60 to 300 ppm, whereas the xenoliths contain ~150 to 4700 ppm. These observations are consistent with the petrographic observation that more chalcopyrite and pentlandite are observed in the Bedded Pyrrhotite Unit xenoliths than in the Bedded Pyrrhotite Unit rocks from the contact aureole.

Platinum-group elements and Ag also show broad correlations with S (Fig. 5A, B, C and D). The Bedded Pyrrhotite Unit rocks from the contact aureole have the lowest concentrations of platinum-group elements for a given S content, and the mafic rocks have the highest concentrations. The massive sulfides have greater Os, Ir, Ru and Rh concentrations than the other rocks types and the massive sulfides lie along the extension of the mafic rock trends (Ir and Rh shown in Fig. 5B, C). In contrast, the massive sulfides have contents similar in Pd, Pt, Au, and Ag to the S-rich xenoliths and mafic rocks (Pd and Ag shown in Fig. 5A, D).

The Bedded Pyrrhotite Unit rocks contain between 5 and 200 ppm As, similar to black shales averages from literature (Fig.6A and Table 1). The mafic rocks contain less As than the Bedded

Pyrrhotite Unit rocks, i.e., between 0.1 and 50 ppm, whereas the massive sulfides contains between 25 and 175 ppm, similar to the Bedded Pyrrhotite Unit. A broad correlation between As and S is shown for all the rock types. In general, the Bedded Pyrrhotite Unit rocks are richer in As than the mafic rocks for the same S content, and the Bedded Pyrrhotite Unit xenoliths appear to be particularly enriched in As.

The Bedded Pyrrhotite Unit rocks contain between 0.2 and ~10 ppm Sb, similar to black shale averages from the literature (Fig. 6B and Table 1). The mafic rocks record between ~0.05 and 4 ppm Sb, and the massive sulfides contain ~0.1 to 2 ppm Sb. There is no obvious correlation between Sb and S for the sample set as whole, but the norites show a moderate correlation.

The Bedded Pyrrhotite Unit rocks contain ~0.1 to 2 ppm Bi, and ~0.07 to 0.6 ppm Te, similar to black shale averages from literature (Fig. 6C, D, and Table 1). Mafic rocks contain ~0.1 to 5 ppm Bi and ~0.3 to 2 ppm Te. Mafic rocks contain higher Te contents for a given S content than the Bedded Pyrrhotite Unit rocks. The massive sulfides lie along the extension of the trend of the Bedded Pyrrhotite Unit.

4.2.2 Normalization to 100 % sulfides

In order to compare the composition of the sulfide component in each rock type the composition of the sulfide component has been calculated from the averages for the rocks containing more than 0.3 wt% S (Table 2). Sulfides from the Bedded Pyrrhotite Unit from the contact aureole contain the lowest Ni and Cu contents at 0.14 and 0.18 wt %, respectively, whereas the Bedded Pyrrhotite Unit xenolith sulfides have a higher Ni and Cu contents 0.75 and 1.7 wt %, respectively. The amount of Ni and Cu in the norite is greater than in the xenolith sulfides, and the gabbronorite contains the highest Ni and Cu contents at 4.5 and 13 wt % respectively. These calculated compositions are in agreement with the observations that the contact aureole does not contain pentlandite and only a little chalcopyrite, the xenolith sulfides

contain some pentlandite and chalcopyrite, and the gabbronorite sulfides have the greatest concentrations of pentlandite and chalcopyrite.

Most of the other chalcophile elements (Co to Se on Fig. 7, Table 2) follow the same order of enrichment with the sulfides from the contact aureole having the lowest concentrations and the gabbronorite having the highest concentrations. In contrast, the elements from Bi to Mo show variable degrees of enrichment. Arsenic and Sb concentrations are the highest in the Bedded Pyrrhotite Unit xenoliths, whereas they are similar for the other rock types. Lead and Bi show similar concentrations in the xenoliths and the norites.

The overall shape of the mantle normalized patterns from all four rock types is similar, with a steady increase from Co through the PGE to Ag (Fig. 7A) with Co/Ag of approximately 300 to 500. From Ag onwards the patterns tend to be approximately flat. Exceptions to these general trends are that the Bedded Pyrrhotite Unit from the contact aureole and the gabbronorite patterns have positive and negative Te anomalies, respectively.

The massive sulfides contain similar amounts of Co, Ni, Ir, Ru, and Rh to the norite sulfides, but they are depleted in most of the other chalcophile elements (Fig. 7B and Table 2). The mantle normalized concentrations increase from Co to Ag, but the pattern is not as steep as for the other sulfides, with a Co/Ag ratio of ~60 versus greater than 250 for the sulfides hosted by other rock types. The massive sulfides mantle normalized pattern show large negative Pt anomaly (Fig. 7B).

4.2.3 Change in sulfide composition with distance from the Bedded Pyrrhotite Unit xenoliths

Queffurus and Barnes (2014) found that the sulfides close to the Bedded Pyrrhotite Unit xenoliths are poorer in Se than those close to the xenoliths. Thériault and Barnes (1997) found that there are more Sb- and As-bearing minerals present in the rocks close to the xenoliths. Our current more detailed sampling shows that As and Sb contents of norite and gabbronorite-hosted sulfides decrease with distance from the Bedded Pyrrhotite Unit xenoliths in the basal Unit I (Fig.

8A, B) whereas for most of the other elements the content increases (Pd and Ir shown in Fig.8C, D).

4. Discussion

5.1 Role of partial melting in contamination processes

Based on decreases in S/Se ratios and in $\delta^{34}\text{S}$ values, and the increase in (Pt+Pd)/S ratios from the Bedded Pyrrhotite Unit xenoliths into the norite and gabbronorite of the Unit I, Queffurus and Barnes (2014) argued that S was added to the mafic magma by partial melts of the Bedded Pyrrhotite Unit carrying sulfide droplets. Our petrographic observations support this model.

Metamorphosed Bedded Pyrrhotite Unit in the contact aureole, close to the contact with the intrusion, records temperatures of 800 to 870°C (Sawyer, 2014). Xenoliths of the Bedded Pyrrhotite Unit would have experienced even hotter temperatures as they were surrounded by mafic magma (~1100-1200°C; typical values for basaltic magma). Petrological features in the Bedded Pyrrhotite Unit xenoliths show pockets and pores of former anatectic melt pseudomorphed by large grains of K-feldspar and quartz that contain inclusions of orthopyroxene, plagioclase and cordierite. These features suggest that xenoliths of the Bedded Pyrrhotite Unit reached temperature of at least 800 to ~950°C in the granulite facies (White et al., 2003; Grant, 2009; Chu and Ague, 2013).

We interpret the rounded sulfide blebs found in the patches of former anatectic melt in the Bedded Pyrrhotite Unit xenoliths, and in the anatectic patches in the norite, as globules of a sulfide melt. We suggest that small droplets of sulfide melt were carried into the mafic magma by the silicate anatectic melt that was expelled from xenoliths of the Bedded Pyrrhotite Unit and in

the process transferred S and semimetals to the mafic magma. Partial melting of sulfide minerals occurred in xenoliths of the Bedded Pyrrhotite Unit. The disrupted sulfide beds in the Bedded Pyrrhotite Unit xenoliths (Fig. 2D of this study; Fig. 2 of Queffurus and Barnes, 2014) reveal microstructures typical of sulfide partial melting, i.e., low interfacial angles between sulfide and silicate phases, and sulfide microveinlets that fill space between silicate grains (Frost et al., 2002; Tomkins et al., 2007). It might be thought that the melting temperature of pyrrhotite is too high (1190°C) for it to have melted. However, the presence of C, sulfosalts, and tellurides lowers the melting temperature of sulfides (Dasgupta et al., 2009, Tomkins et al., 2007), and thus the sedimentary sulfides could have melted.

5.2 Composition of sulfide droplets and massive sulfides

The composition of the sulfide droplets is expected to change from essentially sedimentary to igneous as they equilibrated with the mafic magma. Queffurus and Barnes (2014) modeled the changes in $\delta^{34}\text{S}$, S/Se, and (Pt+Pd)/S using equations 5 and 8 of Lesher and Burnham (2001):

$$C_s = [(C_i R + C_{ss}) D^{\text{sul/sil}}] / (R + D^{\text{sul/sil}})$$

$$I_s = \{I_i C_i [R/(1+R)] + I_{ss} C_{ss} [1/(1+R)]\} / \{C_i [R/(1+R)] + C_{ss} [1/(1+R)]\}$$

where C_s = concentration in the sulfide after equilibration; C_i = concentration of the element in the mafic magma; C_{ss} = concentration of the element in the sedimentary sulfides; R = ratio of silicate liquid to sulfide liquid; $D^{\text{sul/sil}}$ = partition coefficient between sulfide and silicate liquid; I_s = isotopic ratio in the sulfides after equilibration; I_i = isotopic ratio of mafic magma before contamination; and I_{ss} = isotopic ratio of the sedimentary sulfides.

A critical variable in these equations is the ratio of sulfide to silicate liquid. They found for the norites close to the xenoliths R -factors as low as 25 were required to model the sulfide composition, whereas for sulfides from the gabbro-norite R -factors were up to 6000. We have applied this approach to our larger data sets (Table 3).

On the plot $\delta^{34}\text{S}$ vs. Cu in 100% sulfides, a progressive decrease of $\delta^{34}\text{S}$ occurs from the Bedded Pyrrhotite Unit xenoliths to the mafic magma with intermediate values for norites (Fig. 9A). Massive sulfides and norite-hosted sulfides close to xenoliths of the Bedded Pyrrhotite Unit plot at low R-factors (50), whereas most of gabbronorite-hosted sulfides require an R-factor greater than 500 (Fig. 9A).

Nickel, Cu, Co, and trace metals concentrations are highest in gabbronorite-hosted sulfides and lowest in sulfides from the Bedded Pyrrhotite Unit xenoliths; intermediate values occur in the norite-hosted sulfides (only Pd vs. Cu is shown in Fig. 9B). Thériault et al., (1997) proposed that the metal contents of norite and gabbronorite-hosted sulfides result from variations in the degree of magma contamination and R-factor. Modeling of sulfide composition after equilibration of Bedded Pyrrhotite Unit sulfides with the mafic magma is illustrated in figure 9B for Pd. The plot of Pd vs. Cu shows that massive sulfides and norite-hosted sulfides close to xenoliths of the Bedded Pyrrhotite Unit plot at R-factor values of 50 and 500 whereas most of the gabbronorite-hosted sulfides require higher R-factors (>500). Most of the other metals show similar results.

The results for the semimetals are more complex. Despite the fact that the Bedded Pyrrhotite Unit xenoliths are rich in Bi and Te, the sulfide component of the norite is richer in Bi and Te than the xenolith sulfides (Fig. 9C, D). Modeling shows that the concentrations of these elements rise in the sulfides as the R-factor increases, i.e., as the sulfide interacts with more mafic magma. This counter-intuitive result occurs because of the high partition coefficients for these elements into the sulfide liquid (Li and Audétat, 2015; Brenan, 2015). Although the mafic magma has low Bi and Te contents compared with the Bedded Pyrrhotite Unit xenoliths, the high partition coefficients of these elements into sulfides, combined with the high R-factor, results in the sulfide droplets from the Bedded Pyrrhotite Unit xenoliths collecting a large quantity of these elements (Fig. 9C, D).

In contrast, the Bedded Pyrrhotite Unit xenolith-hosted sulfides are richer in As and Sb than the norite or gabbronorite-hosted sulfides (Fig. 9E, F). Modeling of the sulfide compositions indicates that the concentrations of these elements decrease as the sedimentary sulfide interacts with the magma, but the observed sulfide compositions cover a wide range rather than a single trend (Fig. 9E, F). Possibly this wide range of results is due to the extreme sensitivity of the partition coefficients to fO_2 and temperature (Li and Audétat, 2015). At relatively cooler temperatures (900°C) in the xenolith, the partition coefficients would have been high (100-300), whereas at hotter temperatures (1100°C) in the mafic magma, the partition coefficients would have been lower (1-10). Consequently, the effect of higher R-factor in the mafic magma was offset by the lower partition coefficient in the mafic magma.

Thériault and Barnes (1998) show based on petrographic and geochemical observations that massive sulfides surrounding the xenoliths have undergone fractional crystallization. The massive sulfides in our current study are depleted in most of the chalcophile elements relative to Co, Ni, Ir, Ru, and Rh. This type of depletion is commonly observed in monosulfide-solid solution cumulates (Barnes et al., 1997). Given the position of the massive sulfides as narrow rims around the xenoliths; our interpretation is that they represent residual monosulfide-solid solution.

The sulfides hosted by the Bedded Pyrrhotite Unit from the contact aureole show similar chalcophile mantle-normalized patterns to the sulfides in the Bedded Pyrrhotite Unit xenoliths, but are depleted in all the chalcophile elements, except Se, relative to the sulfides in the xenoliths. It is possible that the contact aureole black shales initially had a slightly different composition to the xenoliths, but field observations suggest that they are the same unit. Another possibility is that the xenoliths have preferentially lost S and thus the recalculated sulfide compositions are too high. We do not think that this occurred because the xenolith and the contact aureole rocks contain on average similar amounts of S (~4%), and the xenolith sulfides are 3 to 10 times richer

in most elements than the contact aureole sulfides, thus the xenoliths would have to have lost 66 % or more of their S. In other words, for elements such as Pd and Ni, which are strongly enriched, the xenoliths would have originally contained 24 to 40 % S. Therefore, we argue that the sulfides in the xenoliths have been enriched in chalcophile metals by diffusion of the elements from the mafic magma to the xenolith sulfides prior to the transfer of the sulfide droplets to the mafic magma. The reason for this diffusion was the chemical potential difference between the sulfide droplet in the anatectic melt in the xenolith and the chalcophile elements in the mafic magma.

5.3 Proposed model for S and semimetals contamination of the mafic magma

Based on petrological and geochemical observations, a synthesis model is proposed in Figure 10 to explain S and semimetals contamination of the mafic magma from xenoliths of the Bedded Pyrrhotite Unit:

- 1) Xenoliths of the Bedded Pyrrhotite Unit isolated in the gabbro-norite magma of the basal unit, undergo partial melting at temperatures of $\sim 900^{\circ}\text{C}$ (Fig. 10A). Sulfide layers in the xenoliths undergo partial melting and sulfide droplets are incorporated in pockets of anatectic silicate melt. Simultaneously, diffusion of Ni, Cu, and trace metals may occur from the surrounding mafic magma into the xenoliths of the Bedded Pyrrhotite Unit.
- 2) Rounded droplets of sulfides in the anatectic silicate melt are transferred to the mafic magma (Fig. 10A). As a result, norites close to the xenoliths are enriched in semimetals. After dissolution of the silicate anatectic melt in the mafic magma, the entrained sulfide droplets interact and equilibrate with the mafic magma. This results in a minor metal enrichment of the sulfide droplets because of the low R-factor calculated for norite close to xenoliths of the Bedded Pyrrhotite Unit. Dissolution of sulfide droplets in the mafic magma leads to S contamination of the mafic magma as shown by Queffurus and Barnes

(2014), with progressive decrease of $\delta^{34}\text{S}$ values from the Bedded Pyrrhotite Unit xenoliths to the mafic magma.

- 3) As the degree of partial melting increases, connectivity between pores and pockets of melt is established and large volumes of silicate anatectic melt are segregated to the mafic magma leaving the xenoliths with residual bulk compositions (Fig. 10B). As a consequence, numerous sulfide droplets trapped in the silicate anatectic melt are also transferred to the mafic magma and norites close to the xenoliths record enrichment in semimetals. Accumulation of these sulfide droplets close to xenoliths could result in the formation of massive sulfide. These massive sulfides are poor in metals because there is no interaction between the sulfides and the mafic magma. Massive sulfides may undergo fractional crystallization and as a result they represent a monosulfide-solid solution cumulate, as was suggested by Thériault and Barnes (1998).
- 4) Movement of the magma, perhaps driven by a new injection of magma or by an earthquake in the partially molten system, allows bulk flow and hence transport of the sulfides away from the xenoliths, and this results in a larger-scale contamination of the mafic magma by S and semimetals (Fig. 10C). Hence, platinum-group minerals are found in sulfides from gabbro-norites within the basal Unit I (McSwiggen, 1999; Severson and Hauck, 2003; Table 6a; Cervin, 2011). Enriched metal contents are recorded in sulfides away from the xenoliths, because of the interaction of the entrained sulfides with large volumes of magma, i.e., high R-factor.

5. Conclusions

Sulfur and semimetals contamination of the mafic magma occurs by the transfer of sulfide droplets in a mobile, silicate partial melt of the Bedded Pyrrhotite Unit xenoliths to the mafic magma. This conclusion is supported by:

- 1) Petrological observations of sulfide droplets trapped within former anatectic melt of the xenoliths of the Bedded Pyrrhotite Unit in the xenolith margins and surrounding the Bedded Pyrrhotite Unit xenoliths. We suggest that small droplets of sulfide melt were carried into the mafic magma by anatectic melt that was segregated from the Bedded Pyrrhotite Unit xenoliths.
- 2) Geochemical study which shows that a progressive decrease of $\delta^{34}\text{S}$ values from the Bedded Pyrrhotite Unit xenoliths to the mafic magma together with a progressive decrease of the semimetals content of norite- and gabbro-norite-hosted sulfides occurs with distance from the Bedded Pyrrhotite Unit xenoliths in the basal Unit I.
- 3) Modeling of the composition of sulfides shows that the semimetals content of sulfides in the magma results from their equilibration after transfer to the mafic magma and depends on R-factor, i.e., interaction rate of sulfides with the mafic magma, and partition coefficients for these elements into the sulfide liquid.

In addition, petrographic and geochemical observations show that higher proportions of pentlandite and chalcopyrite occurs in the Bedded Pyrrhotite Unit xenoliths than in the Bedded Pyrrhotite Unit from the contact aureole, and that sulfides in the Bedded Pyrrhotite Unit from the contact aureole are depleted in Ni, Cu, and trace metals relative to the Bedded Pyrrhotite Unit xenolith-hosted sulfides. We propose that these elements diffused from the mafic magma to the Bedded Pyrrhotite Unit xenolith-hosted sulfides at temperatures in excess of 900°C.

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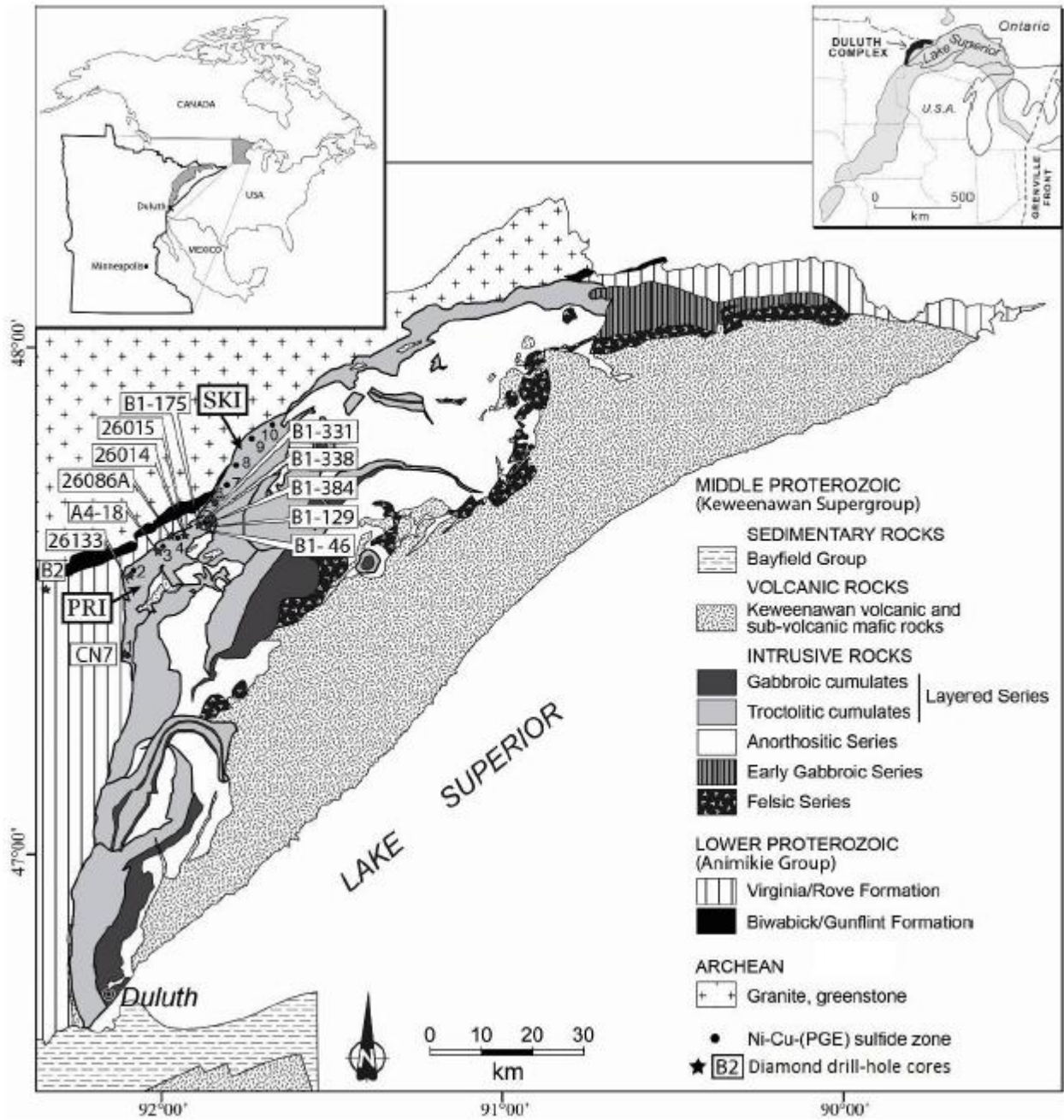


Figure 1

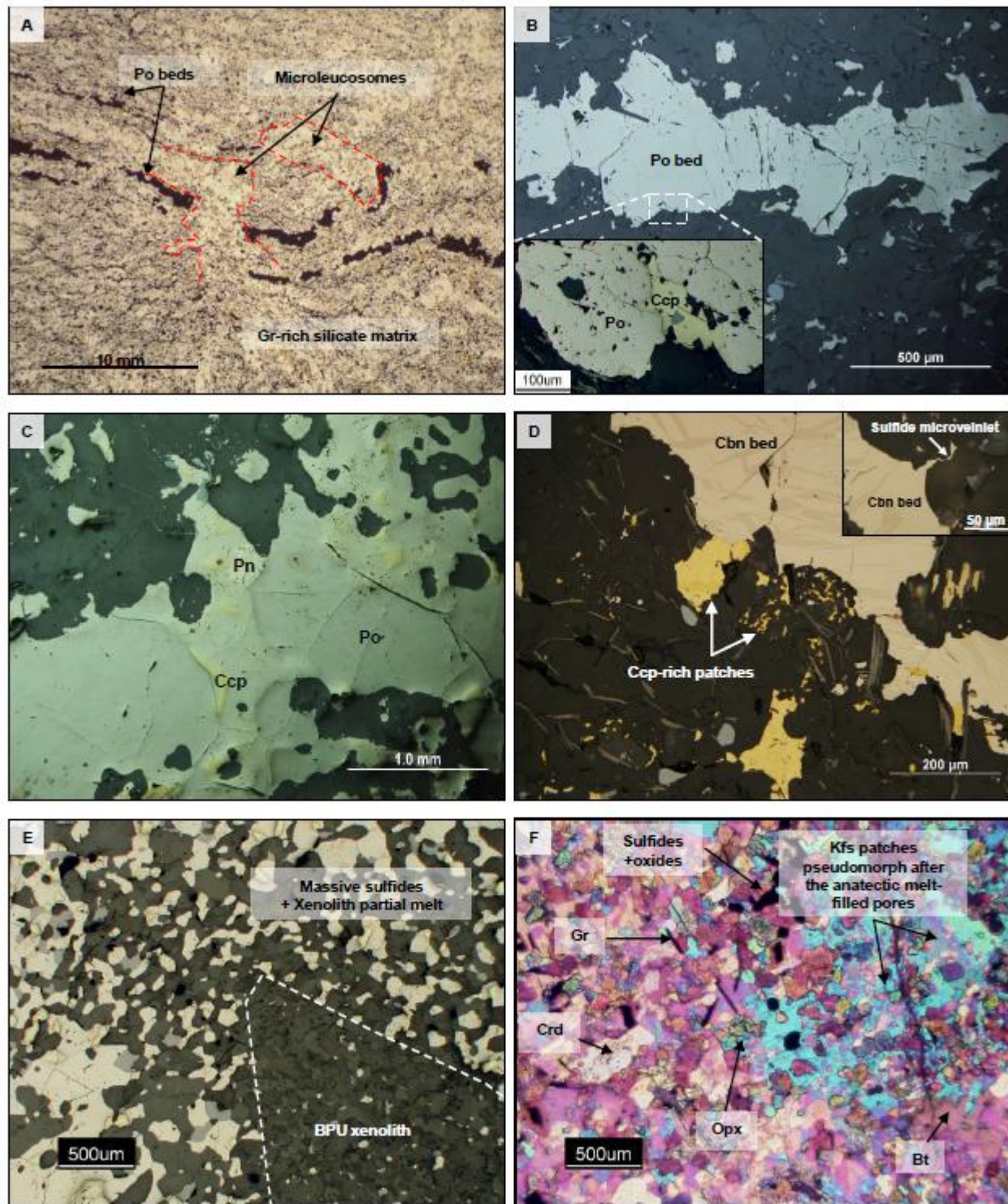


Figure 2

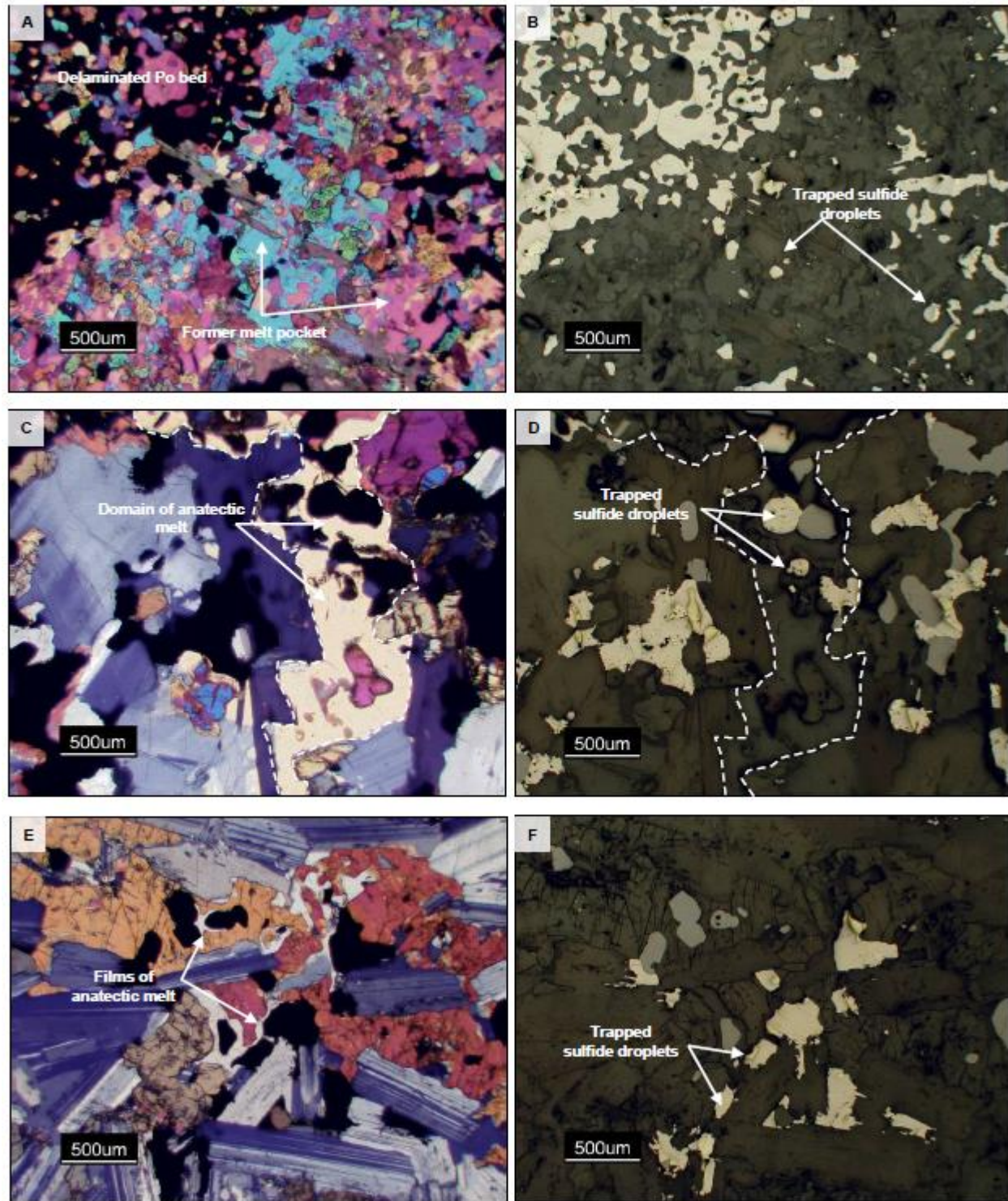


Figure 3

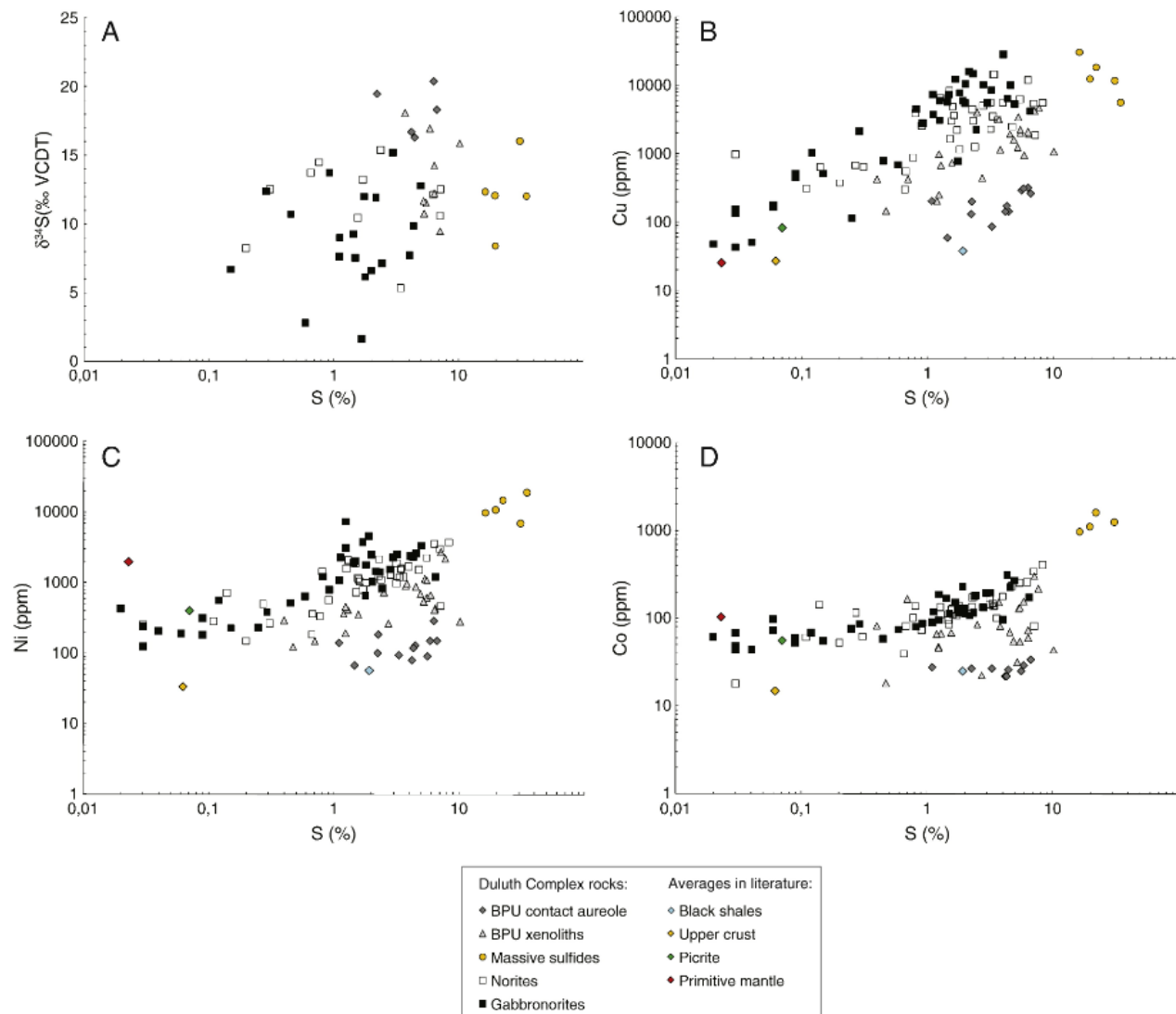


Figure 4

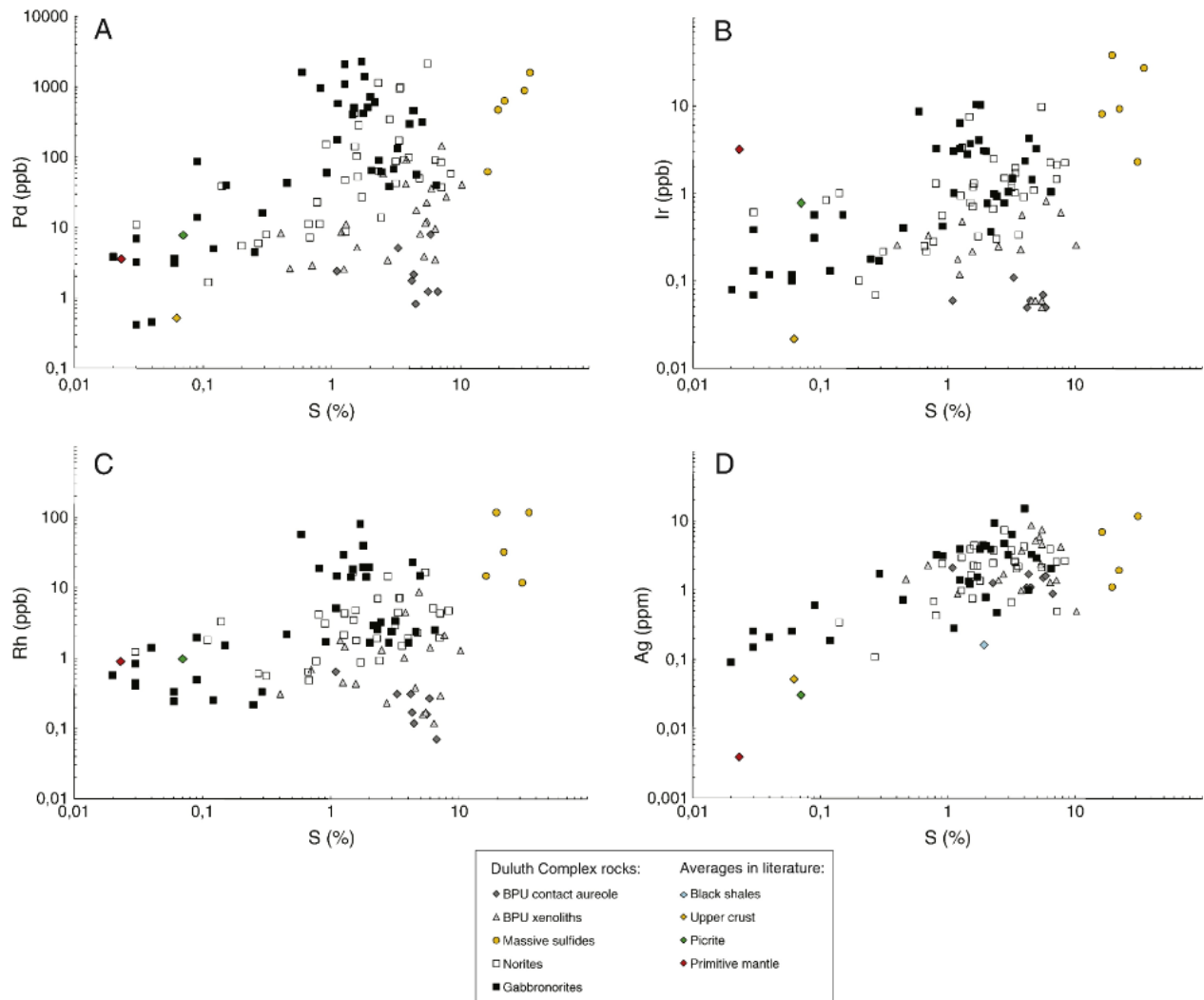


Figure 5

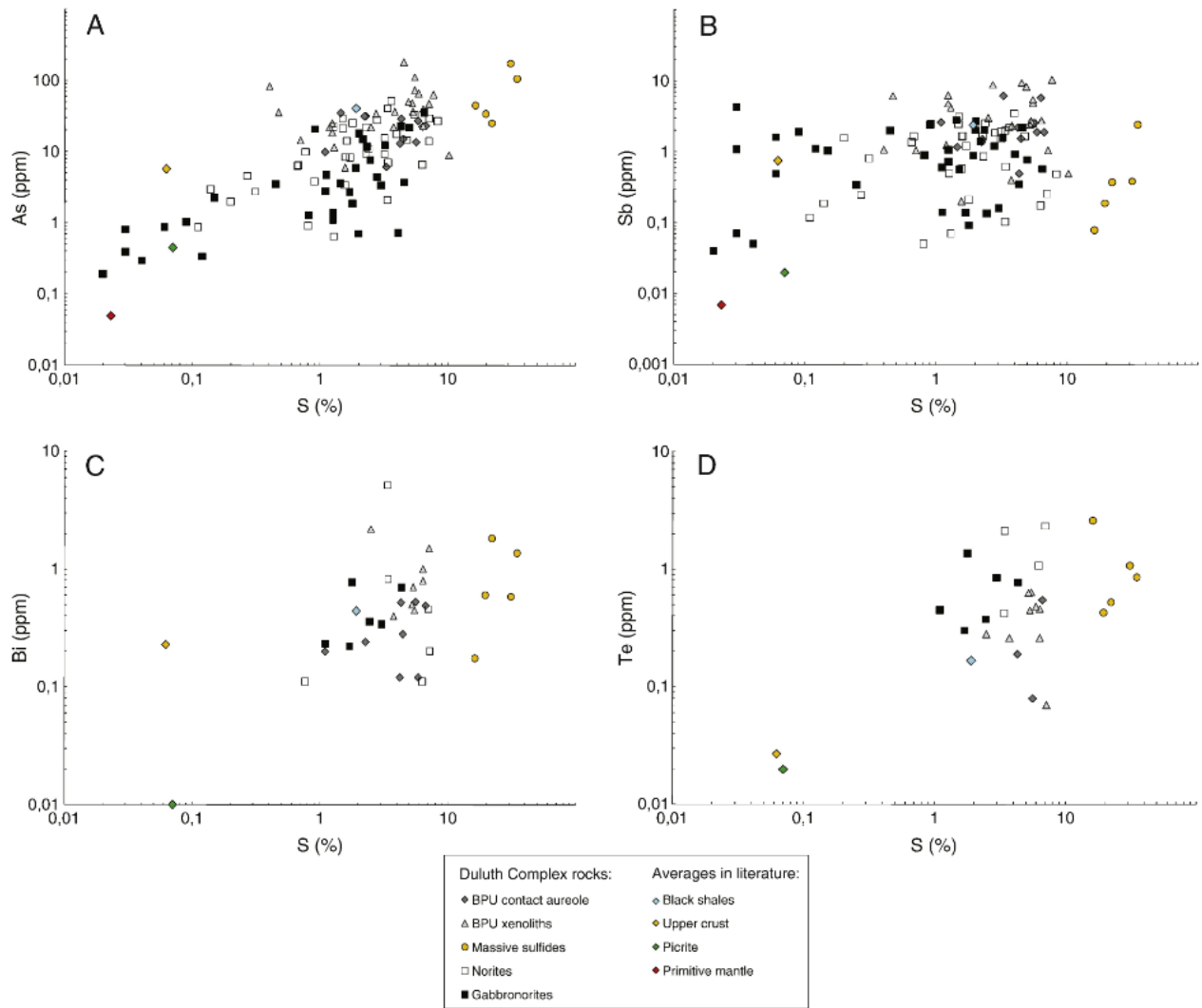


Figure 6

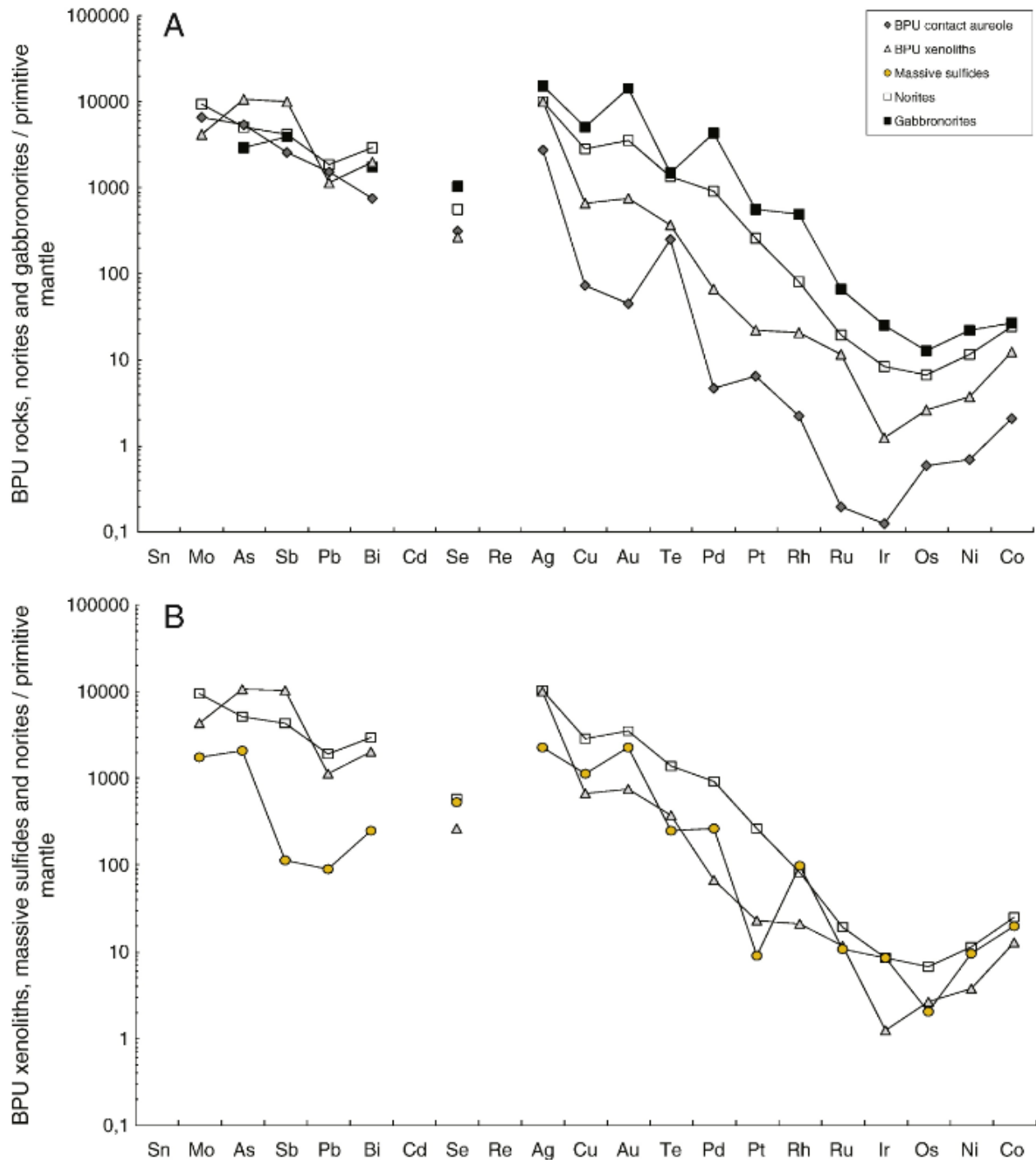


Figure 7

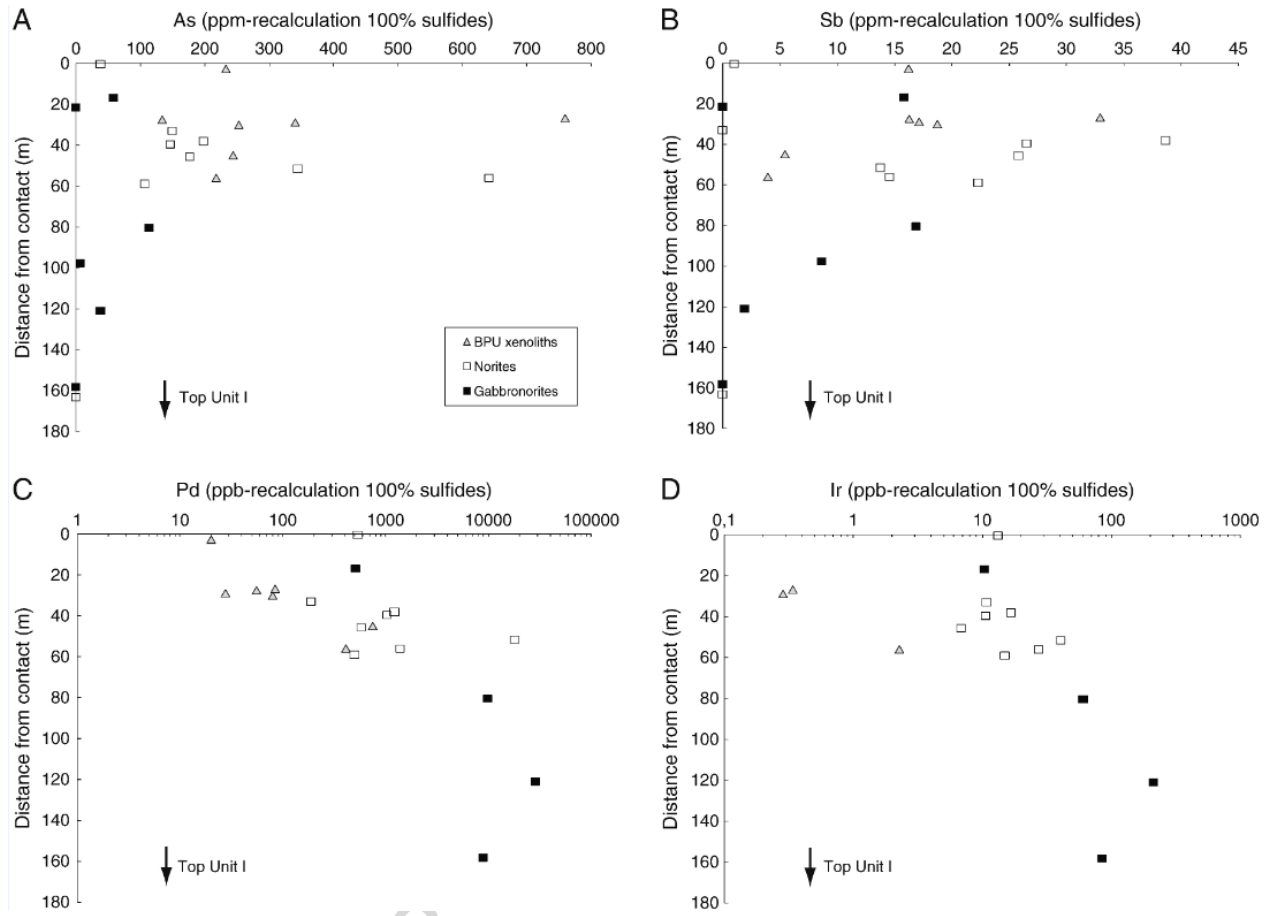


Figure 8

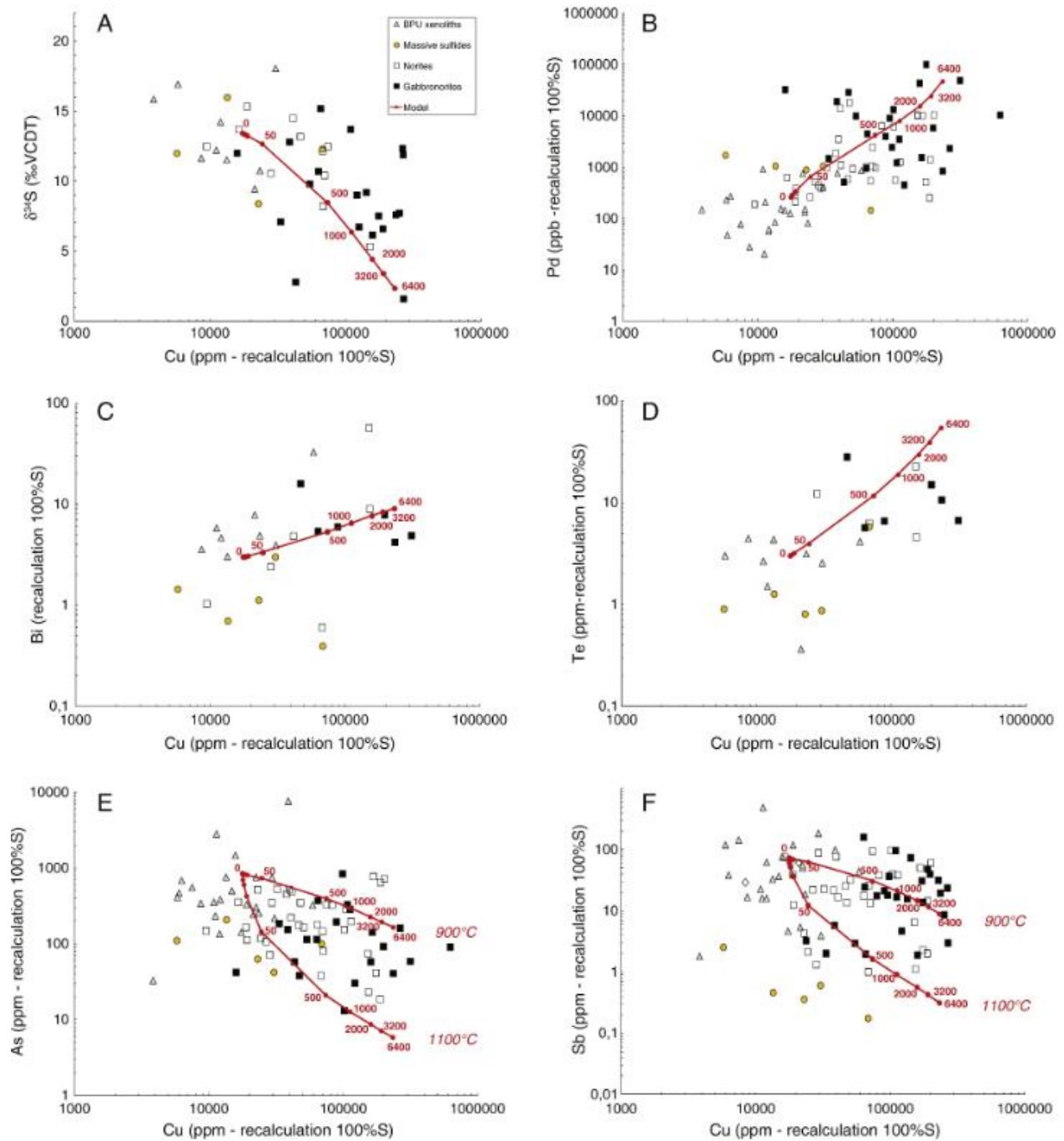


Figure 9

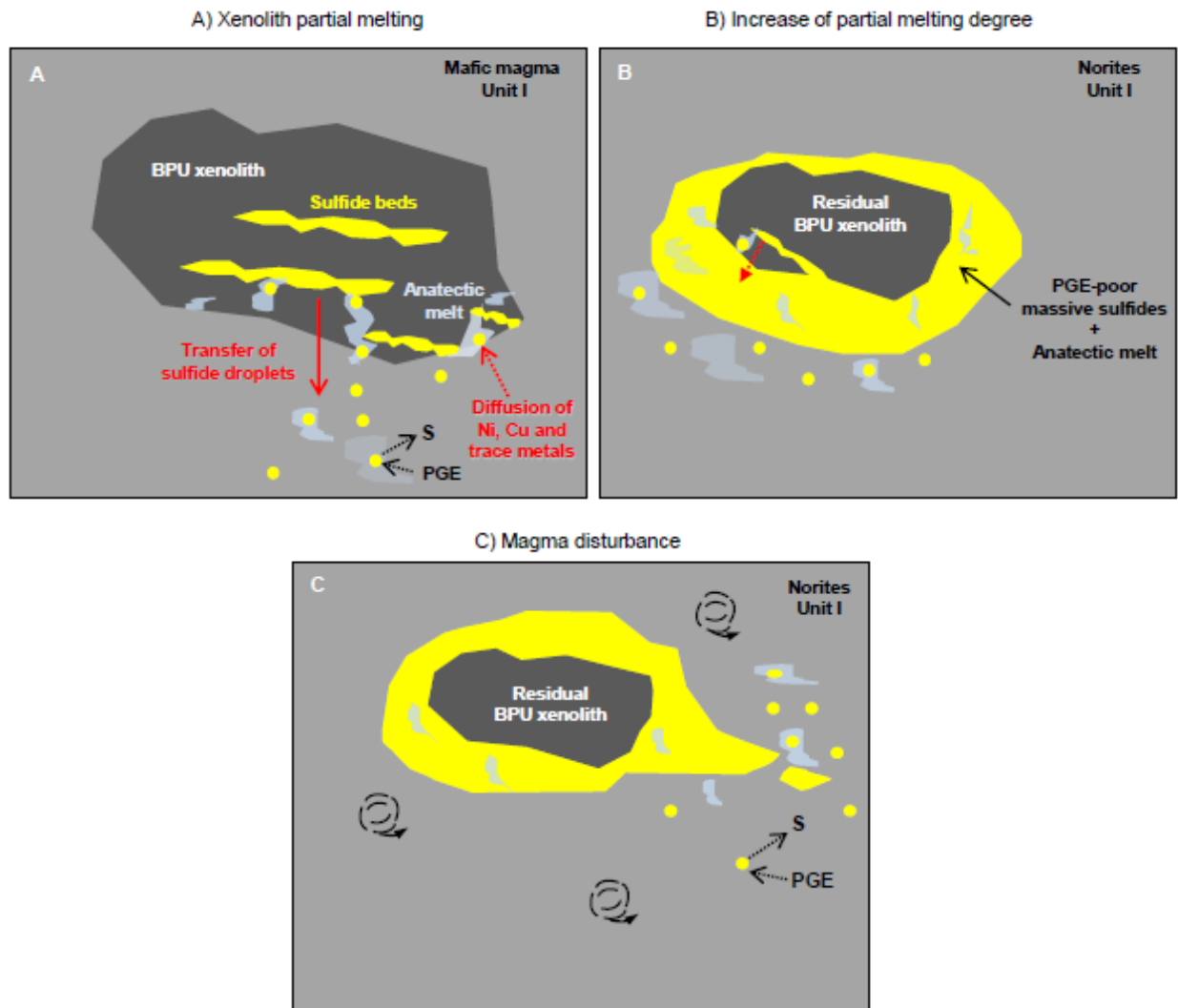


Figure 10

Fig. 1. Geological and location map of the Duluth Complex (modified from Queffurus and Barnes, 2014; Ojakangas et al., 2001; Miller et al., 2002). The right inset shows position of the Duluth Complex in the Midcontinent Rift System.

Abbreviations: PRI = Partridge River intrusion, SKI = South Kawishiwi intrusion. Ni-Cu-PGE deposits : 1, Water Hen; 2, Wyman Creek; 3, Wetlegs; 4, NorthMet (Dunka Road); 5, Mesaba (Babbitt); 6, Serpentine; 7, Dunka Pit; 8, Birch Lake; 9, Nokomis; 10. Diamond drill-hole cores are indicated by stars.

Fig. 2. Photomicrographs of Bedded Pyrrhotite Unit (BPU) Sulfides and silicates textures from the contact aureole and inside the intrusion.

A) Sulfide beds crosscut by anatectic melt in partially molten BPU from the contact aureole. B) Pyrrhotite bed in BPU metamorphosed from the contact aureole. Inset shows details for chalcopyrite grain in pyrrhotite bed. Notice that pentlandite and cubanite are absent of BPU country rocks. C) Sulfide bed in BPU xenolith within the intrusion. D) Delaminated sulfide bed in BPU xenolith. Chalcopyrite-rich patches close to the sulfide bed and sulfide microveinlets. Low angle interfaces occurs between sulfides and silicates. E) Massive sulfides surrounding BPU xenolith within the intrusion. F) Silicate mineralogical assemblage in a partially melted BPU xenolith. Quartz plate is added to better identify the inclusions of former anatectic melt.

Abbreviations (Whitney and Evans, 2010): Silicates: Bt = Biotite; Opx = Orthopyroxene; Crd = Cordierite; Pl = Plagioclase; Kfs = K-Feldspar; Qtz = Quartz. Sulfides: Ccp = Chalcopyrite; Cbn = Cubanite; Po = Pyrrhotite; Pn = Pentlandite. Gr = Graphite.

Fig. 3. Photomicrographs of partial melting textures of Bedded Pyrrhotite Unit (BPU) xenoliths and surrounding zones.

A) Detached pyrrhotite bed in BPU xenoliths in contact with melt filled space between the grains in BPU xenolith. Quartz plate is added to better show anatectic melt pockets. B) Sulfide droplets in pockets of anatectic melt pockets in same area of the photomicrograph A). C) Anatectic melt patches with trapped sulfide droplets in the anatectic melt-rich zone surrounding BPU xenolith. D) Rounded sulfide droplets trapped in xenolith anatectic melt patches from same area as the photomicrograph C). E) Films of anatectic melt that contained sulfide droplets in zones surrounding the BPU xenolith. F) Details of sulfide droplets trapped in small pockets of anatectic melt, same area than photomicrograph E).

Abbreviation: BPU = Bedded Pyrrhotite Unit.

Fig. 4. Plots of (A) $\delta_{34}\text{S}$ isotopic ratio, (B) Ni, (C) Cu and (D) Co vs. S.

Isotopic ratio $\delta_{34}\text{S}$ of mantle is comprised between 0 to 1‰. Averages of black shales, upper crust, picrites, and primitive mantle are shown for reference (Dionne-Foster, 2007; Lyubetskaya and Korenaga, 2007; Hu and Gao, 2008; Henrique-Pinto et al., 2016).

Abbreviation: BPU = Bedded Pyrrhotite Unit.

Fig. 5. Trace metals (A) Pd, (B) Ir, (C) Rh and (D) Ag vs. S for all rocks.

Averages of black shales, upper crust, picrites, and primitive mantle are shown for reference (Dionne-Foster, 2007; Lyubetskaya and Korenaga, 2007; Hu and Gao, 2008; Henrique-Pinto et al., 2016). Abbreviation: BPU = Bedded Pyrrhotite Unit.

Fig. 6. Semimetals (A) As, (B) Sb, (C) Bi and (D) Te vs. S for all rocks.

Averages of black shales, upper crust, picrites, and primitive mantle are shown for reference (Dionne-Foster, 2007; Lyubetskaya and Korenaga, 2007; Hu and Gao, 2008; Henrique-Pinto et al., 2016). Abbreviation: BPU = Bedded Pyrrhotite Unit.

Fig. 7. Mantle normalized chalcophile elements recalculated to 100 % sulfides and plotted in order of compatibility with picrite mantle (Barnes, 2016) of A) BPU rocks, norites and gabbronorites and B) massive sulfides. A) Note that the concentrations of the elements from

Co to Se increase from the contact aureole rocks to the norites through the BPU xenoliths with the highest concentrations in gabbronorites. Arsenic and Sb concentrations are the highest in the xenolith-hosted sulfides whereas the other sulfides from the other rock types contain similar to each other. Lead concentrations are similar in all rock types. B) Note that the massive sulfides contain similar concentrations of Co, Ni, Ir, Ru and Rh to the norite-hosted sulfides but are depleted in most other chalcophile elements. In addition the massive sulfides have negative Pt anomalies. Abbreviation: BPU = Bedded Pyrrhotite Unit.

Fig. 8. Variations of semimetals and PGE contents, A) As, B) Sb, C) Pd and D) Ir in 100% sulfides with distance from contact in the B1-384 borehole, Mesaba deposit.

Contact between Virginia Formation sedimentary rocks and mafic magma corresponds to distance zero. Arsenic, Sb, Pd and Ir values are plotted for rocks in the basal unit of the Partridge River Intrusion, i.e. Unit I. Semimetals contents of sulfides decreases with distance from the Bedded Pyrrhotite Unit xenoliths whereas Pd and Ir contents of the norite and gabbronorite-hosted sulfides increase with distance from xenoliths. Abbreviation: BPU = Bedded Pyrrhotite Unit.

Fig. 9. Plots of Cu vs. A) $\delta_{34}\text{S}$ isotopic ratio, B) Pd, C) Bi, D) Te, E) As and F) Sb in 100% sulfides of rocks within the intrusion and results of the modeling.

Red line represents interaction model from calculation based on equations of Leshner and Burnham, 2001. Numbers along the line represent different R factor values. Models for As and Sb are plotted for partition coefficients at 900°C and 1000°C. Abbreviation: BPU = Bedded Pyrrhotite Unit.

Fig. 10. Proposed model for S and semimetals contamination of the mafic magma.

A) Partial melting of Bedded Pyrrhotite Unit xenoliths in the magma. Sulfides are transferred to the mafic magma via anatectic silicate melt. This transfer leads to S and semimetals contamination of the mafic magma. In addition, diffusive transfer of Ni, Cu and trace metals from the mafic magma to the Bedded Pyrrhotite Unit xenolith occurs. Low R factor are recorded in the mafic magma. B) Massive sulfide accumulation after increase of partial melting degree. Massive sulfides are PGE-poor because of lack of interaction with the magma. C) Magma disturbance, i.e. new injection of magma or possibly related to seismic fault activity, results in sulfide transportation. Platinum-group element enrichment of sulfide droplets occurs because of high R-factor of the mafic magma. Large-scale sulfur and semimetals contamination of the mafic magma occurs in the basal unit. Hence, platinum-group minerals are present in gabbronorites in the whole basal unit I. Abbreviations: BPU = Bedded Pyrrhotite Unit; PGE = Platinum-group elements.

	As (ppm)	Sb (ppm)	Bi (ppm)	Te (ppm)
BPU average – This study	38	4.1	0.6	0.4
SDO-1 average (Henrique-Pinto et al., 2016)	62.6	4.11	0.27	0.131
SCHS-1 average (Henrique-Pinto et al., 2016)	50.1	0.072	0.102	-
SBC-1 average (Henrique-Pinto et al., 2016)	29.4	1.22	0.6	0.184
SH-1 average (Henrique-Pinto et al., 2016)	22.5	1.17	1.19	0.198
Median black shale (Ketris and Yudovich, 2009)	30	5	1.1	2
Upper crust average (Hu and Gao, 2008)	5.7	0.75	0.23	0.027
MORB average (Arevalo and McDonough, 2010)	0.11	0.014	0.01	0.005
Picrites average (Dionne-Foster, 2007)	0.45	0.02	0.01	0.02
Primitive mantle average (Lyubetskaya and Korenaga, 2007)	0.05	0.007	0.004	0.008

Table 1. Average semimetals contents of the Bedded Pyrrhotite Unit in comparison with average and median values of black shale standards (SDO-1, SCHS-1, SBC-1 and SH-1), black shale, upper crust, MORB, picrites and primitive mantle.

Abbreviations: BPU=Bedded Pyrrhotite Unit; MORB=Mid-oceanic ridge basalt; SDO-1=Devonian Ohio Shale; SCHS-1= Carbonaceous black shale; SBC-1= Brush Creek Shale; SH-1= In-house black shale.

Table 2. Average values of chalcophile elements in the Partridge River Intrusion and the contact aureole rocks, recalculated to 100 % sulfides. Abbreviations: BPU=Bedded Pyrrhotite Unit; c.a.=contact aureole; n.d.=not determined.

	M	As	Sb	Pb	Bi	Se	Ag	Cu	Au	Te	Pd	Pt	R	R	O	Ni	Co
	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppm	ppb	ppm	ppb	ppb	ppb	ppb	ppb	ppm	ppm
BPU c.a. (n=11)	19 7	27 3		22 1		3 24	11	1842	40	2	17	43			0. 2	138 1	21 7
BPU xenoliths (n=26)	12 8	54 0		16 5		8 20	40	1669 8	669	3	240	14 9				746 8	13 22
Massive sulfides (n=6)		10 52						2816 3	198 2							189 51	20 91
Norites (n=32)	28 8	25 5		27 5		12 43	41	7184 9	311 9		329 6	17 22				228 26	25 96
Gabbrono rites (n=27)	n.d. .	14 7		n.d. .		7 80	61	1283 00	126 75		159 06	37 29	45 2	33 7		445 77	28 72

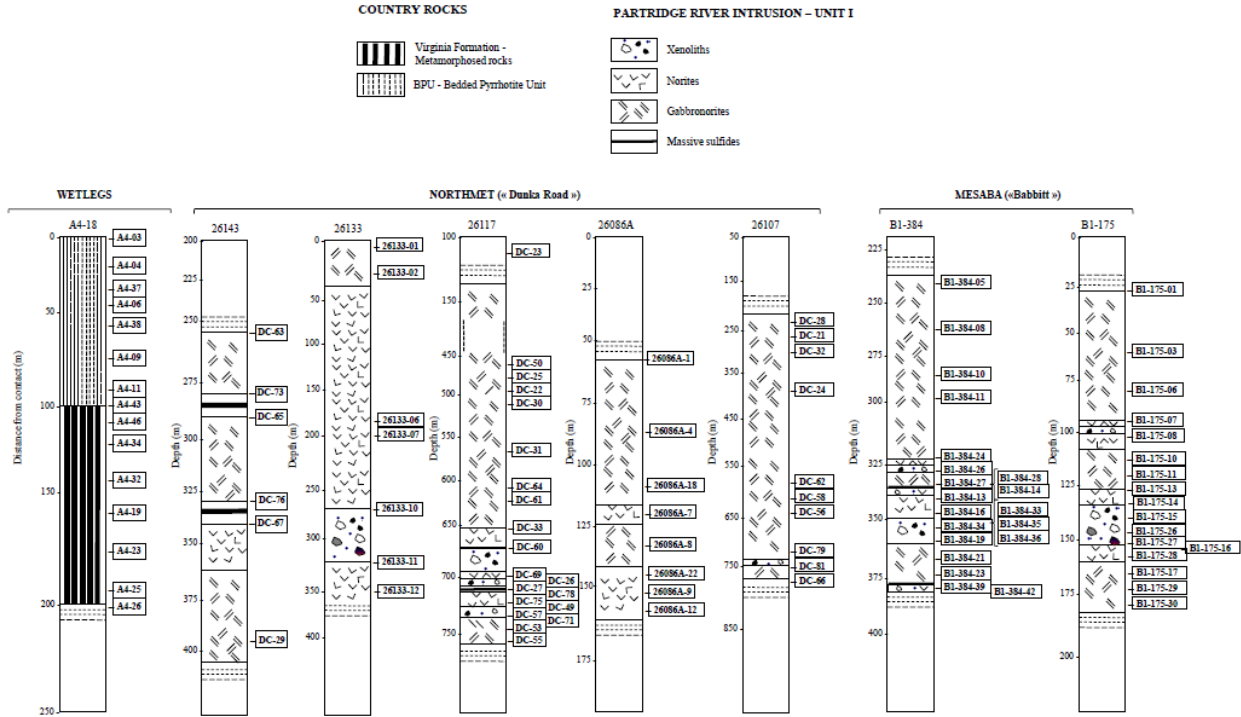
Table 3. Results of modeling of compositions of sulfides.

R Factor:	As (ppm)	As (ppm)	Sb (ppm)	Sb (ppm)	Bi (ppm)	Te (ppm)	Cu (ppm)	Pd (ppb)	$\delta^{34}\text{S}$ (‰)		
0	837.0	837.0	0	0	73.00	73.00	3.00	3.00	17621	263.0	13.40
2	832.3	698.2	5	5	72.56	60.87	3.01	3.04	17903	278.8	13.37
4	827.7	599.1	6	4	72.12	52.20	3.03	3.08	18185	294.5	13.34
10	814.3	420.7	5	5	70.84	36.60	3.07	3.19	19026	341.8	13.24
50	736.7	143.2	1	5	63.43	12.33	3.33	3.96	24508	656.6	12.63
500	398.2		5	20.82	31.13	1.63	5.33	11.70	74097	8	8.48
1000	297.0		0	12.74	21.46	0.92	6.50	18.82	11174	8053.	6.35
2000	226.5		7	8.64	14.74	0.56	7.67	29.77	15881	15674	4.41
3200	195.1		7	7.09	11.74	0.43	8.33	39.16	19139	24604	3.38
6400	166.4		3	5.80	9.00	0.31	9.05	54.08	23276	47329	2.32
Concentrations in sedimentary sulfides	837	837	73	73	3	3	17621	263	-	-	-
Concentrations in silicate magma	0.45	0.45	0.02	0.02	0.01	0.02	150	7.88	-	-	-
Dsulfide/silicate	300	10	300	10	1000	4500	2000	90000	-	-	-
Temperature (°C)	900	1100	900	1100	-	-	-	-	-	-	-
$\delta^{34}\text{S}$ magma	-	-	-	-	-	-	-	-	-	-	1
$\delta^{34}\text{S}$ sediments	-	-	-	-	-	-	-	-	-	-	13.4

Calculations are based on equation 5 in Lesher and Burnham (2001). Grey section corresponds to final concentrations of elements in the sulfide magma after modeling. Values for model calculations are concentrations of elements in the sedimentary sulfides, i.e. average of the BPU xenolith concentrations; concentrations of elements in silicate magma, i.e. average of picrites concentrations (Dionne-Foster, 2007); concentrations of elements in the silicate magma, i.e. average of picrite concentrations (Dionne-Foster, 2007) and partition coefficients between sulfide and silicate melts (Li and Audétat, 2012; Kiseeva and Wood, 2013; Patten et al., 2013; Brenan, 2015; Li and Audétat, 2015; Liu and Brenan, 2015). Abbreviations: BPU=Bedded Pyrrhotite Unit.

Appendix A. Stratigraphic position of the samples in the boreholes sections of Virginia Formation country-rocks and the basal part (Unit I) of the Partridge River Intrusion (modified from Queffurus and Barnes, 2014). The few samples taken from Dunka Pit, A4-15, 26014, 26015, B1-46, B1-129, CN-7 and LE-3 are not indicated.

Abbreviations: PRI = Partridge River intrusion, BPU = Bedded Pyrrhotite Unit



Appendix B1. Whole-rock data for the Bedded Pyrrhotite Unit from the contact aureole. xenoliths of the Bedded Pyrrhotite Unit. norites and gabbro-norites. Abbreviations: BPU=Bedded Pyrrhotite Unit; c.a.=contact aureole.

Sa m pl es	Roc k typ e	De po sit	S and $\delta^{34}\text{S}$		Semimetals				Major and trace elements													
			S (%) S (%) V C D T)	δ 3 4 S (%) V C D T)	A s (p p m)	S b (p p m)	B i (p p m)	T e (p p m)	N i (p p m)	C u (p p m)	C o (p p m)	A g (p p m)	S e (p p m)	Pb (p p m)	M o (p p m)	O s (p p b)	Ir (p p b)	R u (p p b)	R h (p p b)	P t (p p b)	P d (p p b)	A u (p p b)
D C- 70	BP U c.a.	Du nk a Pit	4 . 4 5	1 6. 3	1 5. 1 6	1. 5 3 8	0. 2 8	< 0. 6	1 3 0 0	1 4 0	2 6	1. 1 0 0	1. 6 0 0	24 .5 2	2 8. 0 0	< 1 .1 5	0 .1 0	< 5 .1 0	0 .1 1	4 .9	0. 8	1 .7
B 1- 33 1- 01	BP U c.a.	M esa ba	4 . 3 0	-	2 9. 0 1	0. 5 5 2	0. 0 2	0. 1 9	1 2 0 0	1 7 0	2 2	1. 7 0	-	20 .0 0	2 0. 7 2	0 .3	-	0 .2 2	0 .2 1	3 .1	2. 2 3	2 .3
B 1- 33 8- 01	BP U c.a.	M esa ba	5 . 6 0	-	1 3. 7 6	2. 5 5 3	0. 0 3 8	0. 0 8	9 0 0 0	2 9 5 0	2 5	1. 5 5 0	-	42 .6 5	3 0. 0 0	0 .3 1	0 .1 2	0 .2 2	0 .2 7	9 .7	1. 2 7	2 .7
A 4- 01	BP U c.a.	W etl eg s	6 . 5 3	-	1 0. 0 0	2. 6 0 0	0. 2 0	-	1 8 7	1 9 7	2 8	2. 1 0	-	37 .0 0	3 7. 0 0	0 .2 1	0 .1 2	0 .7	0 .3	9 .4	2. 4 8	1 5 .8
A 4- 03	BP U c.a.	W etl eg s	2 . 2 6	-	3 1. 5 7	1. 5 2 4	0. 2 4 6	< 0. 6	2 8 9	3 1 3	2 7	1. 2 9	1. 1 1	12 .6 4	-	-	-	-	-	-	-	-
A 4- 04	BP U c.a.	W etl eg s	6 . 3 0	2 0. 4	2 2. 4 0	5. 8 5	-	-	1 4 0	2 0 0	-	-	2. 4 1	-	-	-	-	-	-	-	-	-
A 4-	BP U	W etl	4 . 6.	1 3.	1 3.	2. 2 1	0. -	-	8 0	1 4	2 2	1. 1	3. 1	35 .0	2 6.	0 .1	0 .1	0 .1	0 .1	-	1. 8	7 .

09	c.a.	eg s	1 9	7	0 0	0	2			0		0	1	0	0 0	4	1	2	3			1
A 4- 11	BP U c.a.	W etl eg s	6 .8 6 7	1 8. 3	2 2. 8 6	1. 9	0. 4	0. 5	1 5	2 6	3 4	0. 4 9	4. 4 0	38 .0 0	2 3. 8 7	< 0	-	0 .1 1	0 .1 1	6 .1 1	1. 2	3 .5
A 4- 12 A	BP U c.a.	W etl eg s	5 .8 5	-	2 7. 0 0	1. 9	0. 1	-	1 5	3 1	2 9	1. 6 0		27 .0 0	3 1. 0	0 .2	0 .1	0 .2	0 .3	6 .2	8. 0	8 .5
A 4- 38	BP U c.a.	W etl eg s	2 .2 3	1 9. 5	3 1. 4 0	1. 3	-	-	1 0	1 3	-	-	2. 2 3	-	-	-	-	-	-	-	-	-
A 4- 43	BP U c.a.	W etl eg s	1 .4 5	-	3 5. 0 0	1. 1	-	-	6 7	5 9	-	-	1. 4 0	-	-	-	-	-	-	-	-	-
E C- 07 -A	BP U xen olit h	Du nk a Pit	2 .5 0	-	2 2. 0 0	3. 0	2. 2	0. 8	7 3	3 9	8 6	1. 4 0	2. 4 0	15 .8 8	2. 7 5	-	0 .3	1 .1	1 .3	2 .7	5 8. 4	2 .6 6
B 1- 38 4- 14	BP U xen olit h	M esa ba	7 .1 3	9. 5	4 7. 0 0	1. 0	1. 5	0. 0	2 7	4 1	3 0	1. 4 0	6. 1 5	20 .4 9	2 5. 0	0 .3	-	0 .2	0 .3	1 .3	1 4. 8	6 .2
B 1- 38 4- 18	BP U xen olit h	M esa ba	7 .7 0	-	6 3. 2 2	1 0. 4 7	-	-	2 1	4 6	2 1	4. 2 4	-	-	-	< 1 .2	0 .6	< 7 .3	2 .1	4 .0	2 7. 4	1 4. 8
B 1- 38 4- 19	BP U xen olit h	M esa ba	5 .4 8	1 1. 5	1 1. 2 4 2	4. 8	0. 4	0. 6	6 0	1 9	5 4	7. 4 5	2. 4 5	37 .3 9	-	0 .4	0 .1	3 .3	< 0 .1	0 .9	1 2. 5	1 .1 1
B 1- 38 4- 20 7	BP U xen olit h	M esa ba	1 .2 4	-	2 4. 9 5	6. 3	-	-	4 5	9 8	4 8	-	-	-	-	0 .2	-	2 .9	< 2 .2	1 .4	< 9. 1	2 4. 2
B 1-	BP U	M esa	3 .8	1 8.	2 2.	0. 4	0. 4	0. 2	9 0	3 1	1 5	1. 0	3. 7	19 .0	2 4.	< 0	0 .1	2 .1	1 .1	9 .1	4 1.	1 6

38 4- 26	xen olit h	ba	7 4	1	0 0	0 0	0 6	0 0	0 0	2 0	0 1	1 1	2 8	. 6	2	6	0	3	7	. 2		
B 1- 38 4- 33	BP U xen olit h	M esa ba	5 .3 3	1 0. 7	3 6. 4	2. 7 0	0. 7 0	0. 4 5	1 3 2	3 1 7	1 3 9	2. 2 0	2. 2 5	44 .0 0	2 0. 4	0 . 3	- . 2	0 . 2	- . 4	1 . 4	1 1. 5	4 . 3
B 1- 38 4- 34	BP U xen olit h	M esa ba	5 .2 2	1 1. 7	4 8. 0	2. 4 2	0. 5 0	0. 6 3	5 3 5	1 2 2	3 2 1	5. 9 4	2. 7 3	29 .6 8	0 . 2	-	2 . 8	0 . 2	5 1 4	3. 9 .	2 . 0	
B 1- 38 4- 35	BP U xen olit h	M esa ba	6 .3 8	1 2. 2	4 0. 0	2. 8 0	1. 0 0	0. 4 6	4 3 0	1 9 2	7 3 6	1. 3 0	2. 4 3	31 .0 0	2 4. 5	0 . 3	-	-	-	0 . 7	3. 5 9	2 . 9
B 1- 38 4- 36	BP U xen olit h	M esa ba	6 .3 6	1 4. 2	2 3. 0	2. 8 0	0. 8 0	0. 2 6	4 1 0	2 0 6	6 4 7	1. 3 0	2. 8 4	17 .8 5	1 8. 5	0 . 3	-	0 . 2	0 . 1	1 . 7	9. 6 3	6 . 3
B 1- 46 - 02	BP U xen olit h	M esa ba	0 . 4 7	-	3 5. 1	6. 2 4	-	-	1 2 5	1 4 4	1 8	1. 4 4	-	-	-	< 0 . 2	-	< 2 . 7	< 0 . 2	2 . 9	2. 7 9	4 . 9
B 1- 46 - 05 a	BP U xen olit h	M esa ba	2 . 7 2	-	3 4. 1 5	8. 8 8	-	-	2 6 4	4 3 8	2 3	1. 7 3	1. 4 3	-	-	0 . 6	-	2 . 5	0 . 2	1 . 7	3. 5 9	2 . 9
B 1- 46 - 11	BP U xen olit h	M esa ba	4 . 5 5	-	-	-	-	-	-	-	-	-	2. 1 9	-	-	-	-	-	-	-	-	-
26 01 4- 01 a	BP U xen olit h	No rth M et	1 . 2 4	-	1 8. 5 8	4. 8 4	-	-	1 9 5	2 5 1	4 6	-	1. 1 5	-	-	< . 3 5	0 . 1	4 . 0	0 . 5	2 . 2	2. 6 0	6 . 0
26 08	BP U	No rth	3 .	-	3 6.	2. 3	-	-	9 7	1 1	8 2	2. 3	-	-	-	< 0 .	0 . 5	4 . 4	< 1	9 3.	2 .	

6 A- 11	xen olit h	M et	8 1		4 2	3			3	2 3		1				. 2	6	3	5	. 6	9	7
D C- 26	BP U xen olit h	No rth M et	0 . 7 0	-	1 4. 4 3	1. 0 6	-	-	1 5 0	4 2 2	1 6 6	-	-	-	-	0 . 7	0 . 3	< 2 . 9	0 . 7	6 . 0	2. 9	1 4 4
D C- 69	BP U xen olit h	No rth M et	5 . 9 0	1 6. 9	6 5. 2 0	< 0. 2	< 0. 1	0. 4 8	6 7 0	9 3 0	1 5 7	0. 5 0	1. 9 1	0. 15 8. 4 3	2 . 8.	1 . 2	0 . 8	7 . 5	1 . 4	1 . 3 . 1	3 6. 1	2 8 2
D C- 71	BP U xen olit h	No rth M et	1 0 . 2 0	1 5. 9	9. 0 0	0. 5 0	< 0. 1	-	2 8 0	1 0 6 0	4 4 0	-	0. 1 3	19 . 0	1 1. 0 0	< 0 . 6	0 . 3	0 . 3	1 . 3	3 . 6	4 0. 7	4 3 . 0
D C- 78	BP U xen olit h	No rth M et	1 . 5 6	-	5. 9 9	0. 2 0	-	-	3 5 5	7 3 2	6 9	-	-	-	-	0 . 7	0 . 2	1 . 7	0 . 4	9 . 0	5. 3	1 2 2
D C- 81	BP U xen olit h	No rth M et	0 . 4 0	-	8 2. 9 0	1. 0 7	-	-	2 9 1	4 2 0	8 3	-	-	-	-	0 . 6	0 . 3	< 2 . 6	0 . 3	7 . 4	8. 3	7 2
B 1- 17 5- 08	BP U xen olit h	Se rpe nti ne	1 . 2 9	-	1 1. 4 1	4. 2 3	-	-	4 1 4	6 6 2	6 7	8. 7 2	0. 7 8	-	-	0 . 5	0 . 5	8 . 5	1 . 5	9 . 0	1 1. 2	4 . 8
B 1- 17 5- 15	BP U xen olit h	Se rpe nti ne	4 . 5 3	-	1 7 8. 8 7	9. 4 6	-	-	8 6 7	1 9 2 8	7 0	5. 2 5	1. 6 2	-	-	0 . 4	0 . 1	3 . 9	0 . 4	4 . 8	1 7. 9	1 3 . 6
B 1- 17 5- 26	BP U xen olit h	Se rpe nti ne	4 . 8 8	-	5 0. 5 8	8. 2 8	-	-	6 9 7	1 5 9 0	5 5	4. 5 6	1. 0 5	-	-	0 . 3	0 . 1	1 . 9	8 . 8	0 . 9	8. 2	6 . 9
B 1- 17	BP U xen	Se rpe nti	5 . 4	-	7 3. 3	5. 4 8	-	-	1 0 8	2 1 9	1 3 3	0. 8 9	1. 4 2	-	-	0 . 3	0 . 1	1 . 0	0 . 2	2 . 6	2 3. 0	9 . 4

5-27	olih	ne	8		6				6	6							1					
26 13 3- 11	BP U xen olih	W ym an crek	1 . 1 9	- 2. 1 9	2 2. 5	- - -	- -	3 6 6	2 0 2	6 6	- -	- -	- -	- -	- -	0 . 2	8 . 6	1 . 8	3 3 . 4	8. 7	1 3 . 4	
B 1- 38 4- 28	Ma ssiv e sulf ides	M esa ba	1 9 . 7 5	1 2. 0	- -	- -	- -	- -	- -	- -	- -	2 1. 4 4	- -	- -	- -	- -	- -	- -	- -	- -	- -	
B 1- 38 4- 39	Ma ssiv e sulf ides	M esa ba	1 6 . 3 0	1 2. 3	4 3. 9 7	0. 0 8	0. 1 7	2. 5 7	9 8 0 0	3 0 1 8	9 6 0 0	6. 8 4. 5 1	2. 79	5. 0 0	2 3. 0 0	8 . 5 0	8 . 0	3 9 . 5 9	1 4 . 9	5 . 4	6 3. 1 1	2 2 . 1
D C- 73	Ma ssiv e sulf ides	No rth M et	1 9 . 6 4	8. 4	3 3. 2 7	0. 1 9	0. 0 2	0. 6 4	1 0 6 1 1	1 2 9 5 4	1 0 9 7	1. 1 0 0	2 5. 0 0	- -	- -	1 . 2	3 8 . 2	2 6 . 9	1 1 . 6 4	8 . 8 9	4 6 . 9	2 8 6
D C- 74	Ma ssiv e sulf ides	No rth M et	2 2 . 2 0	- 2	2 5. 2 2	0. 3 7 0	1. 0 2	0. 8 5	1 4 5 1 8	1 8 2 0 6	1 5 8 6	1. 9 2. 0 0	2 0	- -	- -	3 . 5	9 . 3	1 8 . 9 0	3 2 . 5 5	3 2 . 6 5	6 2 . 6 5	3 4 7 1
D C- 75	Ma ssiv e sulf ides	No rth M et	3 1 . 2 0	1 6. 0	1 7 4. 7 9	0. 3 9	0. 5 9	1. 0 7	6 8 4 0 9	1 1 3 4 5 9	1 2 4 5 0	1 1 3. 0 0	- -	- -	3 . 8	2 . 3	< 9 . 3 8	1 1 . 8	2 3 . 2 7	8 7 . 6 0	1 3 . 1 0	
D C- 76	Ma ssiv e sulf ides	No rth M et	3 4 . 9 0	1 2. 0	1 4. 3 9	2. 3 9	1. 3 6	0. 8 5	1 8 7 3 3 1	5 4 2 3 3	- -	- -	4 3. 0 0	- -	- -	4 . 6	2 7 . 0 9	3 9 . 6 9	1 1 . 6 9	3 4 . 6 0	1 5 9 0 0	< 2
B 1- 38 4- 04 a	nori te	M esa ba	0 . 0 3	- 5	< 0. 2	< 0. 1	< 0. 0 6	< 0. 0 6	2 6 5 0 0	9 5 0 0	1 8	- -	1. 4 7	7. 00	2. 0 0	0 . 8 6	0 . 6	2 . 5 2	1 . 2	9 . 0	1 0. 9	1 . 2
B 1- 38	nori te	M esa ba	1 . 2	- 2. 0	2 5 0	< 0. 1	< 0. 0	1 5 7	6 5 0	1 3 4	1. 0 0	3. 0 8	17 . 0	< 2	0 . 6	0 . 9	4 . 7	2 . 1	2 . 2	4 7. 9	1 5 .	

4-12			7		0			6	0	0									7		5	
B 1-38 4-13	nori te	M esa ba	1 . 7 2	1 3. 2	8. 2 0	1. 2 0	- - -	- - -	7 8 6	2 1 6	1 0 8	2. 2 7	1. 8 7	- - -	< 0 . 4	0 . 3	4 . 4	0 . 9	1 0 3	2 7. 1	1 2 8	
B 1-38 4-15	nori te	M esa ba	1 . 5 8	-	8. 4 5	1. 6 5	- - -	- - -	1 1 0	4 8 6	1 2 5	2. 2 6	- - -	- - -	< 0 . 4	0 . 7	3 . 4	1 . 8	2 3 5	5 2. 4	2 7. 5	
B 1-38 4-16	nori te	M esa ba	7 . 1 7	1 2. 5	2 9. 0	< 0. 2	0. 2 0	- - -	4 7 0	1 8 3	8 1 0	0. 2. 5	2. 9 8	8. 00	< 2	0 . 7	2 . 1	5 . 1	4 . 3	9 . 3	3 6. 8	3 3 3
B 1-38 4-24	nori te	M esa ba	3 . 1 7	-	9. 1 2	1. 9 1	- - -	- - -	9 7 6	2 2 8	1 4 4	0. 6 6	4. 1 1	- - -	0 . 8	1 . 3	5 . 2	2 . 9	1 3 6	4 3. 1	1 8 7	
B 1-38 4-27	nori te	M esa ba	2 . 3 1	-	2 1. 4	0. 8 6	- - -	- - -	2 1 7	2 9 4	1 7 2	3. 7 2	- - -	- - -	< 1 . 4	2 . 5	7 . 5	7 . 0	1 0 5	1 1 2	4 8 0	
B 1-38 4-30	nori te	M esa ba	2 . 2 9	-	9. 0 9	1. 6 4	- - -	- - -	1 2 4	4 3 6	1 3 4	2. 4 4	- - -	- - -	0 . 6	0 . 7	2 . 3	1 . 9	2 5 2	6 4. 3	2 3 9	
B 1-38 4-42	nori te	M esa ba	6 . 3 1	1 2. 2	6. 5 2	0. 1 7	0. 1 7	1. 0 7	3 5 3	1 1 6	2 5 7	3. 9 0	- - -	- - -	2 . 9	2 . 3	1 . 1	5 . 1	1 9 0	9 2. 0	2 2 1	
26 08 6 A- 05	nori te	No rth M et	2 . 8 0	-	2 8. 0	1. 8 8	- - -	- - -	1 2 6	4 8 9	1 3 5	7. 3 3	- - -	- - -	0 . 7	1 . 5	3 . 7	1 . 6	5 3 0	3 4 3	7 6 6	
26 08 6 A-	nori te	No rth M et	3 . 6 1	-	5 1. 1	2. 2 3	- - -	- - -	1 2 0	3 1 3	1 0 1	2. 2 3	- - -	- - -	1 . 0	0 . 3	9 . 7	1 . 5	1 3 3	9 4. 2	4 8 2	

06																						
26 08 6 A- 07	nori te	No rth M et	1 . 6 2	-	1 4. 3 0	1. 6 9	-	-	1 0 3 1	3 6 1 5	1 1 2	4. 3 9	2. 0 4	-	-	1 . 0	1 . 3	4 . 9	< 0 . 5	1 4 1 8	2 8 2. 3	4 8 . 2
26 08 6 A- 09	nori te	No rth M et	1 . 5 1	-	2 1. 2 4	3. 1 8	-	-	7 2 4	1 6 1 5	8 5	1. 6 5	1. 3 3	-	-	< 0 . 8	0 . 8	2 . 1	3 . 5	1 8 0 . 6	1 4 3. 3	1 7 . 7
26 08 6 A- 10	nori te	No rth M et	5 . 4 9	-	3 3. 0 7	2. 5 0	-	-	2 2 6 4	6 0 2 0	2 5 2	2. 1 0	-	-	-	2 . 7	9 . 7	7 . 8	1 6 . 3	2 4 5 . 4	2 1 2 6. 4	3 1 . 9
26 08 6 A- 12	nori te	No rth M et	0 . 9 1	-	3. 7 9	2. 4 1	-	-	5 6 0 9	2 4 9	7 4	2. 4 0	1. 6 6	-	-	< 0 . 5	0 . 6	4 . 2	3 . 1	3 8 . 6	1 5 1. 2	3 0 . 8
26 08 6 A- 22	nori te	No rth M et	3 . 3 6	-	4 1. 0 0	1. 9 7	-	-	1 2 1 2	3 4 3 8	1 4 9	2. 6 9	2. 5 8	-	-	< 0 . 5	1 . 0	5 . 2	4 . 4	3 5 . 2	1 7 1. 7	2 9 . 3
D C- 27	nori te	No rth M et	0 . 7 7	1 4. 5	1 0. 0	< 0. 2	0. 1 1	-	3 3 0	8 6 0	1 0 1	0. 7 0	0. 9 2	6. 00	6. 0 0	< 0 . 6	0 . 3	< 2 . 6	0 . 9	1 . 8	2 2. 7	-
D C- 49	nori te	No rth M et	0 . 2 0	8. 2	1. 9 6	1. 5 9	-	-	1 5 1	3 7 3	5 2	-	-	-	-	< 0 . 5	0 . 1	< 5 . 0	< 0 . 2	2 . 5	5. 4	< 2
D C- 52	nori te	No rth M et	0 . 6 6	1 3. 7	6. 2 8	1. 3 5	-	-	1 8 6	2 9 3	4 0	-	-	-	-	< 0 . 9	0 . 3	4 . 9	0 . 7	3 . 7	1 1 . 1	< 2
D C- 54	nori te	No rth M et	0 . 3 1	1 2. 5	2. 7 7	0. 8 0	-	-	2 6 5	6 2 2	6 2	-	-	-	-	0 . 7	0 . 2	< 5 . 0	0 . 6	6 . 9	8. 1	2 3
D C- 57	nori te	No rth M	1 . 4	-	2 8. 9	2. 4 7	-	-	1 8 8	8 1 4	9 6	3. 8 8	-	-	-	2 . 5	7 . 6	7 . 7	1 6 . .	5 5 4	4 2 3.	3 0 9

		et	9		6				5	0							9	.5	5			
D C- 60	nori te	No rth M et	1 .5 7	1 0. 4	3. 4 2	0. 5 7	- -	- -	1 1 4 0 0	3 0 7 5	1 2 7	0. 7 5	- -	- -	< 1 .3	1 .2 0	< 5 .8	4 .8 1	5 0 2 7	1 0 4. 7	1 8 5	
D C- 67	nori te	No rth M et	3 .4 2	5. 3	6. 8 9	0. 6 2	5. 1 9	2. 1 0	1 5 7 0 6	1 3 9 8	1 3 8	2. 0 3	- -	- -	4 .1	2 .0 3	5 .3 1	7 .1 5	8 3 7 4. 2 3	9 7 6 2 3	2 7 3	
D C- 68	nori te	No rth M et	3 .3 9	-	2. 1 1	0. 1 0	0. 8 2	0. 4 2	1 5 5 3	1 4 0 7 5 2	1 3 7	2. 4 8	- -	- -	2 .9	1 .7	< 5 .3	7 .2	6 4 2 2 .4	9 3 2. 3 3	2 1 3 3	
B 1- 17 5- 03	nori te	Se rpe nti ne	3 .9 8	-	1 7. 5 2	3. 5 3	- -	- -	1 6 8 0	5 4 2 4	1 7 9	4. 2 1	4. 5	- -	- -	0 .5	0 .9	6 .9 0	2 .0 0	2 6 0 3	1 0. 1	4 7 1
B 1- 17 5- 07	nori te	Se rpe nti ne	7 .0 8	1 0. 6	1 3. 7 8	0. 2 5 6	0. 4 6	2. 3 4	3 0 3 7	5 3 9 7	3 4 4	2. 5 6	5. 7 7	- -	- -	0 .8	1 .5	3 .5 0	2 .0 7	6 4. 3	8 3 9	2 3 9
B 1- 17 5- 09	nori te	Se rpe nti ne	8 .3 5	-	2 6. 5 3	0. 4 9	- -	- -	3 7 6 8	5 4 9 1	4 0 4	2. 6 9	4. 8 7	- -	- -	2 .5	2 .3	4 .1 7	4 .7 6	1 0 9 1 6	5 9. 1 1	4 5 1
B 1- 17 5- 10	nori te	Se rpe nti ne	3 .1 7	-	1 5. 3 3	1. 5 1	- -	- -	1 8 6 7	5 4 5 0	1 7 0	3. 7 1	5. 0 1	- -	- -	0 .7	1 .2	3 .5 5	3 .5 8	3 8 8. 1	8 0 2	9 0 2
B 1- 17 5- 13	nori te	Se rpe nti ne	4 .7 8	-	1 4. 5 8	1. 6 4	- -	- -	1 5 1 1	2 4 7 1	2 3 3	-	5. 0 3	- -	- -	1 .8	1 .1	5 .8 3	2 .3 2	1 6 2	4 9. 4	1 9 3
B 1- 17 5- 14	nori te	Se rpe nti ne	2 .4 0	1 5. 3	1 0. 7 0	2. 5 7	- -	- -	1 1 1 4	1 2 2 6	1 7 7	-	2. 1 4	- -	- -	0 .3	0 .3	< 3 .9	0 .9 6	2 .6 8	1 3. 8	7 .6 6

B 1- 17 5- 28	nori te	Se rpe nti ne	0 . 6 8	-	6. 3 5	1. 6 3	-	-	3 6 4	5 3 7	8 1	-	1. 0 8	-	-	0 . 5	0 . 2	2 . 3	0 . 5	2 . 3	7. 1	< 1 . 0
C N- 7- 3	nori te	Wate r- he n	0 . 2 7	-	4. 6 1	0. 2 5	-	-	4 9 7	6 7 3	1 1 5	-	-	-	-	1 . 4	0 . 1	9 . 4	0 . 6	1 0 . 5	5. 9	3 . 8
C N- 7- 6	nori te	Wate r- he n	1 . 8 0	-	2 5. 0 6	0. 2 1	-	-	1 0 0 2	1 1 2 1	1 4 2 2	1. 3 5	-	-	-	-	-	-	-	-	-	-
26 13 3- 06	nori te	Wym an crek	1 . 2 9	-	0. 6 5	0. 0 7	-	-	2 0 7 1	6 5 2 5	1 1 2 2	2. 9 9	-	-	-	0 . 9	3 . 4	4 . 3	4 . 3	4 3 . 2	8. 7	1 6 . 9
26 13 3- 07	nori te	Wym an crek	0 . 8 1	-	0. 9 0	0. 0 5	-	-	1 4 6 2	3 8 2 8	1 3 9	0. 4 3	-	-	-	1 . 4	1 . 3	7 . 8	4 . 1	5 6 . 7	1 1 . 2	4 . 2
26 13 3- 10	nori te	Wym an crek	0 . 1 4	-	2. 9 2	0. 1 9	-	-	7 0 3	6 3 0	1 4 4	-	-	-	-	1 . 7	1 . 0	5 . 4	3 . 3	3 7 . 8	3 8. 7	4 . 2
26 13 3- 12	nori te	Wym an crek	0 . 1 1	-	0. 8 6	0. 1 2	-	-	2 7 9	3 0 9	6 2	-	-	-	-	1 . 3	0 . 8	7 . 1	1 . 8	1 8 . 2	1. 7	1 . 2
L E- 3- 01	Ga bbr ono rite	Lo ng ear	0 . 0 2	-	0. 1 9	0. 0 4	-	-	4 3 0	4 8	6 1	-	-	-	-	1 . 1	0 . 1	6 . 7	0 . 6	7 . 8	3. 8	0 . 4
B 1- 38 4- 05	Ga bbr ono rite	Mesa ba	1 . 7 7	1 2. 0	< 5 0.	< 0. 2	< 0. 1	< 0. 0 6	6 5 0	7 6 0	1 2 9	-	0. 4 2	< 5 2	< 2 7	2 . 1	4 . 1	1 . 9	1 . 3	1 0 . 4	4 2 . 9	1 9 . 1
B 1-	Ga bbr	Mesa	1 . 1	6. 1	1. 8	0. 0	0. 7	1. 3	1 8	7 7	1 1	3. 8	1 0.	<5	-	2 . 0	1 0	2 2	3 9	4 2	1 3	1 6

38 4- 08	ono rite	ba	8 0		6	9	7	6	0 4	1 8	6	6	9 1			0	.	.	.	2	9	9
B 1- 38 4- 10	Ga bbr ono rite	M esa ba	4 . 0 4	7. 7	0. 7	0. 9	-	-	2 3 3 7	2 2 4 2	9 6	1 4. 9 8	-	-	-	1 . 8	2 . 4	1 . 8	1 . 7	5 8 . 6	2 9 6 1	1 1 9 . 2
B 1- 38 4- 11	Ga bbr ono rite	M esa ba	1 . 9 1	-	5. 8 3	0. 8 7	-	-	4 5 2 4	5 7 6 7	2 3 0 2	4. 4 2	-	-	-	1 . 8	3 . 1	1 3 . 5	1 4 . 2	5 9 . 7	5 1 0. 6	2 6 . 9
B 1- 38 4- 21	Ga bbr ono rite	M esa ba	0 . 2 9	1 2. 4	< 5	< 0. 2	< 0. 1	< 0. 6	3 9 0 0	2 0 7 0	8 7	-	0. 7 2	7. 00	< 2	< 0 . 5	0 . 2	< 2 . 1	0 . 3	5 . 1	1 6. 5	7 . 5
B 1- 38 4- 23	Ga bbr ono rite	M esa ba	2 . 8 1	-	4. 3 9	1. 2 0	-	-	1 5 5 6	1 0 1 5 3	1 3 1	4. 7 0	-	-	-	0 . 6	0 . 8	5 . 1	1 . 7	1 6 . 8 3	3 9 . 1	1 2 . 1
26 08 6 A- 01	Ga bbr ono rite	No rth M et	0 . 0 9	-	1. 0 3	1. 9 1	-	-	3 1 2	5 1 2	5 3	-	0. 5 0	-	-	0 . 3	0 . 6	2 . 1	1 . 9	2 4 . 8	8 8. 5	7 . 2
26 08 6 A- 03	Ga bbr ono rite	No rth M et	1 . 2 5	-	1. 4 2	1. 0 6	-	-	3 0 7 5	5 8 3 4	9 7	3. 8 9	-	-	-	2 . 0	6 . 4	8 . 6	2 . 8	4 4 . 0	1 0 8 5. 3	9 3 9
26 08 6 A- 04	Ga bbr ono rite	No rth M et	1 . 2 5	-	1. 0 9	0. 7 2	-	-	7 3 7 9	3 0 6 5	1 8 8	1. 4 3	3. 3 3	-	-	1 . 8	3 . 3	1 5 . 8	2 9 . 5	7 1 . 6	2 0 8 3. 1	5 7 . 9
26 08 6 A- 08	Ga bbr ono rite	No rth M et	0 . 8 2	-	1. 2 7	0. 8 9	-	-	1 2 3 3	4 3 8 6	8 0	3. 2 3	2. 8 0	-	-	0 . 9	3 . 3	9 . 2	1 . 2	1 7 . 1	9 6 1. 4	2 9 . 9
26 08 6	Ga bbr ono	No rth M	0 . 4	1 0. 7	3. 4 8	1. 9 7	-	-	5 2 0	7 7 2	5 8	0. 7 2	0. 7 1	-	-	1 . 1	0 . 4	1 . 2	2 . 2	9 . 9	4 2. 8	5 . 4

A-18	rite	et	5														8					
26086A-21	Gabronrite	NoMeth	501	128	2168	0.78	-	-	3382	5231	272	2.88	-	-	-	0.23	3<147	14.3	54.3	315.6	129.6	
D C-21	Gabronrite	NoMeth	0.25	-	<0.33	0.34	-	-	228	112	76	-	-	-	<0.9	0.2	<0.2	0.25	4.5	4.5	-	
D C-22	Gabronrite	NoMeth	0.01	-	<0.33	0.47	-	-	295	169	84	-	-	-	<0.5	0.0	<0.4	<0.2	<0.4	<2.8	2	
D C-23	Gabronrite	NoMeth	0.01	-	<0.5	<0.2	<0.1	<0.06	3100	2700	66	-	0.16	<0.5	<0.2	<0.3	0.39	4.8	0.8	6.7	10.6	32
D C-24	Gabronrite	NoMeth	0.06	-	0.86	1.62	-	-	189	164	74	-	-	-	1.1	0.1	<0.7	0.2	3.1	3.7	48	
D C-25	Gabronrite	NoMeth	<0.1	-	0.34	0.31	-	-	198	129	62	-	-	-	<0.4	0.3	<0.0	0.5	<0.3	6.1	31	
D C-28	Gabronrite	NoMeth	0.01	-	<0.39	0.45	-	-	266	62	74	-	-	-	<0.3	0.4	<0.1	0.5	<0.1	<2.8	19	
D C-29	Gabronrite	NoMeth	0.03	-	0.82	4.35	-	-	123	155	44	-	-	-	<0.6	0.4	<0.5	0.4	8.5	6.9	59	
D C-30	Gabronrite	NoMeth	<0.1	-	<0.5	<0.2	<0.1	<0.06	150	30	36	-	0.02	<0.5	-	<0.7	0.3	<0.3	1.5	<0.0	42	
D C-31	Gabronrite	NoMeth	<0.1	-	<0.44	0.64	-	-	97	87	62	-	-	-	0.5	0.1	<0.0	0.9	2.6	2.0	18	
D C-32	Gabron	NoM	0.0	-	<0.5	1.07	-	-	244	132	69	-	-	-	<0.1	0.1	<0.1	0.4	5.3	3.3	89	

	rite	et	3		0										9		7					
D C- 33	Ga bbr ono rite	No rth M et	0 . 0 6	- . 4 3	< 0. 4 3	0. 5 0	- . - .	- . - .	1 9 0	1 7 3	9 7 7	- . - .	- . - .	- . - .	< 0 . 1 8	0 . 4 1	< 0 . 3 7	< 0 . 4 7	3. 1 .	2 6 5		
D C- 50	Ga bbr ono rite	No rth M et	0 . 0 9	- . - .	< 5 .	< 0. 2	< 0. 1	< 0. 0	1 8 0	4 5 0	6 5 0	- . - .	0. 1 8	< 5 .	< 2 .	< 1 . 3 2	< 0 . 2 2	0 . 5 5	7 . 4 2	1 4 2	< 2 .	
D C- 53	Ga bbr ono rite	No rth M et	0 . 1 5	6. 7	2. 2 3	1. 0 3	- . - .	- . - .	2 3 0	5 1 4	5 5	- . - .	- . - .	- . - .	< 0 . 6 8	0 . 5 0	< 1 . 5 2	1 . 0 5	1 . 0 2	3 9. 0	- .	
D C- 55	Ga bbr ono rite	No rth M et	1 . 1 1	9. 0	2. 3 4	0. 1 4	0. 2 3	0. 4 5	1 0 7	3 6 5	8 9 2	0. - .	- . - .	- . - .	0 . 9	1 . 0	6 . 9	5 . 1	4 . 1	1 4. 3	3 2 2	
D C- 56	Ga bbr ono rite	No rth M et	1 . 4 4	9. 2	3. 5 2	2. 8 6	- . - .	- . - .	1 9 0	5 5 0	1 7 0	1. - .	- . - .	- . - .	1 . 6	2 . 8	1 . 9	1 . 4	1 . 7	4 0 2. 1	4 7 6	
D C- 58	Ga bbr ono rite	No rth M et	2 . 0 0	6. 6	0. 7 1	2. 6 5	- . - .	- . - .	2 4 3	1 0 2	1 1 3	0. - .	- . - .	- . - .	1 . 9	3 . 0	1 . 3	1 . 3	1 . 7	7 1 2. 0	1 4 0	
D C- 59	Ga bbr ono rite	No rth M et	0 . 1 2	- .	0. 3 4	1. 1 0	- . - .	- . - .	5 6 7	1 0 1	6 8	- . - .	- . - .	- . - .	< 1 . 4	0 . 1	< 3 . 1	0 . 3	< 5 . 4	5. 1	- .	
D C- 61	Ga bbr ono rite	No rth M et	1 . 5 0	7. 5	< 0. 3 8	0. 5 6	- . - .	- . - .	2 0 5	7 1 9	1 1 3	1. - .	- . - .	- . - .	2 . 5	3 . 8	1 . 2	1 . 2	1 . 3	5 0 1. 3	5 8 9	
D C- 62	Ga bbr ono rite	No rth M et	1 . 6 9	1. 6	2. 8	0. 4	0. 2	0. 0	3 7 4	1 2 2	1 5 0	1. - .	- . - .	- . - .	3 . 2	1 . 5	3 . 2	7 . 9	5 . 7	2 2 5 0. 6	4 2 1 4	
D C- 63	Ga bbr ono rite	No rth M et	2 . 4 5	7. 1	7. 3	0. 4	0. 6	0. 8	8 2 4	2 2 1	1 8 0	0. - .	- . - .	- . - .	1 . 2	0 . 9	< 5 . 0	3 . 2	1 . 6	6 2. 9	1 8 4	
D	Ga	No	0	2.	<	<	<	<	6	6	7	-	3.	<	<	4	8	3	5	2	1	8

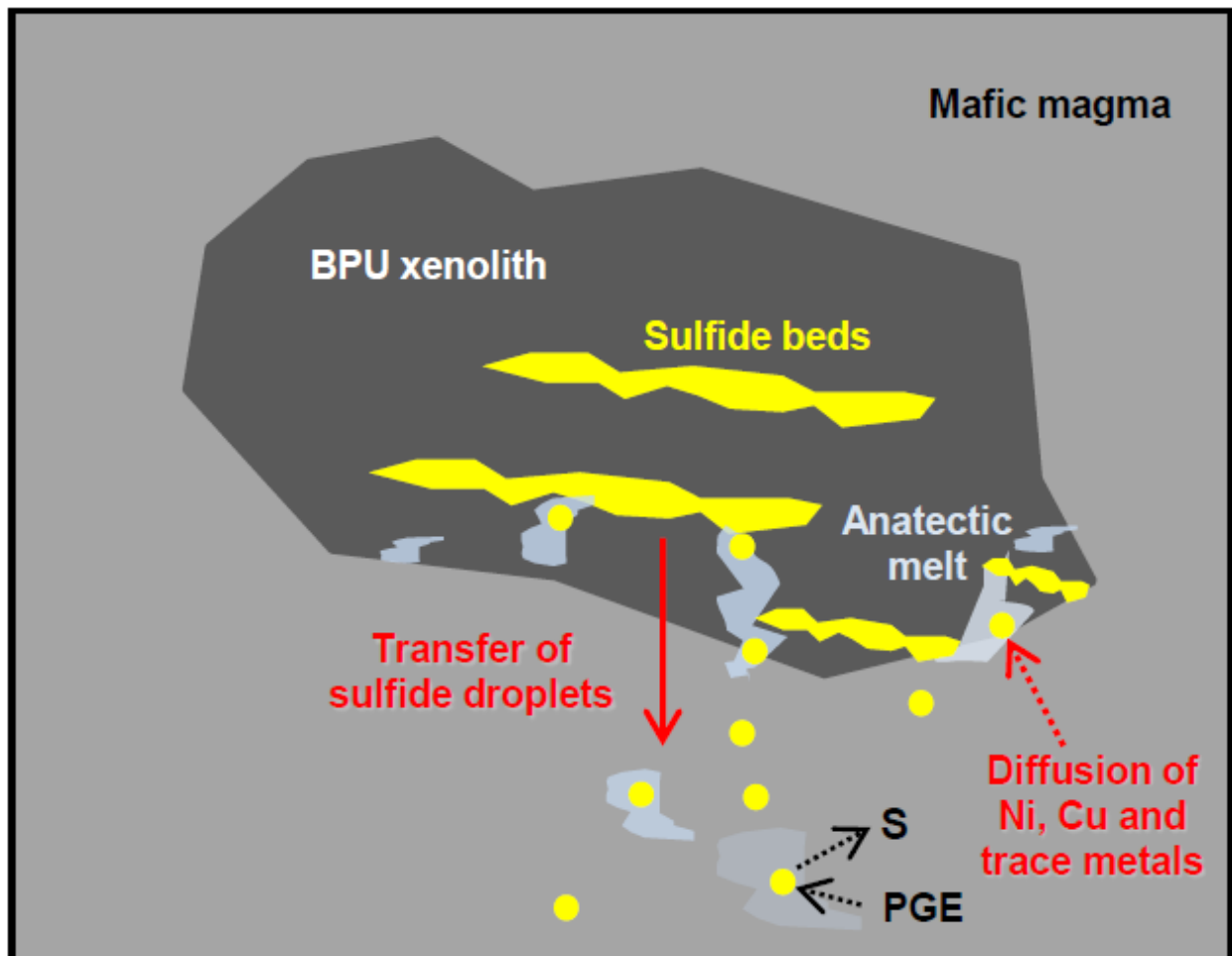
C-64	bbr ono rite	rth M et	. 5 9	8	5	0. 2	0. 1	0. 0 6	4 0	9 0	5		3 0	5	2	. 1	. 6	1 .5	8 .0	9 .1	6 3 0. 4	8 7
D C- 65	Ga bbr ono rite	No rth M et	2 .1 7	1 1. 9	1 4. 8 3	1. 3 8	-	-	1 4 6 5	1 5 6 2	1 0 9	3. 9 1	-	-	-	2 .3	0 .4	< 5 0	2 .9	3 7 1 .1	6 0 0. 7 0	2 6 5 0
D C- 66	Ga bbr ono rite	No rth M et	4 .3 2	9. 8	2. 2. 5 8	0. 3 5	0. 6 9	0. 7 7	2 2 8 3	6 3 1 5	3 1 0	1. 0 2	-	-	-	2 .2	4 .3	2 0 .3	2 3 .0	3 5 .1	4 6 3. 7	4 5 1
D C- 79	Ga bbr ono rite	No rth M et	1 .1 2	7. 6	4. 6 8	0. 6 0	-	-	2 2 4 6	7 1 2 0	1 2 0	-	-	-	-	1 .3	3 .0	1 0 .1	1 5 .0	1 6 3 .9	5 8 3	6 6 7
B 1- 17 5- 01 .	Ga bbr ono rite	Se rpe nti ne	3 .2 1	-	1 2. 3 6	1. 5 8	-	-	2 5 0 6	8 2 4 0	1 9 8	6. 3 3	5. 4 8	-	-	< 0 .9	1 .5	4 .6	3 .4	1 8 .4	1 3 3	4 5 .7
B 1- 17 5- 06	Ga bbr ono rite	Se rpe nti ne	0 .9 2	1 3. 7	2 0. 8 4	2. 4 4	-	-	7 8 3	2 7 1 9	8 9	3. 1 8	1. 2 5	-	-	< 0 .7	0 .4	3 .1	1 .7	1 7 4	6 0. 0 4	2 0 4
B 1- 17 5- 11	Ga bbr ono rite	Se rpe nti ne	2 .0 2	-	1 7. 8 6	2. 0 4	-	-	1 0 1 7	5 3 6 2	1 2 7	4. 3 3	2. 6 3	-	-	0 .6	0 .8	1 .9	1 .7	8 5 .6	6 5. 8	2 0 9
B 1- 17 5- 16	Ga bbr ono rite	Se rpe nti ne	6 .5 2	-	3 5. 1 3	0. 5 8	-	-	1 1 9 4	4 1 9 0	1 7 3	2. 0 1	2. 1 4	-	-	2 .3	1 .0	1 .0	2 .5	< 1 .4	3 9. 2 9	3 5 .9
B 1- 17 5- 17	Ga bbr ono rite	Se rpe nti ne	4 .5 8	-	3 7 3	2. 2 1	-	-	2 5 4 1	9 7 9 7	2 2 5	3. 3 1	5. 2 2	-	-	0 .9	1 .4	2 .8	2 .4	2 1 .8	5 5. 8	2 0
B 1-	Ga bbr	Se rpe	2 .1	-	1 1.	2. 0	-	-	1 4	1 4	1 1	9. 2	3. 2	-	-	0 .1	1 .1	4 .1	2 .1	7 0	9 1.	3 6

17 5- 29	ono rite	nti ne	3 3		7 0	2			2 7	2 8 3	5	2	9			8	0	0	5	.	8	.
B 1- 17 5- 30	Ga bbr ono rite	Se rpe nti ne	3 . 0 0	1 5. 2	3. 3 1	0. 3 6	0. 3 4	0. 8 6	2 2 5 7	5 3 1 8	1 9 3	3. 2 1	3. 9 2	-	-	0 . 5	1 . 0	2 . 5	2 . 4	2 . 5 5	6 8. 7	2 5 . 2
26 13 3- 01	Ga bbr ono rite	W ym an cre ek	0 . 0 4	-	0. 3 0	0. 0 5	-	-	2 0 5	5 1	4 3	-	-	-	-	0 . 8	0 . 1	7 . 1	1 . 4	7 . 0	0. 5	0 . 5
26 13 3- 02	Ga bbr ono rite	W ym an cre ek	0 . 0 3	-	0. 3 9	0. 0 7	-	-	2 3 6	4 3	4 8	-	-	-	-	0 . 1	0 . 5	0 . 8	4 . 3	0. 4	0 . 3	

Appendix B2. Values obtained for reference materials.

LabMaTer S determinati on by infrared spectrometr y S-C analyser	S (%)								
KPT-1 this work	1.08								
<i>KPT-1</i> (<i>Webb et</i> <i>al. 2006</i>)	<i>1.029+</i> <i>/.034</i>								
University of Waterloo Whole-rock $\delta^{34}\text{S}$	$\delta^{34}\text{S}$ (‰ VCDT)								
DC-64 (This study)	3.0								
DC-64 (Duplicate this study)	2.8								
DC-64 (Thériault and Barnes, 1998)	2.5								
B1-384-26 (This study)	18.4								
B1-384-26 (Duplicate this study)	18.6								
B1-384-26 (Queffurus and Barnes, 2014)	17.3								
LabMaTer Semimetals determined by black shale method	As (ppm)	Sb (ppm)	Bi (ppm)	Te (ppm)					
SDO-1 this	68.41	4.69	0.37	0.29					

<i>(Webb et al., 2006)</i>	-71	-102	/-5	0.15	0.53	1.13	-0.15	/-0.994	/-0.118
LabMaTer Ni-FA-ICP-MS	Os (ppb)	Ir (ppb)	Ru (ppb)	Rh (ppb)	Pt(ppb)	Pd (ppb)	Au (ppb)		
OKUM - this study	0.7	0.9	4.5	1.4	10.4	12.0	1.0		
<i>OKUM (Savard et al., 2010)</i>	0.98+/- 0.34	0.99 +/- 0.07	4.25+/- 0.3	1.40 +/- 0.13	11+/- 0.6	11.7+ /-0.5	1.4		
SLg-1 black shale - this study FA	<0.16	0.05	0.38	0.52	3.85	1.37	1760		
SLg-1 black shale - this study ID	<0.03	0.06	0.66	-	3.06	0.98	-		
<i>SLg-1 (Li et al., 1998)</i>	<i>n.d.</i>	0.02+/- 0.013	0.27+/- 0.03	0.32+/- 0.05	1.39+ /-0.12	1.49+ /-0.13	1690+ /-900		
<i>SLg-1 (Petrov et al., 2004)</i>	<i>n.d.</i>	<i>n.d.</i>	1+/-0.4	-	2.2+/- 0.5	2.3+/- 0.6	2500+ /-300		



Graphical abstract

Highlights

- We examine interactions between black shale xenoliths and mafic magma.
- Sulfide droplets are entrained in the xenolith melt and transferred to the magma.
- Sulfur and semimetals contamination of the mafic magma is shown and modeled.

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